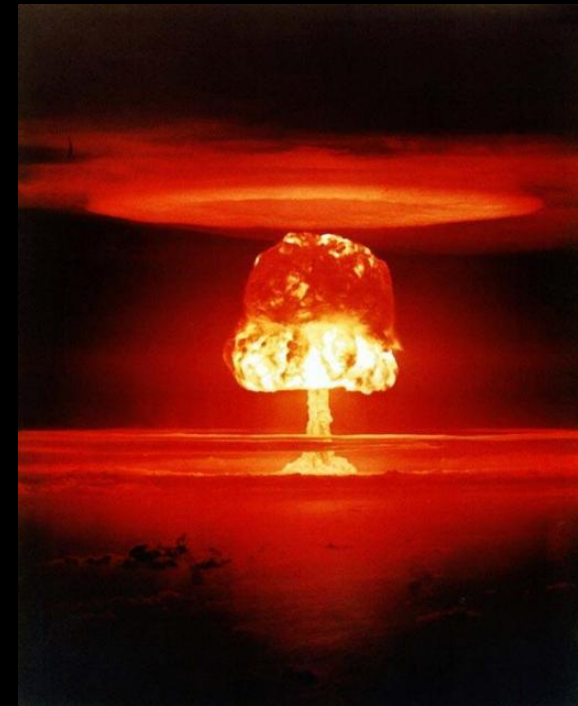
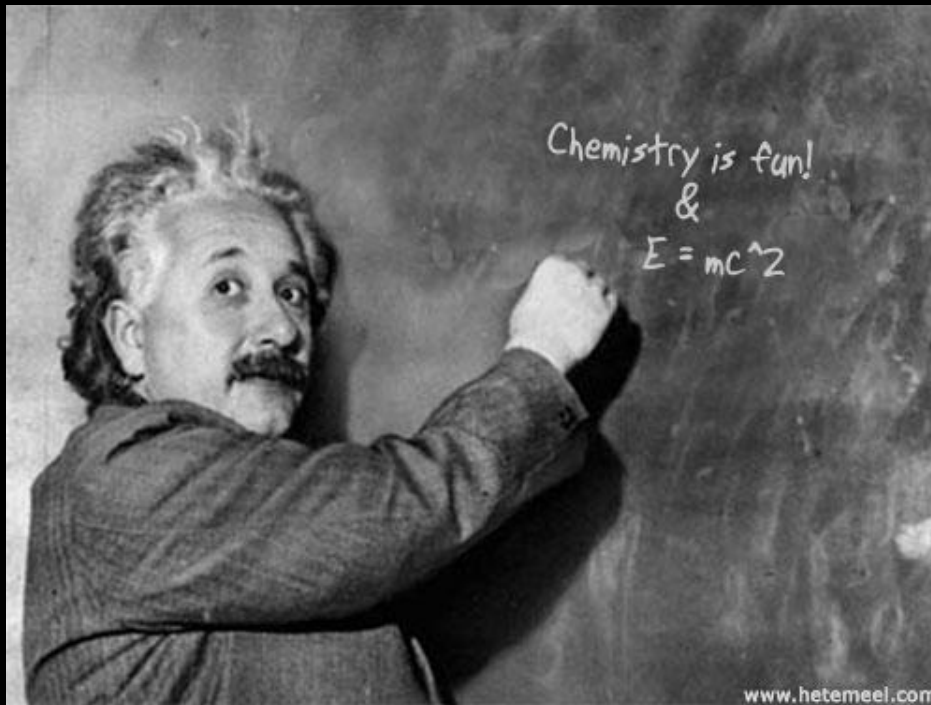
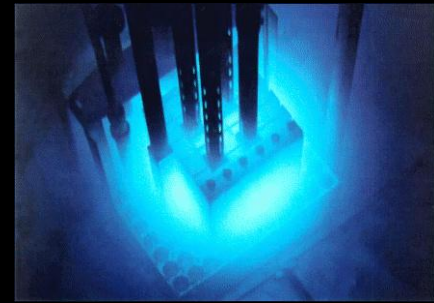
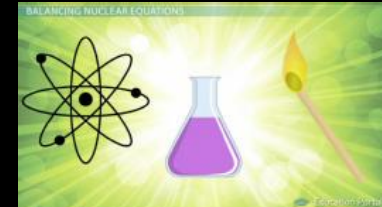


Nuclear Chemistry



Nuclear Chemistry



Nuclear chemistry: study of changes in the atomic nucleus

Recall:

Z: Atomic Number; number of protons in nucleus

A: Atomic Mass number; sum of protons and neutrons in nucleus

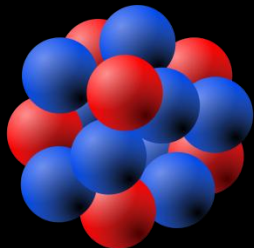
Isotopes: atoms with same Z, but different A

Nuclide: a single atom type

Nuclide Symbol: A_ZX where X = symbol of the element

Nuclide Name: Element - A (Example: ${}^{14}_6\text{C}$ = Carbon-14)

Nucleon: collective name for protons & neutrons

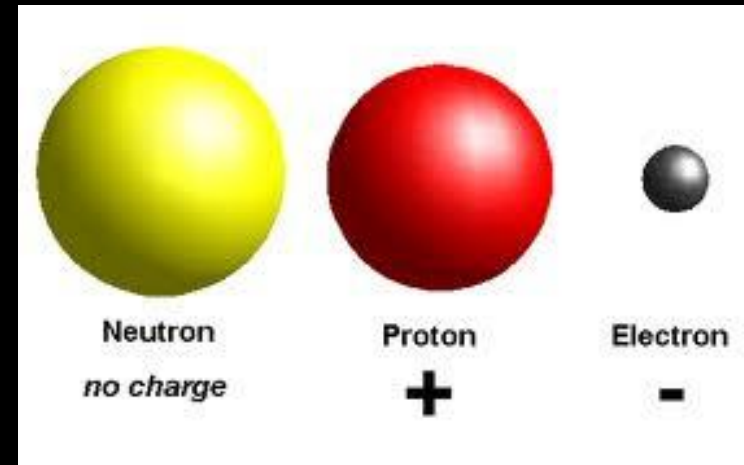
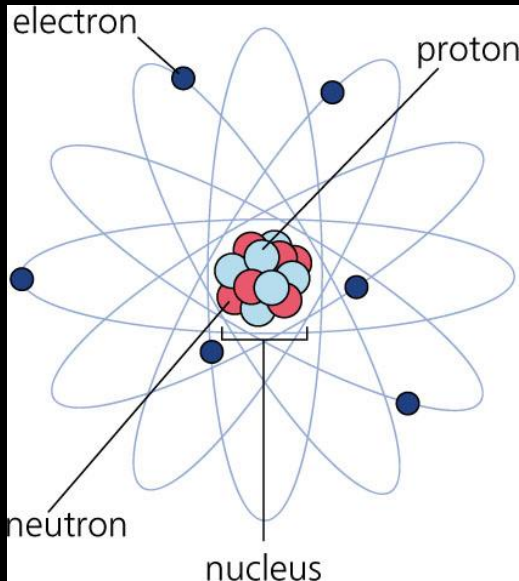


Proton: positive charge

Neutron: no charge

Primary Subatomic Particles

Particle	Location (Nucleus)	Charge	Mass (g)	Mass (amu)
Neutron (n^0)	Inside	0	1.675×10^{-24}	1.00867 (~1)
Proton (p^+)	Inside	+1	1.673×10^{-24}	1.00728 (~ 1)
Electron (e^-)	Outside	-1	9.11×10^{-28}	0.000549 (~0)



Mass Order: $n > p \gg e$

The Nucleus

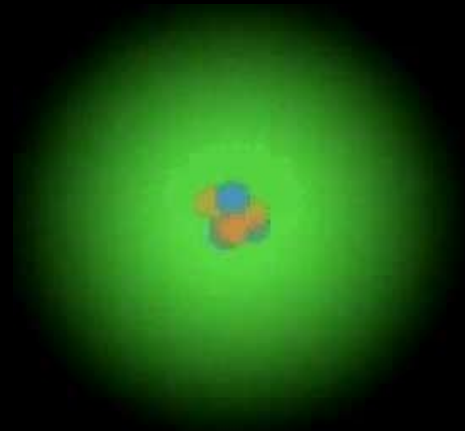
Most mass concentrated in a small, dense core (the nucleus)

Positively charged

Nucleus Radius: $\sim 10^{-15}$ m

Atomic Radius: $\sim 10^{-10}$ m

Atomic Density: $\sim 1.8 \times 10^{14}$ g/cm³



For He atom: If nucleus equals the size of a penny
Atomic sphere diameter equals the size of the pentagon



Density of Neutron Star

Supernova collapse creates neutron star (most dense stars known)

What is the density of the neutron star:

Mass: 1.99×10^{30} kg

Diameter: 26 km

$$D = \frac{1.99 \times 10^{30} \text{ kg}}{\frac{4}{3} \pi (26 / 2 \text{ km})^3}$$

$$D = 5.2 \times 10^{17} \text{ kg/m}^3$$



What is the density of a uranium - 235 nucleus:

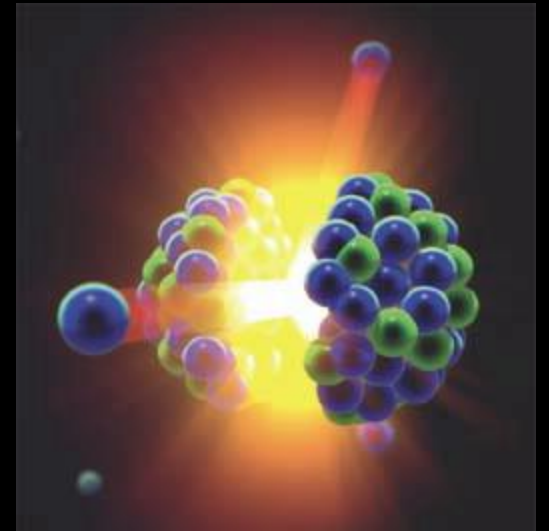
Mass: 235 amu (1.66×10^{-27} kg / amu)

Radius: 7.5×10^{-15} m

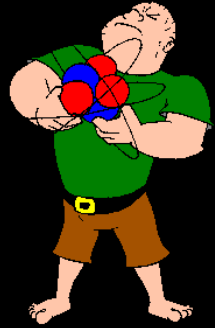
$$D = \frac{235 \text{ amu} (1.66 \times 10^{-27} \text{ kg / amu})}{\frac{4}{3} \pi (7.5 \times 10^{-15} \text{ m})^3}$$

$$D = 2.2 \times 10^{17} \text{ kg/m}^3$$

Nucleus is about 2 x density of neutron star



Strong Nuclear Force



Holds strongly repulsive positive charges (protons) together
Only relevant at extremely short (nuclear) distances

Theoretical mass of the Helium - 4 Atom:

Total mass: sum of mass of 2 protons, 2 neutrons and 2 electrons

Total mass: $(2 \times 1.0072 \text{ amu}) + (2 \times 1.0087 \text{ amu}) + (2 \times 0.00055 \text{ amu})$

Total mass: 4.0331 amu

Value of Helium – 4 atom by mass spectroscopy

4.0026 amu

Difference: $4.0331 \text{ amu} - 4.0026 \text{ amu} = 0.0305 \text{ amu}$

Mass Defect:

Mass loss represents conversion to energy

This energy holds nucleus together

Nuclear Binding Energy (of He - 4)

Mass to Energy conversion defined by Einstein's $E = mc^2$

To use energy term of Joules ($1 \text{ J} = 1 \text{ kg m}^2 / \text{s}^2$), need to express mass in kg:

$$0.0305 \text{ g/mol} \times 1 \text{ kg} / 1000 \text{ g} = 3.05 \times 10^{-5} \text{ kg} / \text{mol}$$

$$E = mc^2 \text{ (where } c = 2.998 \times 10^8 \text{ m/s)}$$

$$E = \frac{3.05 \times 10^{-5} \text{ kg}}{\text{mol}} \times \frac{(2.998 \times 10^8 \text{ m})^2}{\text{s}^2} = 2.74 \times 10^{12} \text{ kg m}^2 \text{ s}^{-2} \text{ mol}^{-1}$$

$$E = 2.74 \times 10^{12} \text{ J mol}^{-1}$$

Molar Binding Energy:

$$E = 2.74 \times 10^{12} \text{ J mol}^{-1} \times \frac{1 \text{ mol}}{6.02 \times 10^{23} \text{ nuclei}} = 4.55 \times 10^{-12} \text{ J}$$

Energy in terms of electron volts:

$$E = 4.55 \times 10^{-12} \text{ J} \times \frac{1 \text{ eV}}{1.602 \times 10^{-19} \text{ J}} = 2.84 \times 10^7 \text{ eV} \rightarrow 28.4 \text{ MeV}$$

Problem: Binding Energy (MeV) of F - 19

Theoretical mass of the Fluorine-19 Atom:

Total mass: sum of mass of 9 protons, 10 neutrons and 9 electrons

Total mass: (9 x 1.0072 amu) + (10 x 1.0087 amu) + (9 x 0.00055 amu)

Total mass: 9.0648 amu + 10.087 amu + 0.0050 amu (only allowed 2 sig figs here)

Total mass: 19.157 amu

Mass spectroscopy value of fluorine – 19 atom:

18.9984 amu

Difference: 19.157 amu – 18.9984 amu = 0.163 amu

Energy Equivalent in Joules:

$$E = \frac{1.63 \times 10^{-4} \text{ kg}}{\text{mol}} \times \left(\frac{2.998 \times 10^8 \text{ m}}{\text{s}} \right)^2 = 1.47 \times 10^{13} \text{ kg m}^2 \text{ s}^{-2} \text{ mol}^{-1} = 1.47 \times 10^{13} \text{ J mol}^{-1}$$

Molar Energy Equivalent:

$$E = 1.47 \times 10^{13} \text{ J mol}^{-1} \times \frac{1 \text{ mol}}{6.02 \times 10^{23} \text{ nuclei}} = 2.43 \times 10^{-11} \text{ J}$$

Energy in electron volts:

$$E = 2.43 \times 10^{-11} \text{ J} \times \frac{1 \text{ eV}}{1.602 \times 10^{-19} \text{ J}} = 1.51 \times 10^8 \text{ eV} \rightarrow 151 \text{ MeV}$$

Nuclear Stability

Nucleus is stable: does not change (transform) without outside energy

~ 250 nuclei are stable

~ 3100 nuclei known

Nucleus is unstable: changes composition by emitting nuclear particles

radioactivity: emission of particles to move towards a more stable nucleus

radioactive: an atom that will spontaneously emit nuclear particles

radioactive decay: emission of nuclear particle

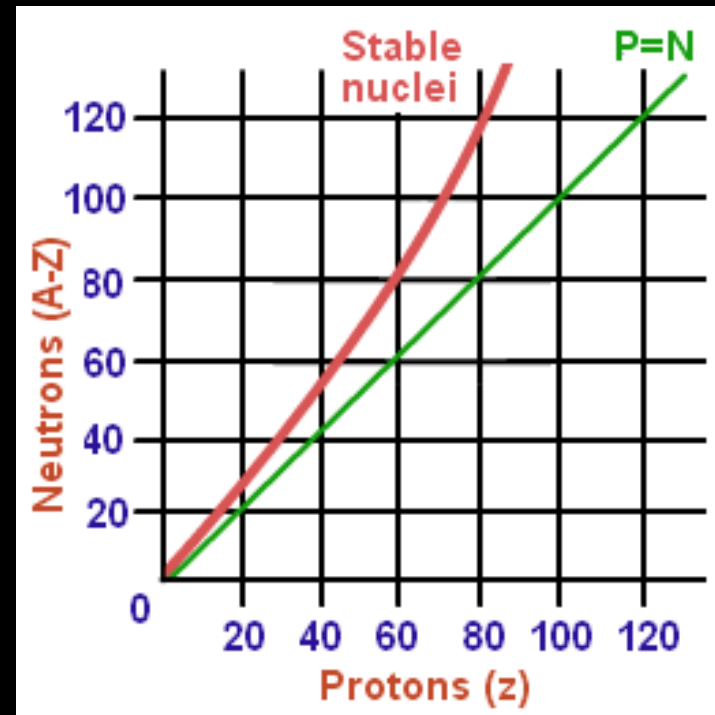
radioisotope: nucleus that spontaneous decays

Larger Z nuclei have more:

protons

proton repulsions

neutrons to overcome electrostatics



Nuclear Stability

All nuclei with $Z > 83$ are unstable

Nuclei likely to be stable:

Even # n

Even # p

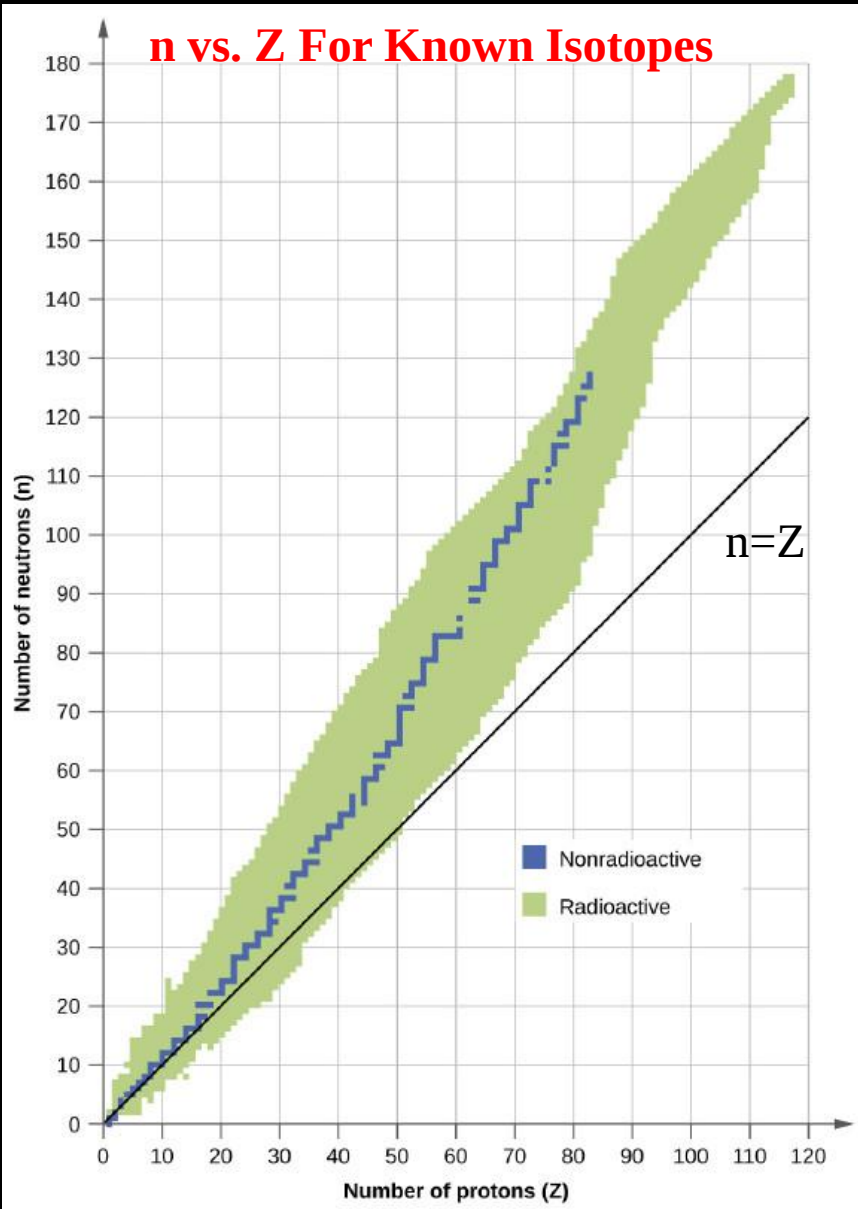
Even # both

“Magic Numbers” (protons or neutrons)
(2, 8, 20, 28, 50, 80, 126)

Nucleons stable against nuclear decay

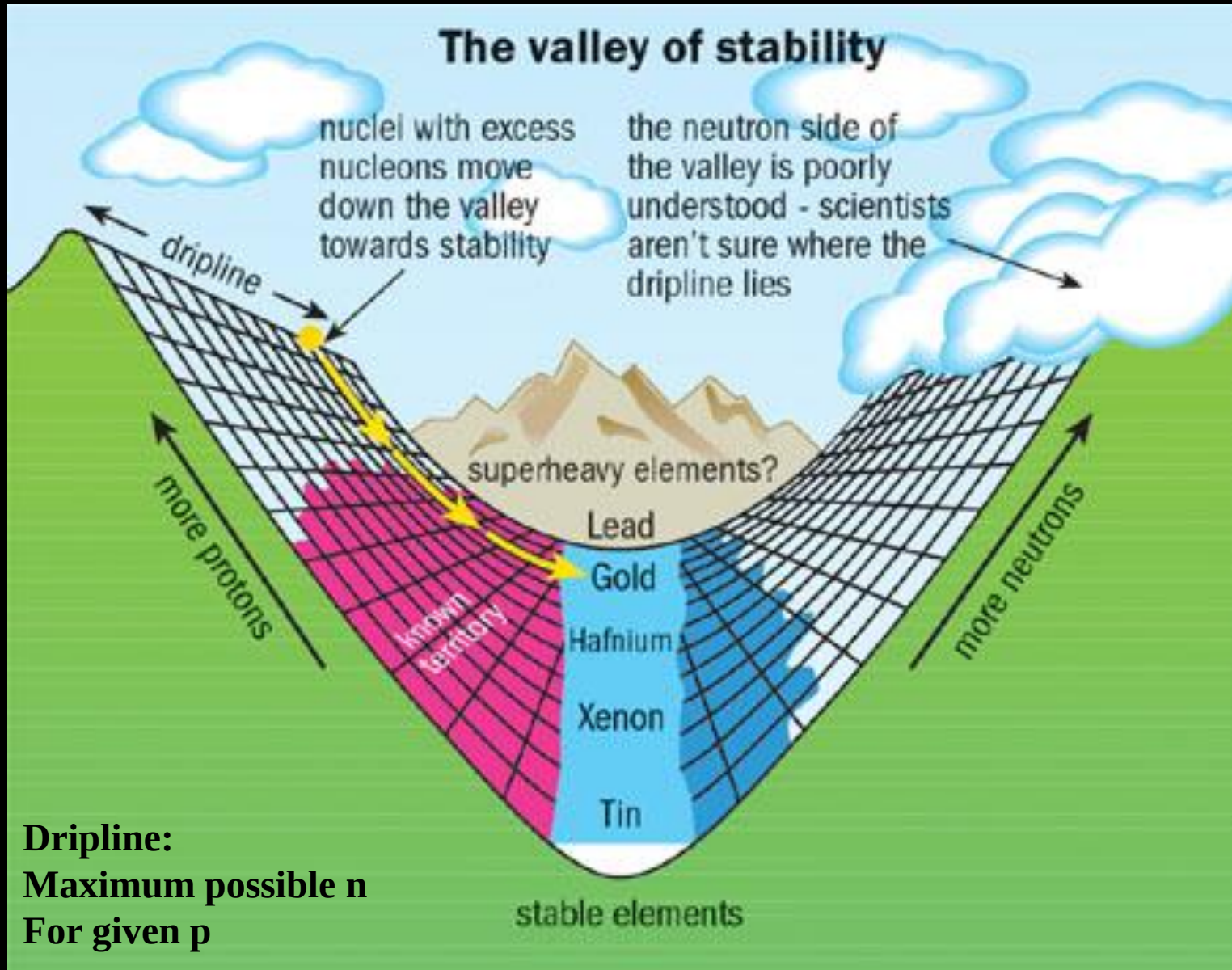
“Double Magic”

Magic number of both n and Z

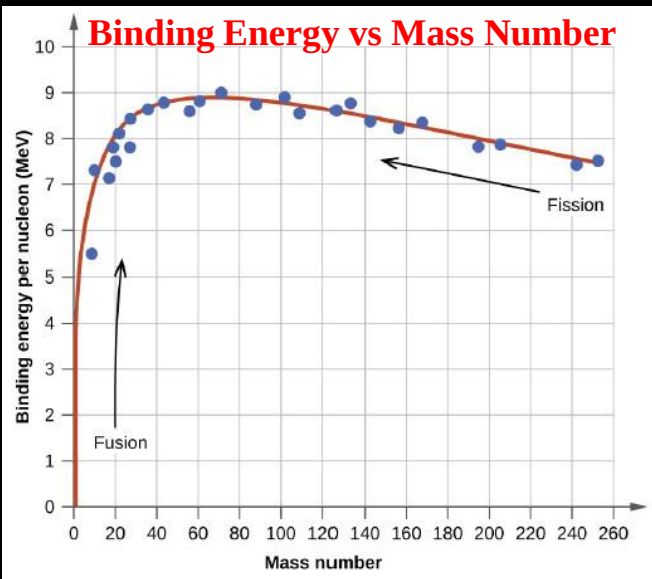


Number of Stable Isotopes	Proton Number	Neutron Number
157	even	even
53	even	odd
50	odd	even
5	odd	odd

Nuclear Stability



Nucleon Binding Energy



$$\text{Nucleon Binding Energy} = \frac{\text{Nuclear Binding Energy}}{\text{\# Nucleons}}$$

For ${}^4_2\text{He}$: Nuclear Binding Energy of 28.4 MeV

$$\frac{28.4 \text{ MeV}}{4 \text{ nucleons}} = 7.10 \text{ MeV}$$

Nucleon Binding Energy greatest about mass number 56
Stars do not create elements (Fusion) beyond Iron

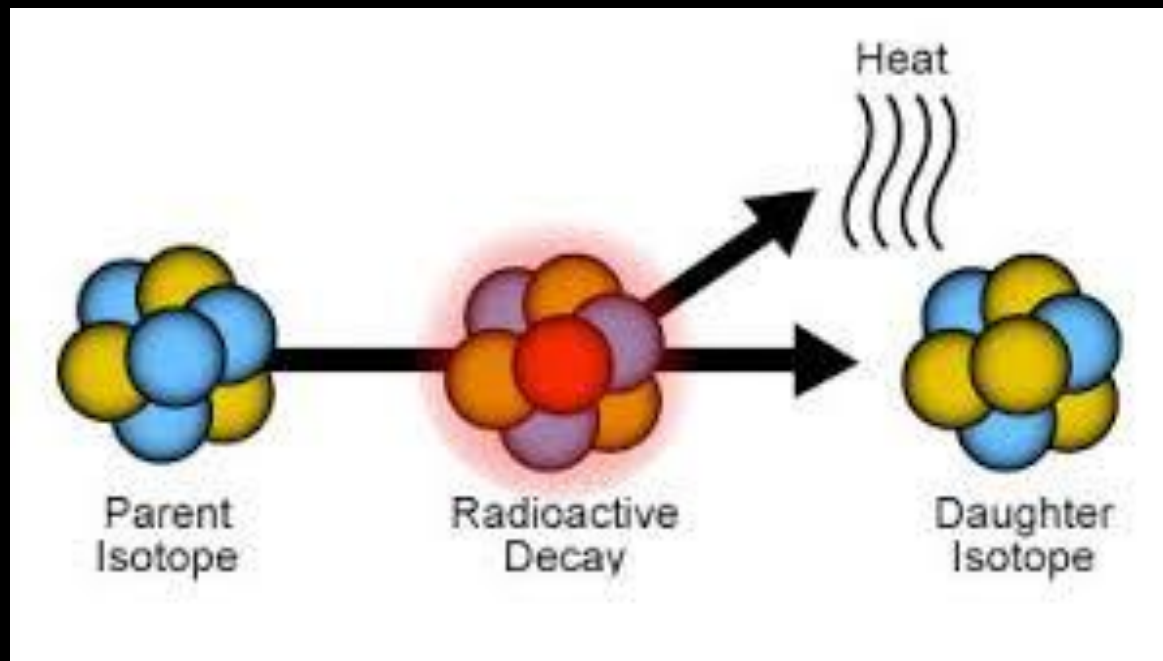
For Fe 56: Nucleon Binding Energy: 8.820 MeV
~ 25% larger than He-4

Nuclear Decay Events

Decay: Radioisotopes move towards a more stable nucleus

Emission of material from nucleus

Changes nucleus



Heat from radioactive decay: adds heat to earth's core

Nuclear Reactions



Nuclear reaction represents the decay process of nuclear rearrangement

Nuclear reaction conservation laws:

sum of mass numbers for reactants and products identical

sum of charges for reactants and products identical

Radioactive decay of polonium-210 can be summarized by the following:



What is element X? **Pb** What is element Y? **He**

amass = 4 **anumber = 2**

The radioactive particle (He nucleus without electrons): α particle

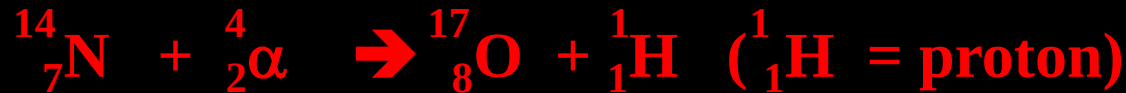
All of the Helium found on earth comes from radioactive decay

Historical Nuclear Reactions

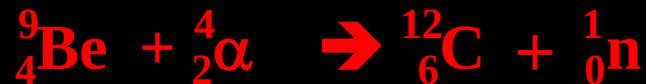
First naturally occurring radioisotope (1898) Marie & Pierre Curie:



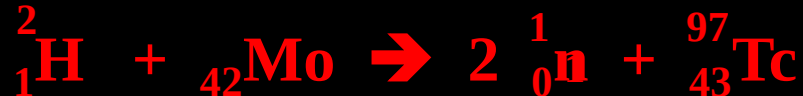
First artificial nuclide prepared (1919) Ernst Rutherford:



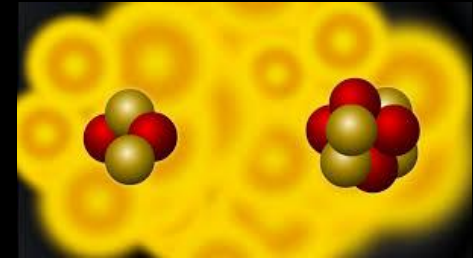
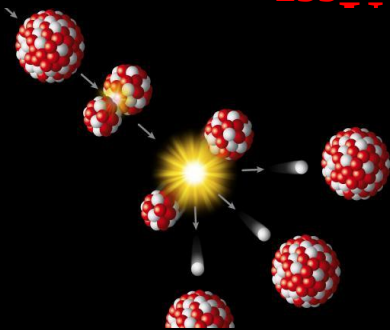
Neutron discovered (1932) by James Chadwick



First artificial element (1937) Technetium:



First controlled sustained chain reaction (1942) at U of Chicago

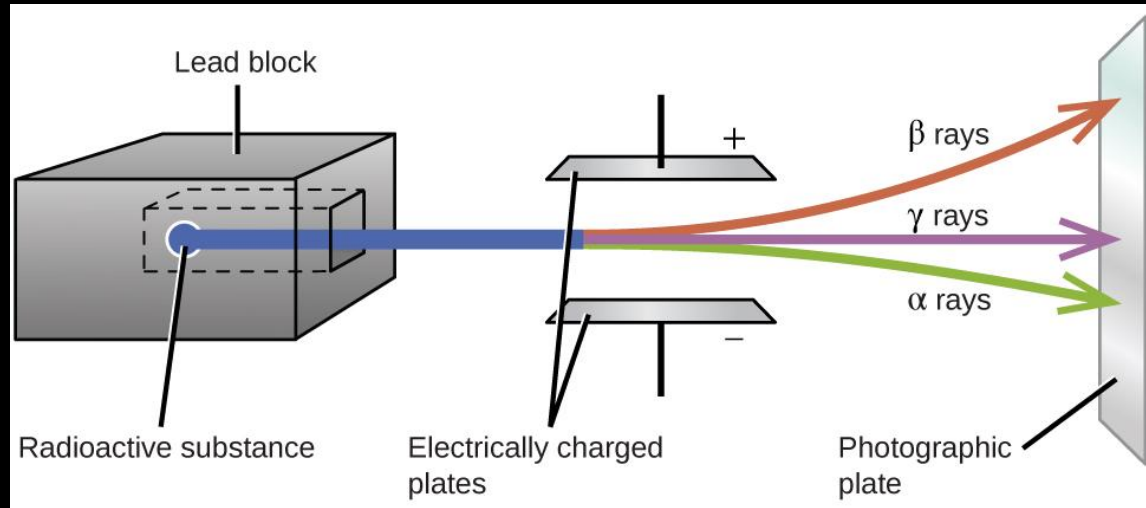


Nuclear Decay Events

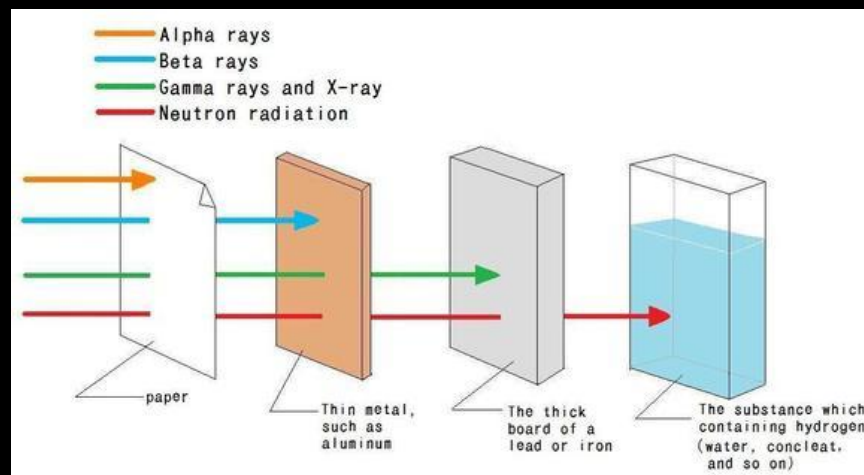
Name	Symbol(s)	Representation	Description
Alpha particle	${}^4_2\text{He}$ or ${}^4_2\alpha$		(High-energy) helium nuclei consisting of two protons and two neutrons
Beta particle	${}^0_{-1}\text{e}$ or ${}^0_{-1}\beta$		(High-energy) electrons
Positron	${}^0_{+1}\text{e}$ or ${}^0_{+1}\beta$		Particles with the same mass as an electron but with 1 unit of positive charge
Proton	${}^1_1\text{H}$ or ${}^1_1\text{p}$		Nuclei of hydrogen atoms
Neutron	${}^1_0\text{n}$		Particles with a mass approximately equal to that of a proton but with no charge
Gamma ray	γ		Very high-energy electromagnetic radiation

Types of Radioactive Decay

Rutherford's Determination of radioactive charge



Relative Penetration Energy



Types of Radioactive Decay

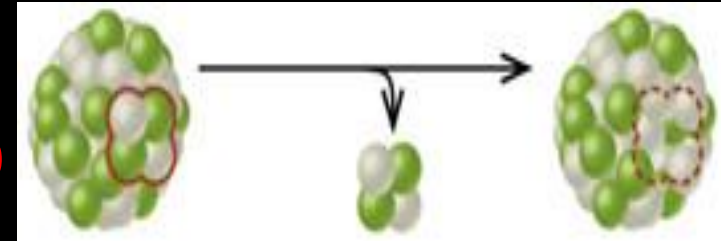
Alpha Particle

Parent nucleus: Emits 2 protons & 2 neutrons (He nucleus)

Daughter nucleus: $A - 4$; $Z - 2$

has higher n:p ratio than parent

Typically from heavier nuclei ($A > 200$; $Z > 83$)



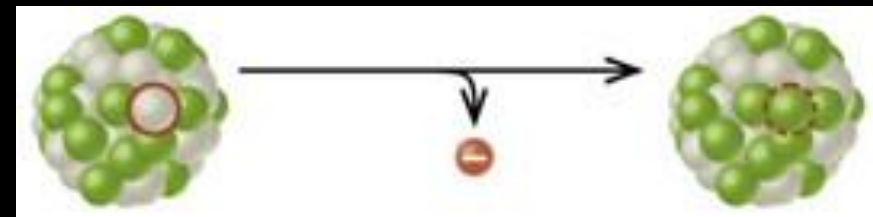
Beta Particle

Parent nucleus: Emits electron

Daughter nucleus: A unchanged; $Z + 1$

has lower n:p ratio than parent

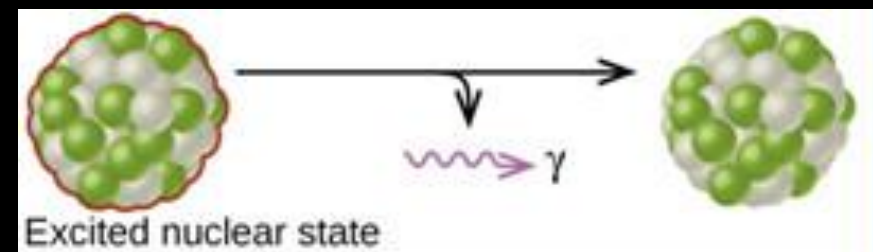
Typically from heavier nuclei ($A > 200$; $Z > 83$)



Gamma Emission

Parent nucleus: emits high energy photon

No effect on either A or Z



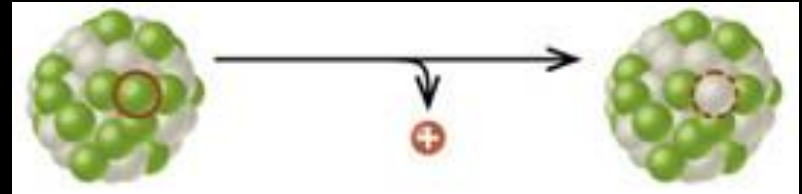
Types of Radioactive Decay

Positron Emission

Parent Nucleus: emits positron

low n:p ratio

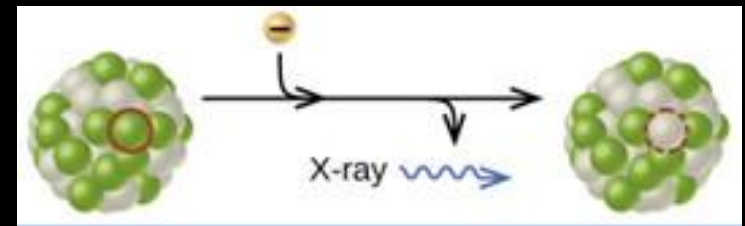
Daughter Nucleus: closer to stability band



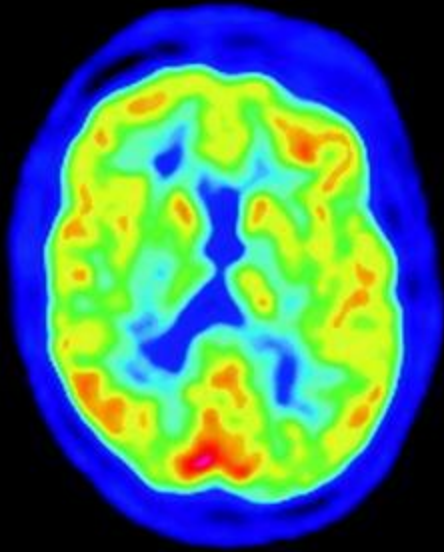
Electron Capture

Parent Nucleus: captures inner shell electron

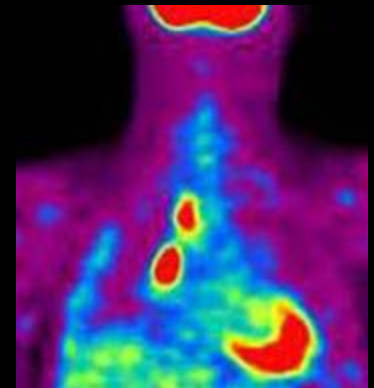
Daughter: A unchanged; Z -1



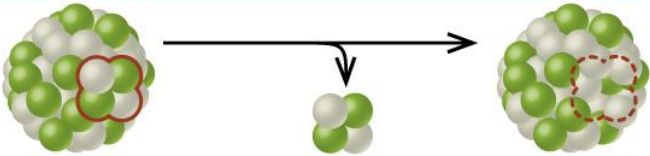
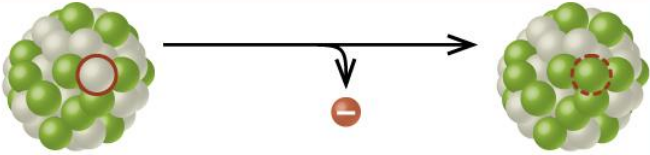
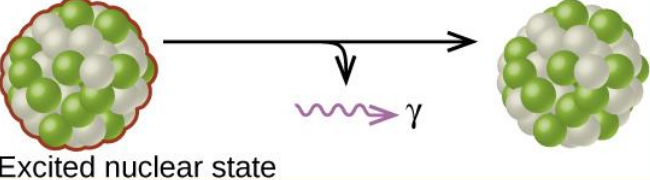
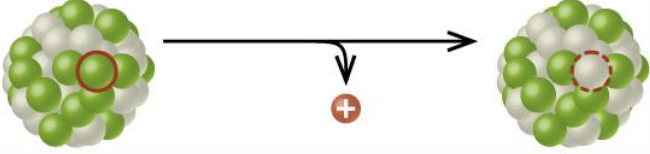
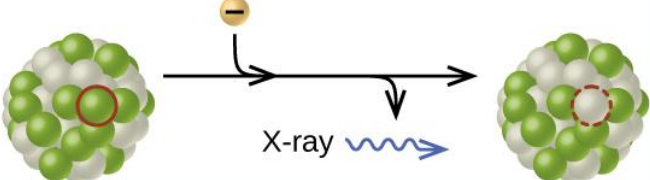
Positron Emission Tomography



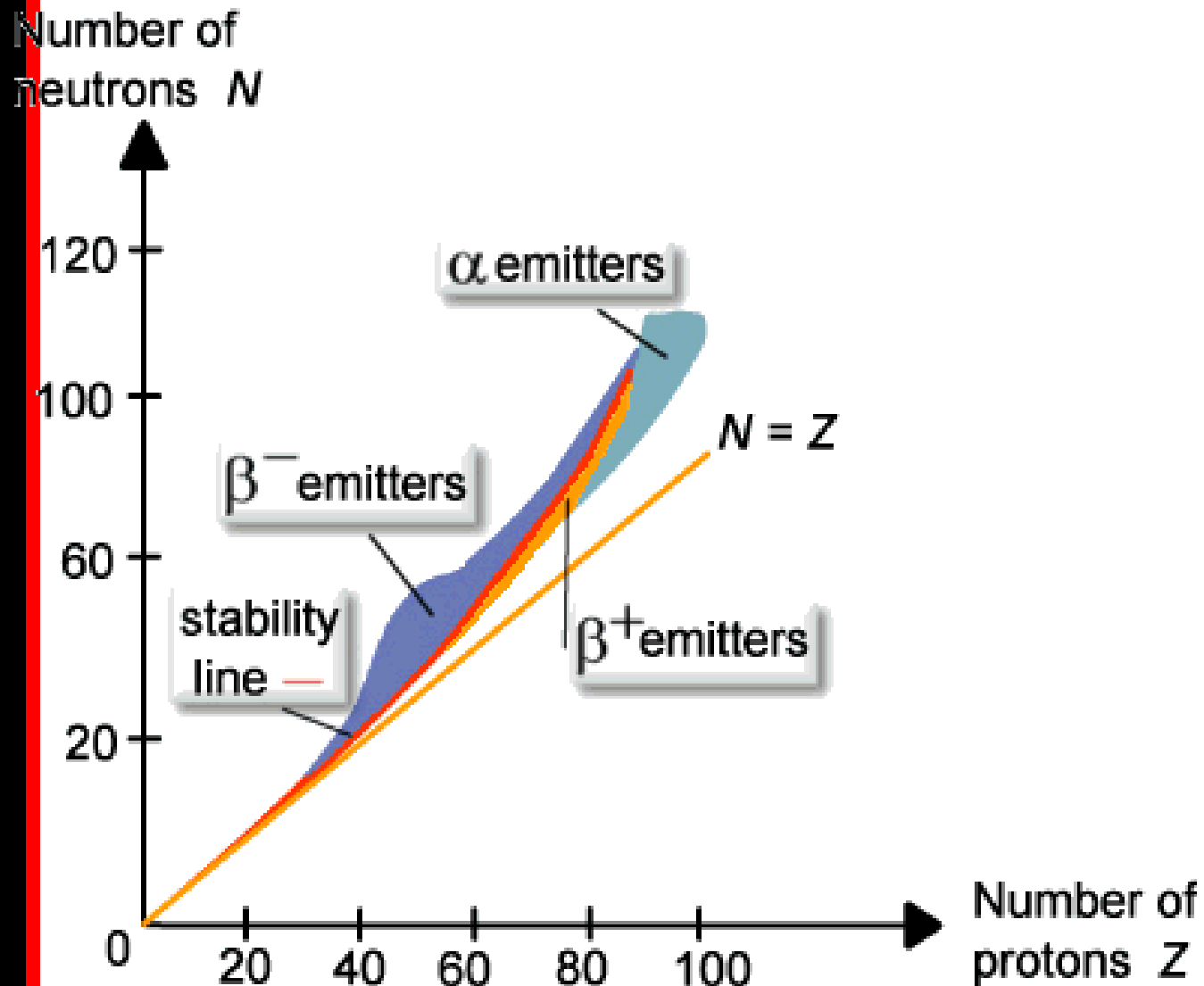
Radioactive F-18 incorporated into glucose
Radioactive decay (gamma radiation) detected
Used to study brain chemistry
highlights increased metabolism
Used to spot cancerous growth



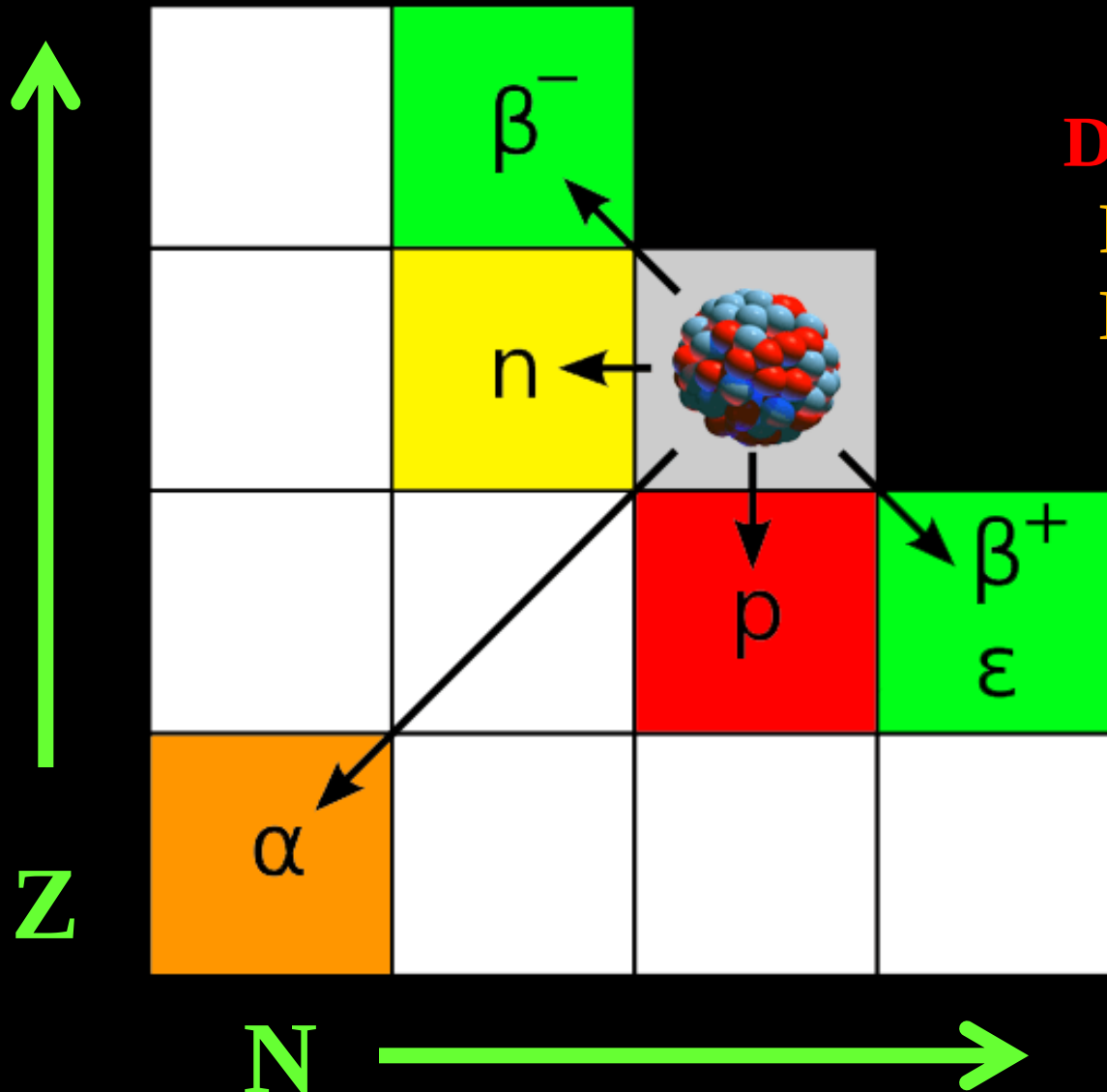
Nuclear Decay Events

Type	Nuclear equation	Representation	Change in mass/atomic numbers
Alpha decay	${}^A_ZX \rightarrow {}^4_2\text{He} + {}^{A-4}_{Z-2}Y$		A: decrease by 4 Z: decrease by 2
Beta decay	${}^A_ZX \rightarrow {}^0_{-1}e + {}^{A}_{Z+1}Y$		A: unchanged Z: increase by 1
Gamma decay	${}^A_ZX \rightarrow {}^0_0\gamma + {}^A_ZY$	 Excited nuclear state	A: unchanged Z: unchanged
Positron emission	${}^A_ZX \rightarrow {}^0_{+1}e + {}^{A}_{Z-1}Y$		A: unchanged Z: decrease by 1
Electron capture	${}^A_ZX + {}^0_{-1}e \rightarrow {}^{A}_{Z-1}Y + \gamma$	 X-ray	A: unchanged Z: decrease by 1

Nuclear Decay Events

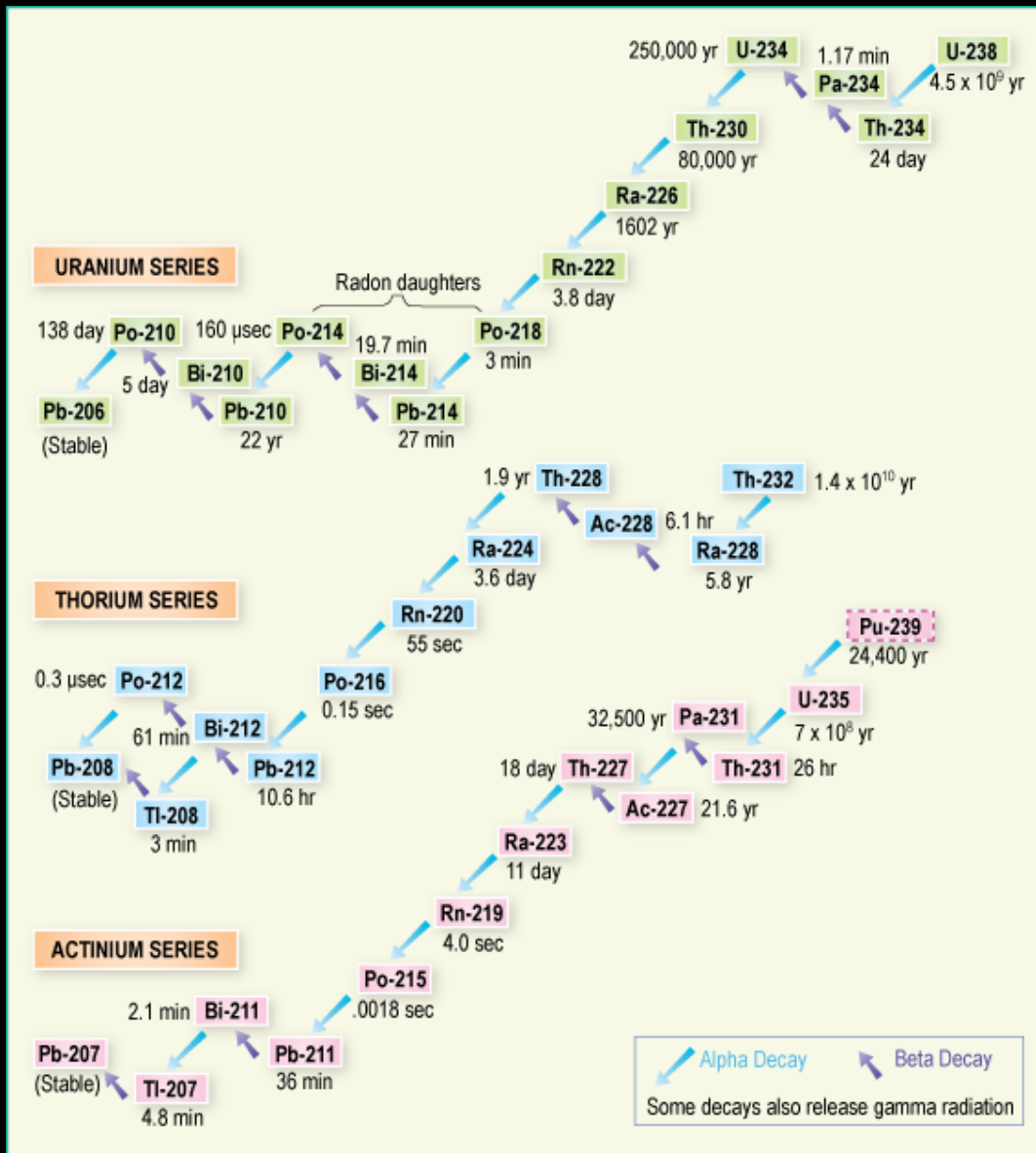


Nuclear Decay Events



Decay alters nucleus
Parent decays
Daughter results

Naturally Occurring Radioactive Decay Series



Chains of successive decays

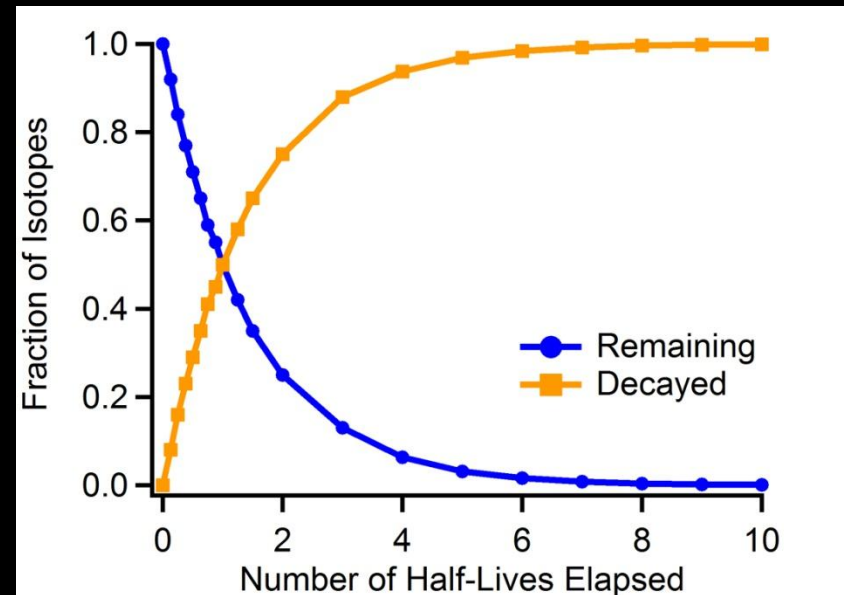
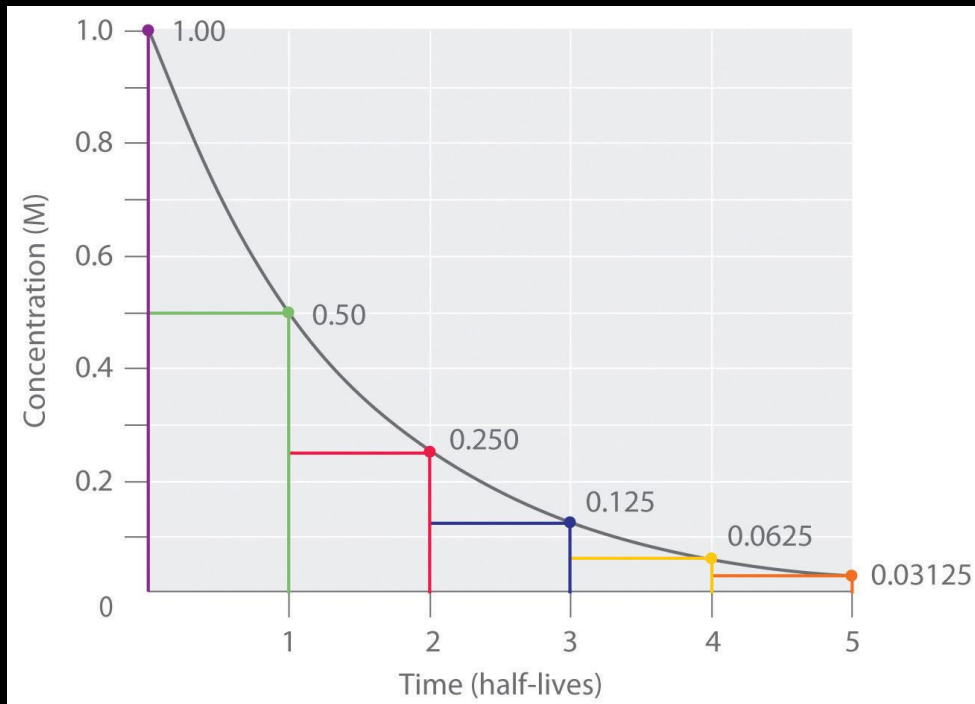
Named for Parent Isotope

All decay to stable lead

Radioactive Half-Lives

Radioactive decay follows first order kinetics

Half-time ($t^{1/2}$): time for one-half of sample to decay



N = number of atoms
n = isotope decay constant
t = time

$$N_t = N_0 e^{-kt}$$

$$t = -\frac{1}{k} \ln \left(\frac{N_t}{N_0} \right)$$

Radioactive Half-Lives



Cobalt -60 decays with a half-life of 5.27 years to produce Ni- 60.

The decay (rate) constant:

$$\lambda = \frac{\ln 2}{t_{1/2}} = \frac{0.693}{5.27 \text{ y}} = 0.132 \text{ y}^{-1}$$

Calculate amount remaining after 15 years:

$$\frac{N_t}{N_0} = e^{-\lambda t} = e^{-(0.132 \text{ y})(15.0 \text{ y})} = 0.138 \rightarrow 13.8 \%$$

How long to decay to 2 % of the original sample

$$t = - \frac{1}{\lambda} \ln \left(\frac{N_t}{N_0} \right) = - (1 / 0.132 \text{ y}^{-1}) \ln (0.0200 N_0 / N_0)$$

$$t = 29.6 \text{ y}$$

Radioactive Half-Life Problem

Radon-222 has a half-life of 3.823 days. How long will it take a sample of radon-222 with a mass of 0.750 g to decay to 0.100 g of the original radon-333

Fraction Remaining:

$$0.100 / 0.750 = 0.1333$$

Determine number of half-times:

$$(1/2)^n = 0.1333$$

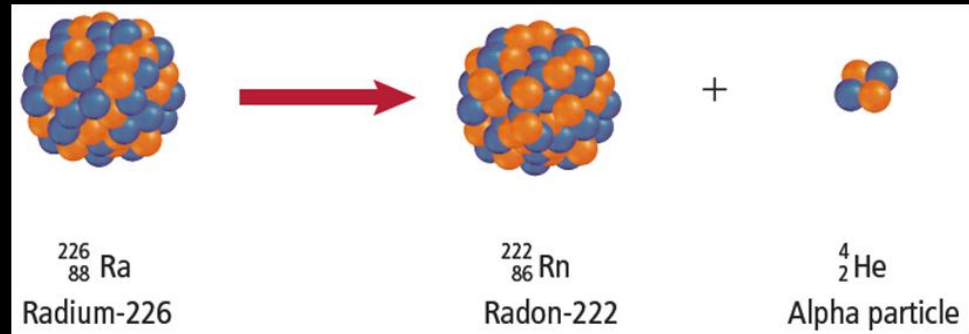
$$n \log (0.5) = \log (0.1333)$$

$$n (-0.301) = -0.876$$

$$n = 2.91$$

Determine time elapsed:

$$2.91 \text{ half-lives} \times \frac{3.823 \text{ days}}{\text{half-live}} = 11.1 \text{ days}$$



Medically Important Radioactive Half-Lives

Type ^[1]	Decay Mode	Half-Life	Uses
F-18	β^+ decay	110. minutes	PET scans
Co-60	β decay, γ decay	5.27 years	cancer treatment
Tc-99m	γ decay	8.01 hours	scans of brain, lung, heart, bone
I-131	β decay	8.02 days	thyroid scans and treatment
Tl-201	electron capture	73 hours	heart and arteries scans; cardiac stress tests

Radioisotopes in the News



**Xe-129 found in Mars atmosphere
Implies ancient nuclear explosion**



**He-3 found on the moon
Potential fusion energy source**



**Sr-90 found in atmosphere
Indicates nuclear processes
Health concern since animals incorporate Sr for Ca**



**Co-60 gamma ray source
Used in radiation therapy to kill cancer cells**

Half-Life Demo



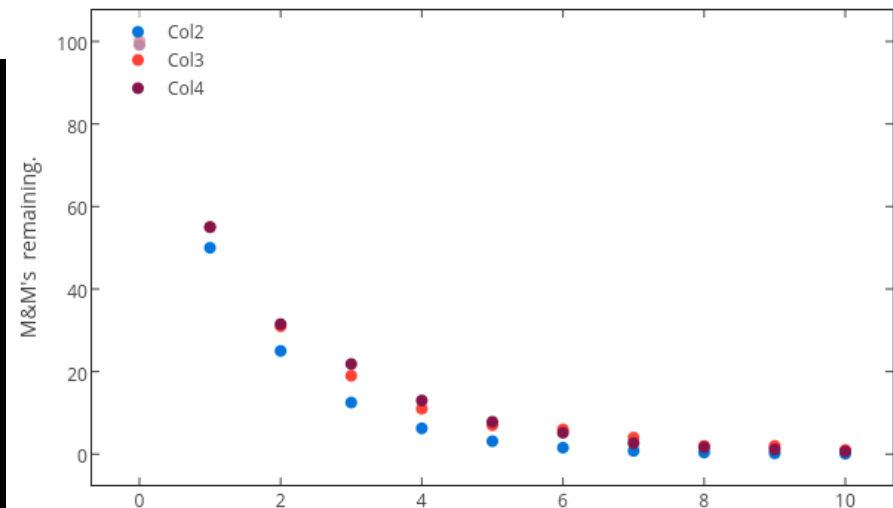
EXPERIMENT

The following experiment is designed to simulate a natural occurrence.

I. Data Collection:

- Quality control: Check each m&m to be sure that it has an "m" on one side *only* – eat any that don't! You may also consume any m&m's that are not whole.
- Count the m&m's in the cup and record as the beginning number in the table at the right.
- Shake the cup and carefully dump the m&m's on the napkin on your desktop. Remove those candies that landed with the "m" showing and eat them! Count the ones remaining and record this number in the chart at the right. Place these candies back into the cup.
- Repeat step c until no m&m's are left.
- Graph the data from the table below – label and use an appropriate scale.

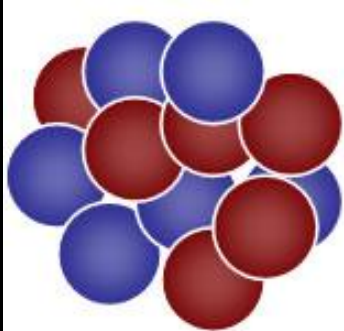
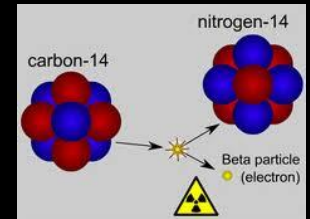
Toss	# remaining
Beginning	
1	
2	
3	
4	
5	
6	
7	
8	
9	
10	
11	
12	



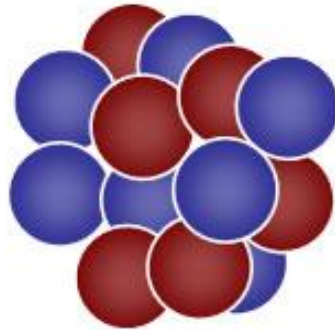
Radiometric Dating

Radioactive decay has become a powerful dating technique

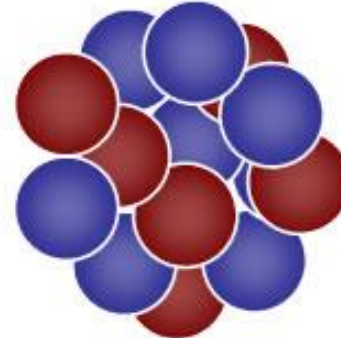
Carbon-14: useful for former living materials
reliable to 30,000 years ago
limit 50,000 years ago



carbon-12
98.9%
6 protons
6 neutrons



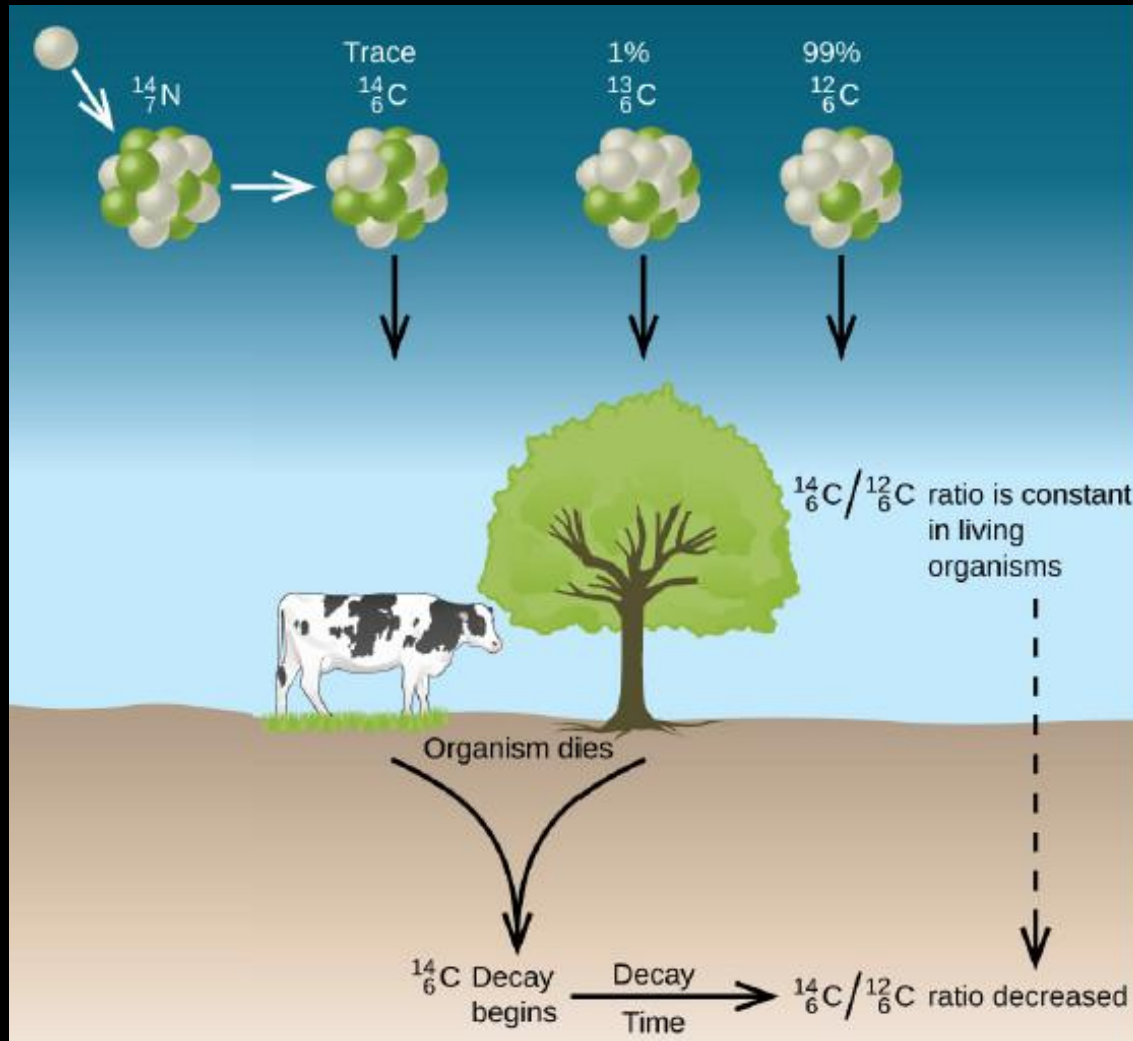
carbon-13
1.1%
6 protons
7 neutrons



carbon-14
<0.1%
6 protons
8 neutrons



Radiometric Dating: Carbon-14



Half-life C-14:
5730 years

Six half-lives:
34,380 years

During Life:
C-14 Constant
C-12

After death
As C-14 decays
C-14 Decreases
C-12

Ratio $\sim$ Age

Increasing CO_2 levels may affect C-14 data; validate with other methods

Radiometric Dating: Carbon-14

A piece of the dead sea scrolls has an activity of 10.8 disintegrations per gram of C-14. If initial activity was 13.6 disintegrations per minute, estimate scroll age.

Determine λ :

$$\lambda = \ln 2 / t_{1/2}$$

$$\lambda = 0.693 / 5730 \text{ y}$$

$$\lambda = 1.21 \times 10^{-4} \text{ y}^{-1}$$

Estimate age:

$$t = -1/\lambda \ln (\text{rate}_t / \text{rate}_0)$$

$$t = - (1/ 1.21 \times 10^{-4} \text{ y}^{-1}) \ln (10.8 \text{ dis/min/g C} / 13.6 \text{ dis/min/g C})$$

$$t = 1905 \text{ years} \rightarrow 1910 \text{ years}$$



Radiometric Dating:

Seeds from Egyptian tombs have a C-14 disintegration rate of 9.07 dis/min/g C. Determine the age of the seeds

Determine λ :

$$\lambda = \ln 2 / t_{1/2}$$

$$\lambda = 0.693 / 5730 \text{ y}$$

$$\lambda = 1.21 \times 10^{-4} \text{ y}^{-1}$$

Estimate age:

$$t = -1/\lambda \ln (\text{rate}_t / \text{rate}_0)$$

$$t = - (1/ 1.21 \times 10^{-4} \text{ y}^{-1}) \ln (9.07 \text{ dis/min/g C} / 13.6 \text{ dis/min/g C})$$

$$t = 3347 \text{ years} \rightarrow 3350 \text{ years}$$



Radiometric Dating: Other nucleotides

Isotope selected depends on date range needed

ISOTOPES		HALF-LIFE OF PARENT (YEARS)	EFFECTIVE DATING RANGE (YEARS)	MINERALS AND OTHER MATERIALS THAT CAN BE DATED
PARENT	DAUGHTER			
Uranium-238	Lead-206	4.5 billion	10 million- 4.6 billion	Zircon Uraninite
Potassium-40	Argon-40	1.3 billion	50,000 - 4.6 billion	Muscovite Biotite Hornblende Whole volcanic rock
Rubidium-87	Strontium-87	47 billion	10 million - 4.6 billion	Muscovite Biotite Potassium feldspar Whole metamorphic or igneous rock
Carbon-14	Nitrogen-14	5730	100 -70,000	Wood, charcoal, peat Bone and tissue Shell and other calcium carbonate Groundwater, ocean water, and glacier ice containing dissolved carbon dioxide

Radiometric Dating: Rocks

Ratio of parent / daughter used to estimate age

Rock sample contains 9.58×10^{-5} g of U-235 and 2.51×10^{-5} g of Pb-206.

Determine the age of the rock. (Assume all Pb-206 produced by decay of U-238.)

Amount of U-238 currently in the rock:

$$9.58 \times 10^{-5} \text{ g U} \times \frac{1 \text{ mol U}}{238 \text{ g U}} = 4.03 \times 10^{-7} \text{ mole U-238}$$

Amount of Pb produced converted to moles uranium:

$$2.51 \times 10^{-5} \text{ g Pb} \times \frac{1 \text{ mol Pb}}{206 \text{ g Pb}} \times \frac{1 \text{ mol U-238}}{1 \text{ mol Pb}} = 1.22 \times 10^{-7} \text{ mole U-238}$$

Total U-238 present in original sample:

$$(4.03 + 1.22) \times 10^{-7} \text{ mole} = 5.25 \times 10^{-7} \text{ mole U-238}$$

Amount of time passed for this decay process:

$$t = -1/\lambda \ln (N_t/N_0)$$

$$t = (1/1.54 \times 10^{-10} \text{ y}^{-1}) \ln \left(\frac{4.03 \times 10^{-7} \text{ mol U}}{5.25 \times 10^{-7} \text{ mol U}} \right)$$

$$t = 1.7 \times 10^9 \text{ years}$$

Need Lamda:

$$\lambda = \ln 2 / 4.5 \times 10^9 \text{ y}$$

$$\lambda = 1.54 \times 10^{-10} \text{ y}^{-1}$$

Problem: Radiometric Dating of Rocks

Rock sample contains 6.14×10^{-4} g of Rb-87 and 3.51×10^{-5} g of Sr-87.

Determine the age of the rock. (Half-life of Rb-87 = 4.7×10^{10} y.)

Amount of Rb-87 currently in the rock:

$$6.14 \times 10^{-4} \text{ g Rb} \times \frac{1 \text{ mol Rb}}{87 \text{ g Rb}} = 7.06 \times 10^{-6} \text{ mole Rb}$$

Amount of Sr produced converted to moles Rb:

$$3.51 \times 10^{-5} \text{ g Sr} \times \frac{1 \text{ mol Sr}}{87 \text{ g Sr}} \times \frac{1 \text{ mol Rb}}{1 \text{ mol Sr}} = 4.03 \times 10^{-7} \text{ mole Rb}$$

Total Rb present in original sample:

$$7.06 \times 10^{-6} + 4.03 \times 10^{-7} \text{ mole} = 7.46 \times 10^{-6} \text{ mole Rb}$$

Amount of time passed for this decay process:

$$t = -1/\lambda \ln (N_t/N_0)$$

$$t = - (1/1.47 \times 10^{-11} \text{ y}^{-1}) \ln \left(\frac{7.06 \times 10^{-6} \text{ mol Rb}}{7.46 \times 10^{-6} \text{ mol Rb}} \right)$$

$$t = 3.7 \times 10^9 \text{ years}$$

Need Lamda:

$$\lambda = \ln 2 / 4.7 \times 10^{10} \text{ y}$$

$$\lambda = 1.47 \times 10^{-11} \text{ y}^{-1}$$

Transmutation of Elements

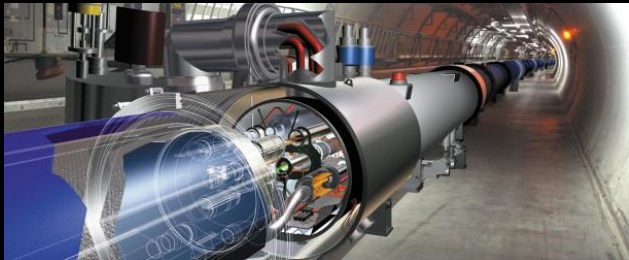
Transmutation:

Changing one nucleotide to another via nuclear reactions

Requires enormous kinetic energy:

Originally called “atom smashers”

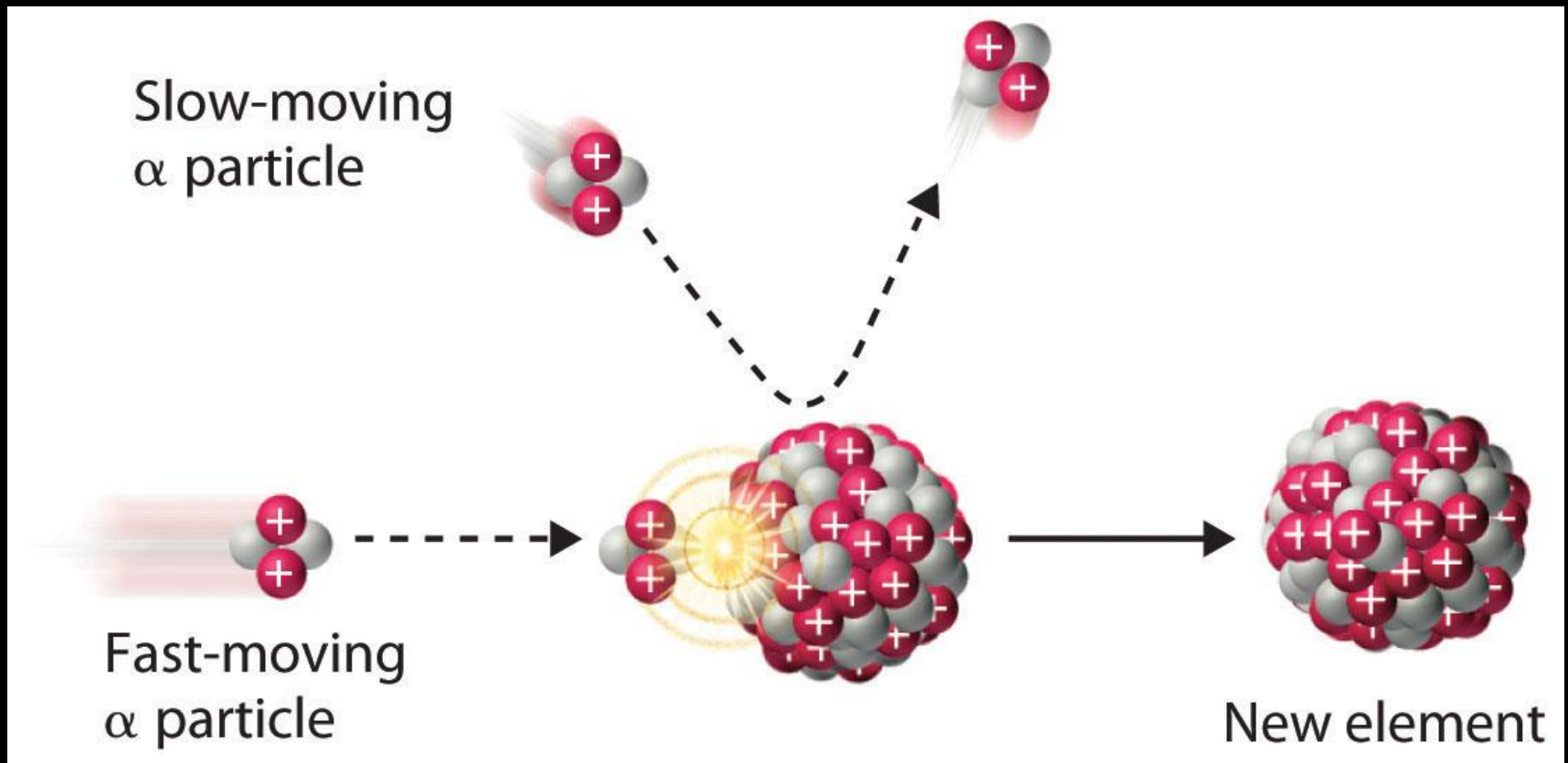
Now: particle accelerators or colliders



Transmutation of Elements

Requires enormous kinetic energy

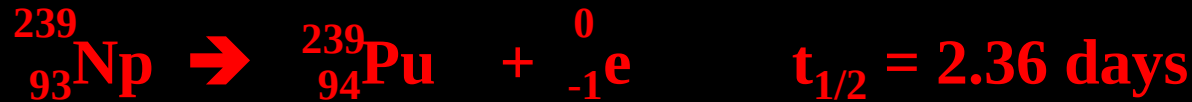
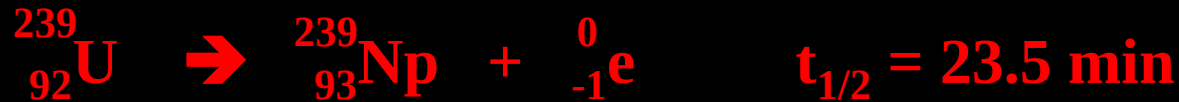
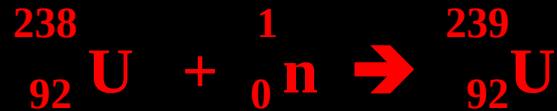
Must overcome forces holding nucleus together



Transmutation of Elements

All elements beyond Uranium are man-made

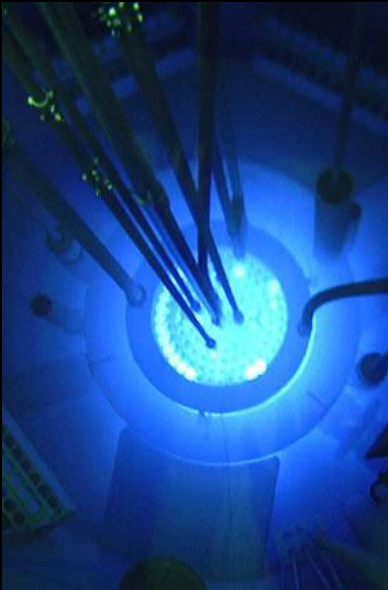
First man-made element: Neptunium (Neptune follows Uranus)



Most Plutonium created in “breeder reactors” using U-238



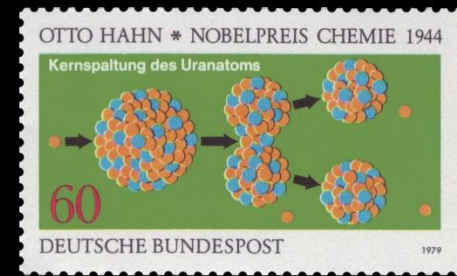
Transmutation of Elements



Name	Symbol	Atomic Number	Reaction
americium	Am	95	${}^{239}_{94}\text{Pu} + {}^1_0\text{n} \longrightarrow {}^{240}_{95}\text{Am} + {}^0_{-1}\text{e}$
curium	Cm	96	${}^{239}_{94}\text{Pu} + {}^4_2\text{He} \longrightarrow {}^{242}_{96}\text{Cm} + {}^1_0\text{n}$
californium	Cf	98	${}^{242}_{96}\text{Cm} + {}^4_2\text{He} \longrightarrow {}^{243}_{97}\text{Bk} + 2{}^1_0\text{n}$
einsteinium	Es	99	${}^{238}_{92}\text{U} + 15{}^1_0\text{n} \longrightarrow {}^{253}_{99}\text{Es} + 7{}^0_{-1}\text{e}$
mendelevium	Md	101	${}^{253}_{99}\text{Es} + {}^4_2\text{He} \longrightarrow {}^{256}_{101}\text{Md} + {}^1_0\text{n}$
nobelium	No	102	${}^{246}_{96}\text{Cm} + {}^{12}_6\text{C} \longrightarrow {}^{254}_{102}\text{No} + 4{}^1_0\text{n}$
rutherfordium	Rf	104	${}^{249}_{98}\text{Cf} + {}^{12}_6\text{C} \longrightarrow {}^{257}_{104}\text{Rf} + 4{}^1_0\text{n}$
seaborgium	Sg	106	${}^{249}_{98}\text{Cf} + {}^{18}_8\text{O} \longrightarrow {}^{263}_{106}\text{Sg} + 4{}^1_0\text{n}$
meitnerium	Mt	107	${}^{209}_{83}\text{Bi} + {}^{58}_{26}\text{Fe} \longrightarrow {}^{266}_{109}\text{Mt} + {}^1_0\text{n}$

[illegible]

Nuclear Fission

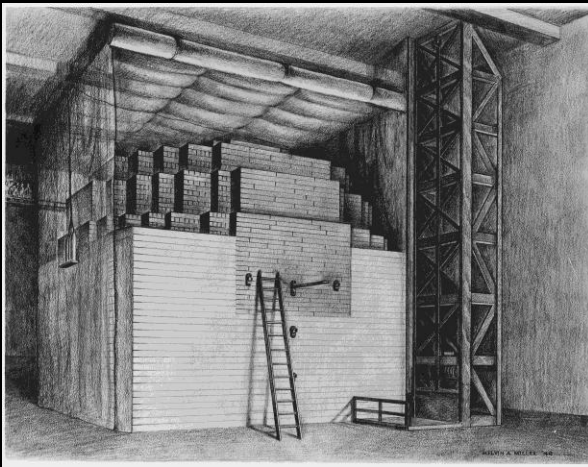


Fission (splitting):

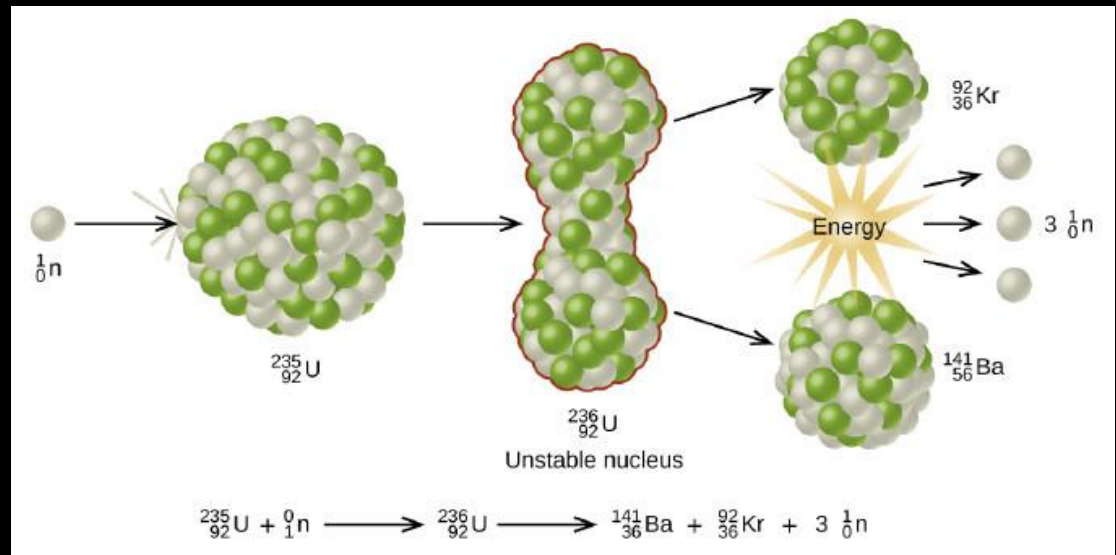
Heavy element nuclei with small nuclear binding energy spontaneously decompose to nuclei with smaller binding energy

Typically, a random, low probability event

First artificial fission (Hahn, 1939)



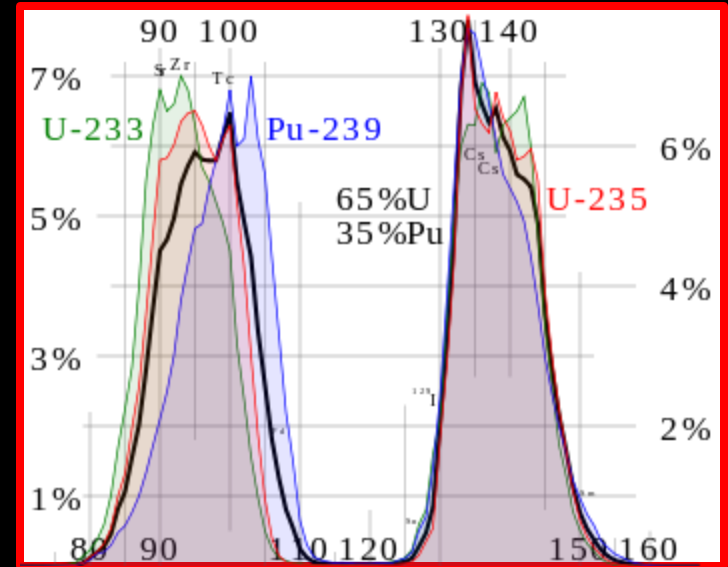
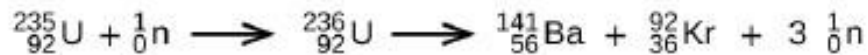
12/2/1942



Process releases large amount of energy

1 kg U-235: 2.5×10^6 more energy than 1 kg coal

Nuclear Fission: Typical products

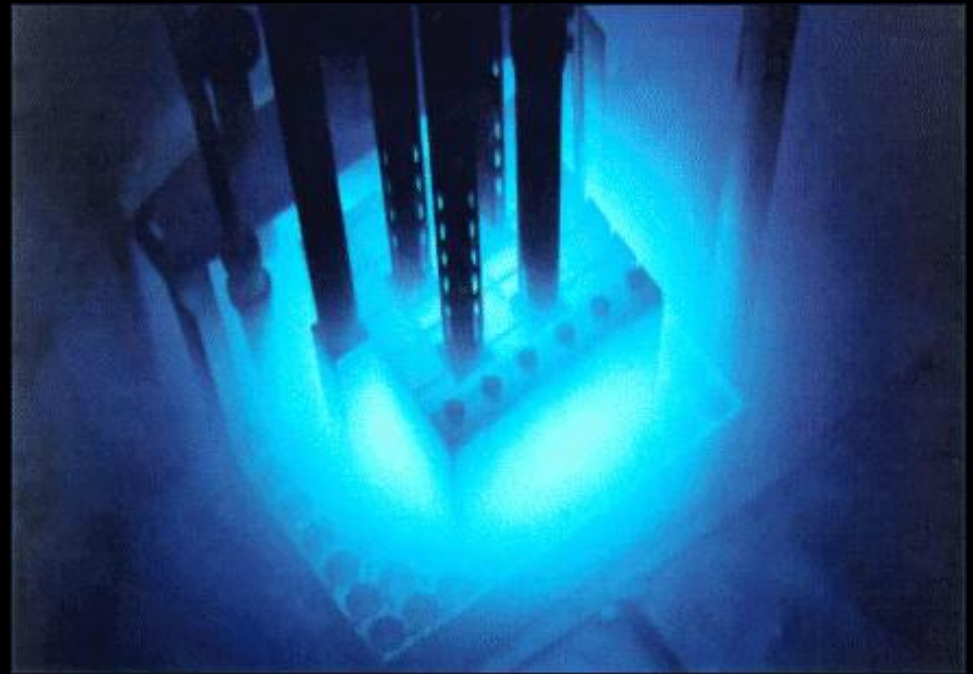
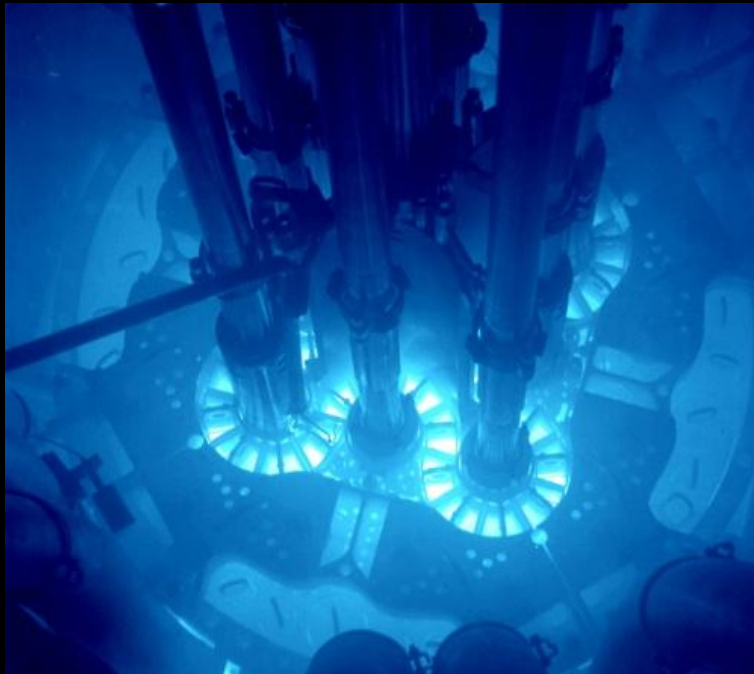
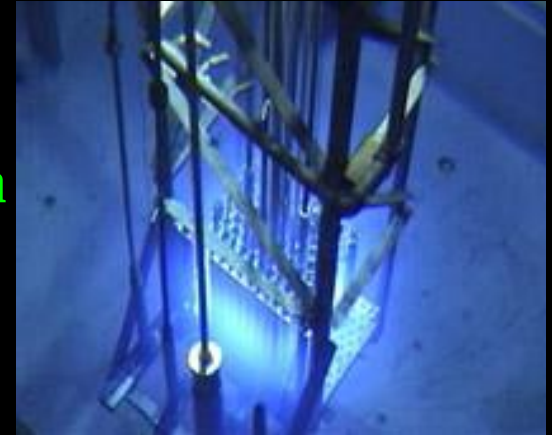


Fission can produce elements from Zn to Lanthanides
Most appear in two peaks: 85-105 and 130-150



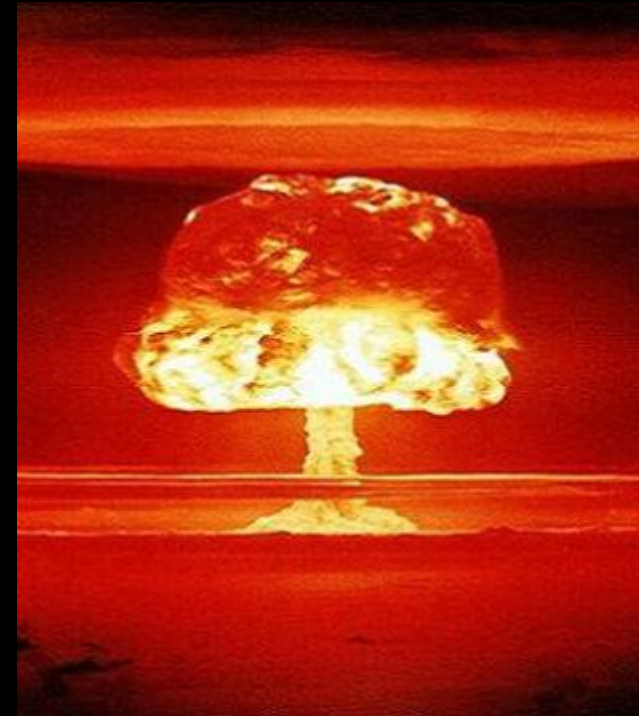
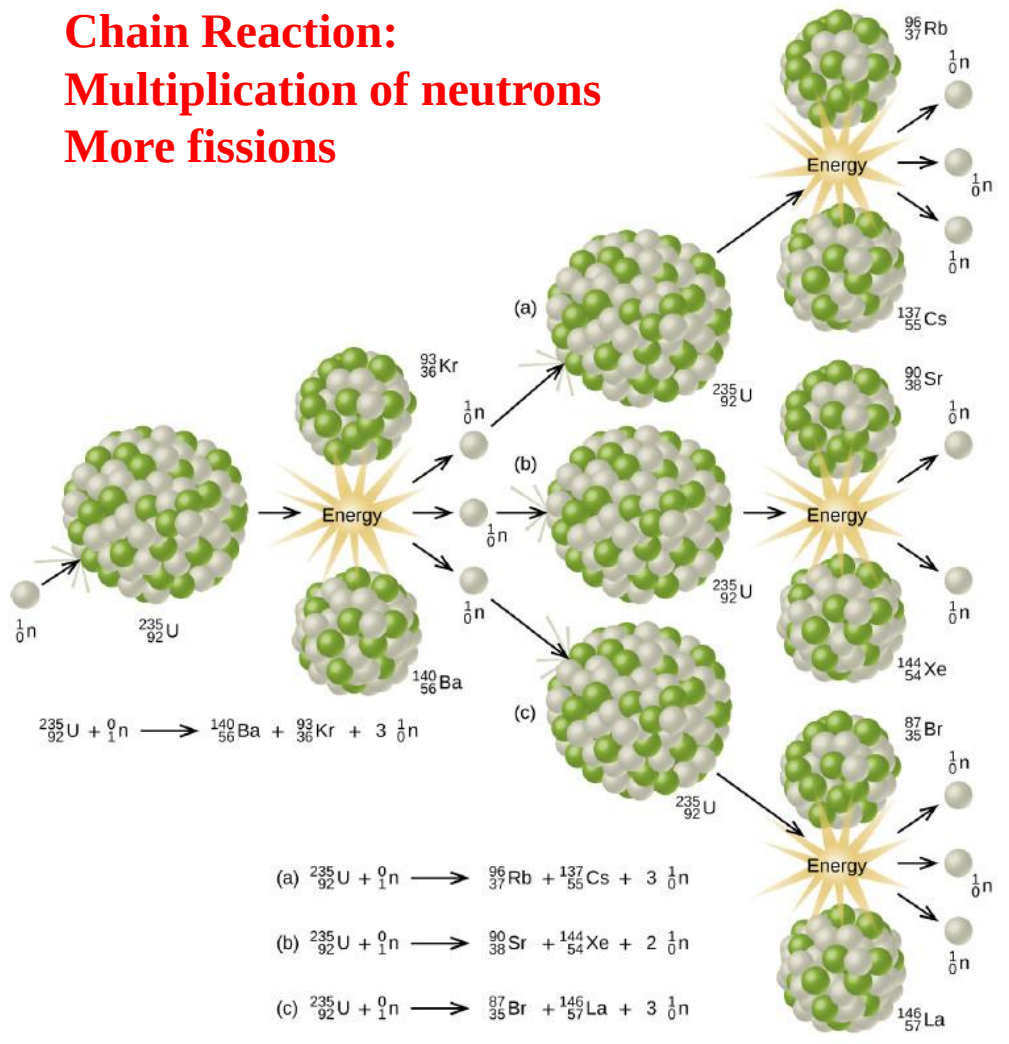
Cherenkov Radiation

Blue glow that results from electrically charged particle moving thru media faster than speed of light in that media



Nuclear Fission: Chain Reaction

Chain Reaction:
Multiplication of neutrons
More fissions



Releases enormous (explosive) energy

Nuclear Fission: Definitions

Fissile: material that will undergo fission as a result of any neutron source

U-233 (from Thorium) U-235 Pu-239 Pu-241

Most actinides with odd number of neutrons

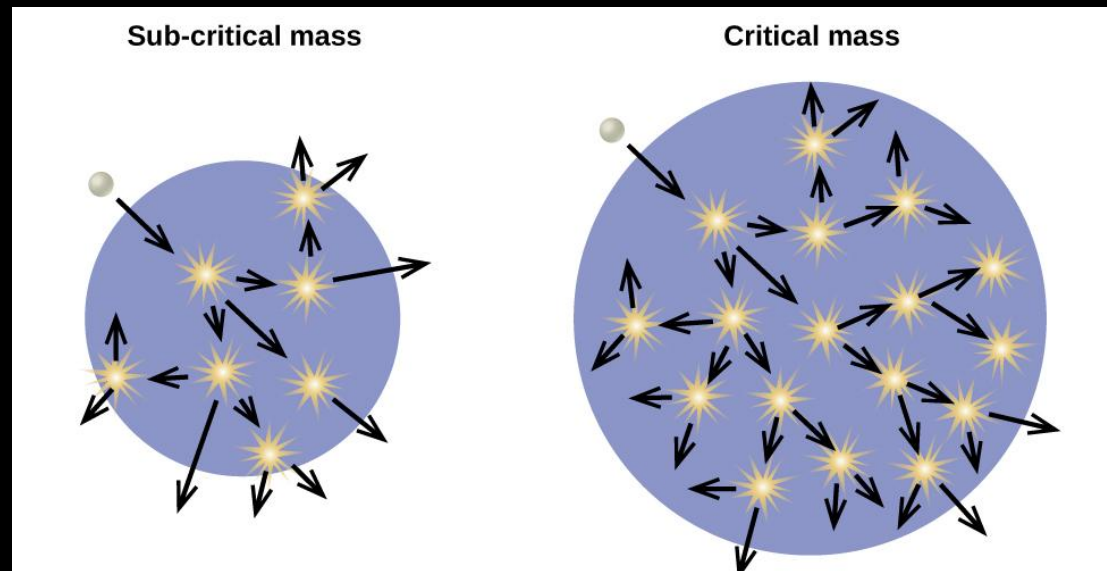
Fissionable: material that requires a high energy neutron source.

Pu-240 U-238

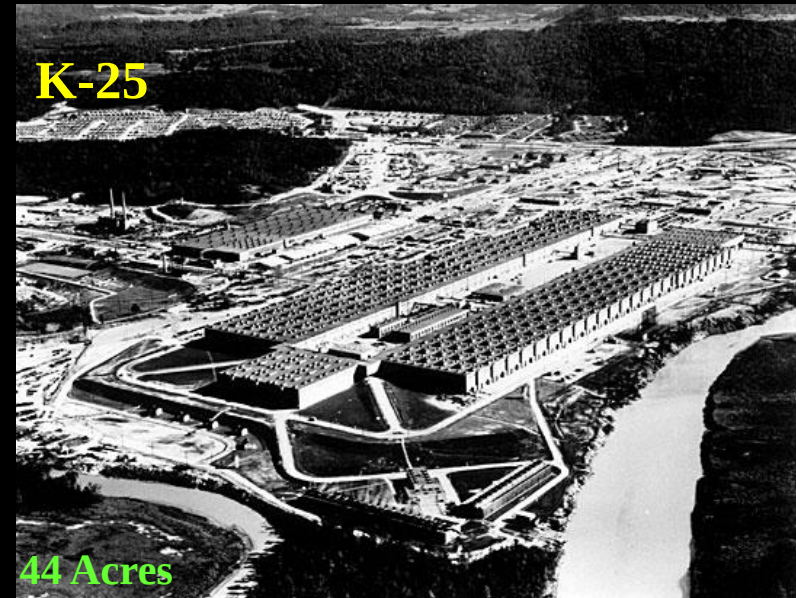
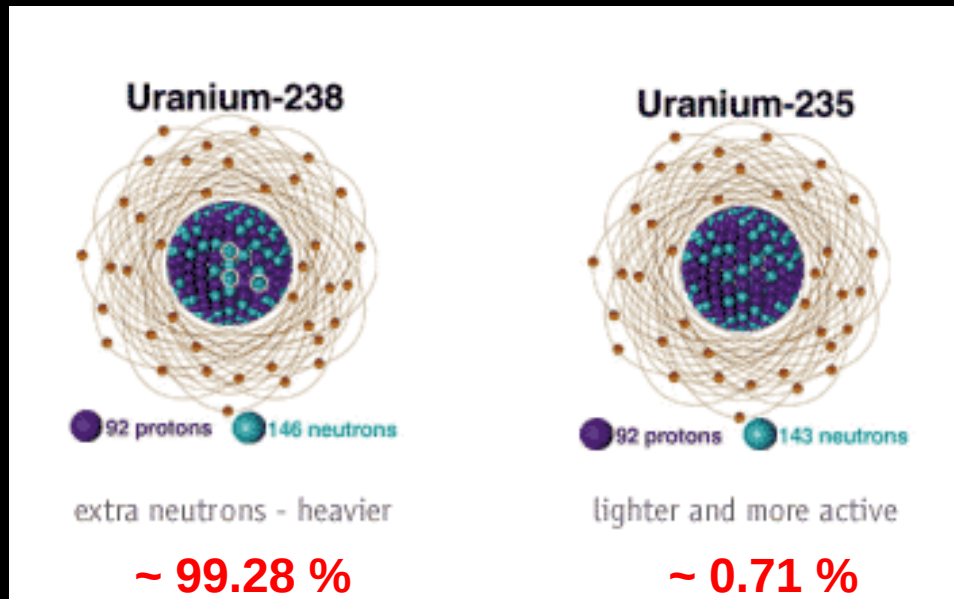
Fission is self-sustaining when neutrons produced exceed neutrons absorbed

Critical Mass: amount of material needed to sustain a chain reaction

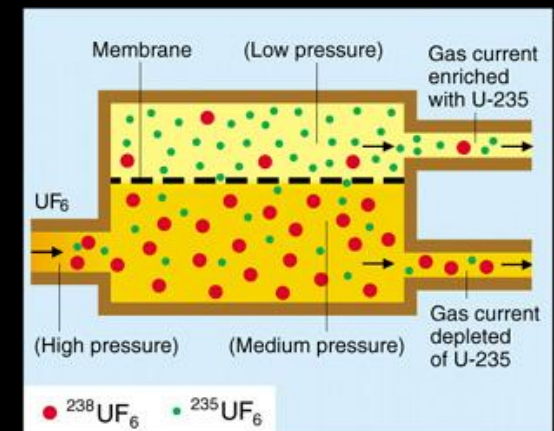
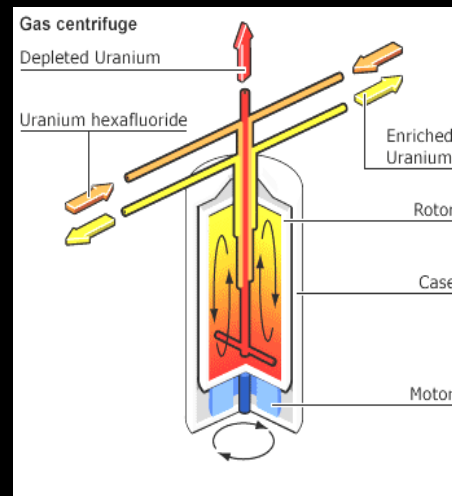
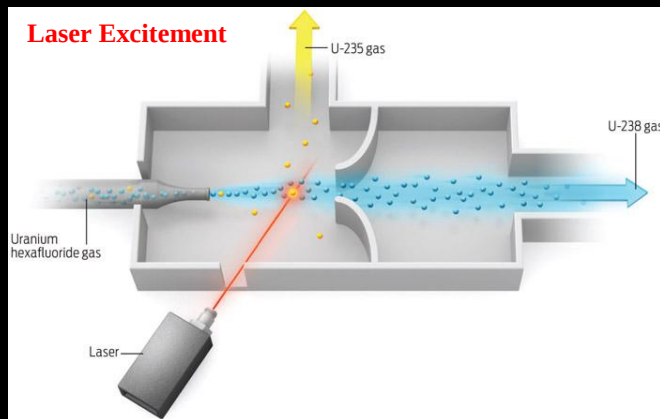
Sub-critical mass: less than critical mass (not self-sustaining)



Isotopes of Uranium



Since Z is the same, chemical properties of the isotopes are identical

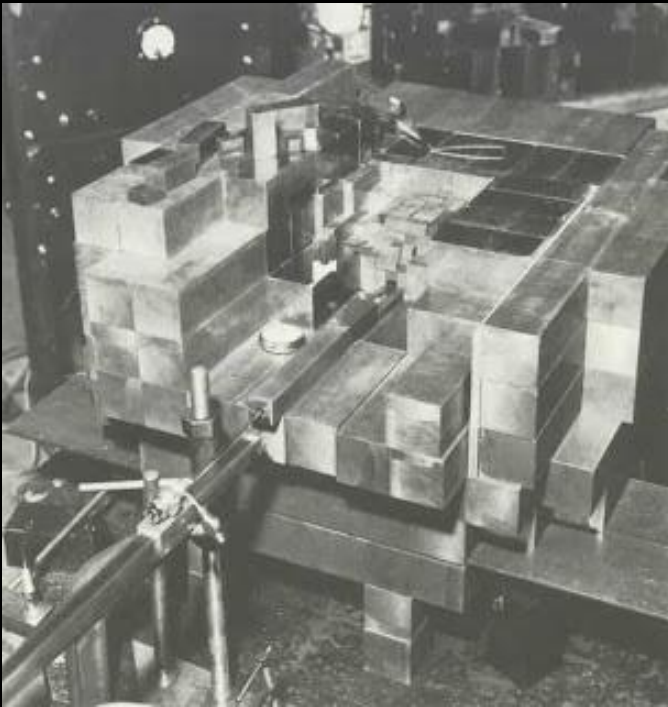


Tickling The Dragon's Tail

Manhattan Project Experiment to determine critical mass
Sub-critical masses brought together on a rail

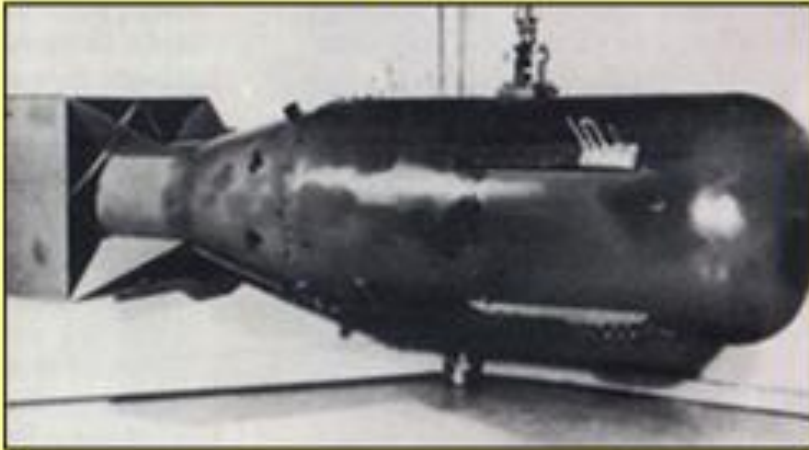
Device failed, Louis Sloton held Uranium spheres apart
Died 5 days latter

Refused medications to allow detailed observations



Atomic Bomb

Little boy : Hiroshima

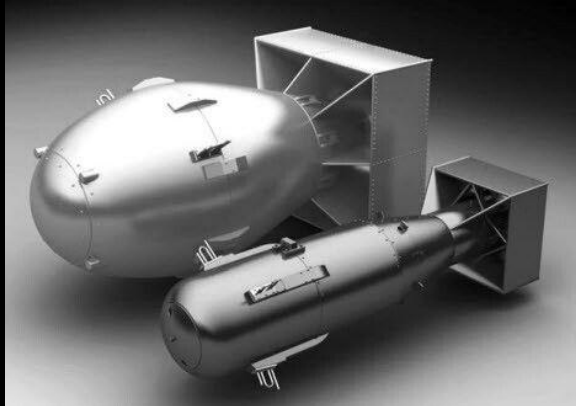


Fission uranium-235
Weight : 4400 kg
Power : 15,000 tons of TNT

Fat Man : Nagasaki

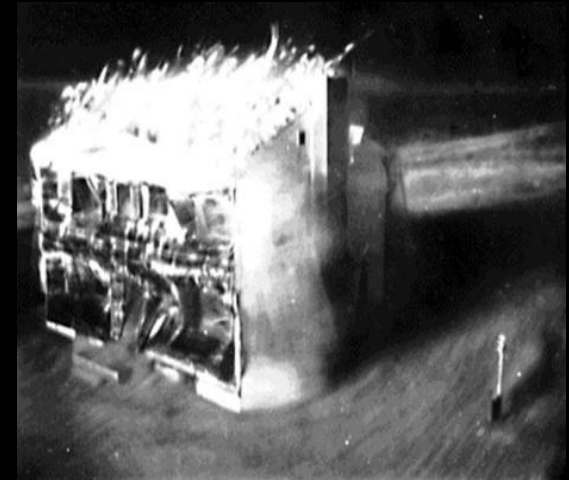
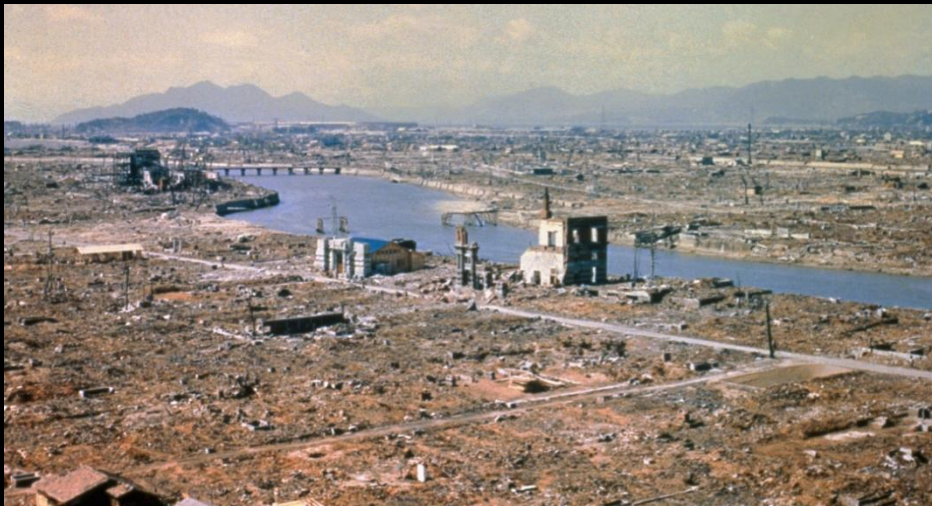
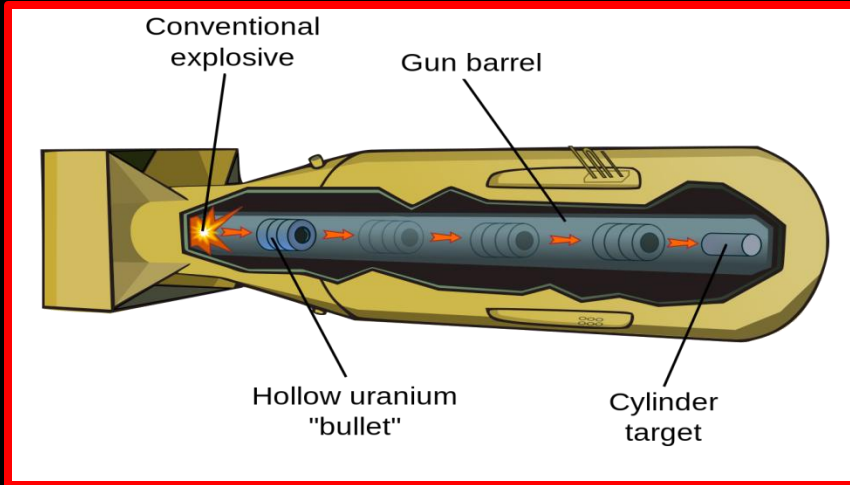


Fission plutonium-239
Weight: 4535 kg
Power : 21,000 tonnes de TNT



Enola Gay

Atomic Bomb



Fission Reactors

Nuclear reactions that is controlled and sustained can be harnessed for power

Reactors control fission of U-235 or Pu 239 with 5 components:

Nuclear fuel

Moderator (slows neutrons)

Reactor coolant

Control rods

Radiation shield and containment

Nuclear fuel is separated (diluted) to prevent chain reaction

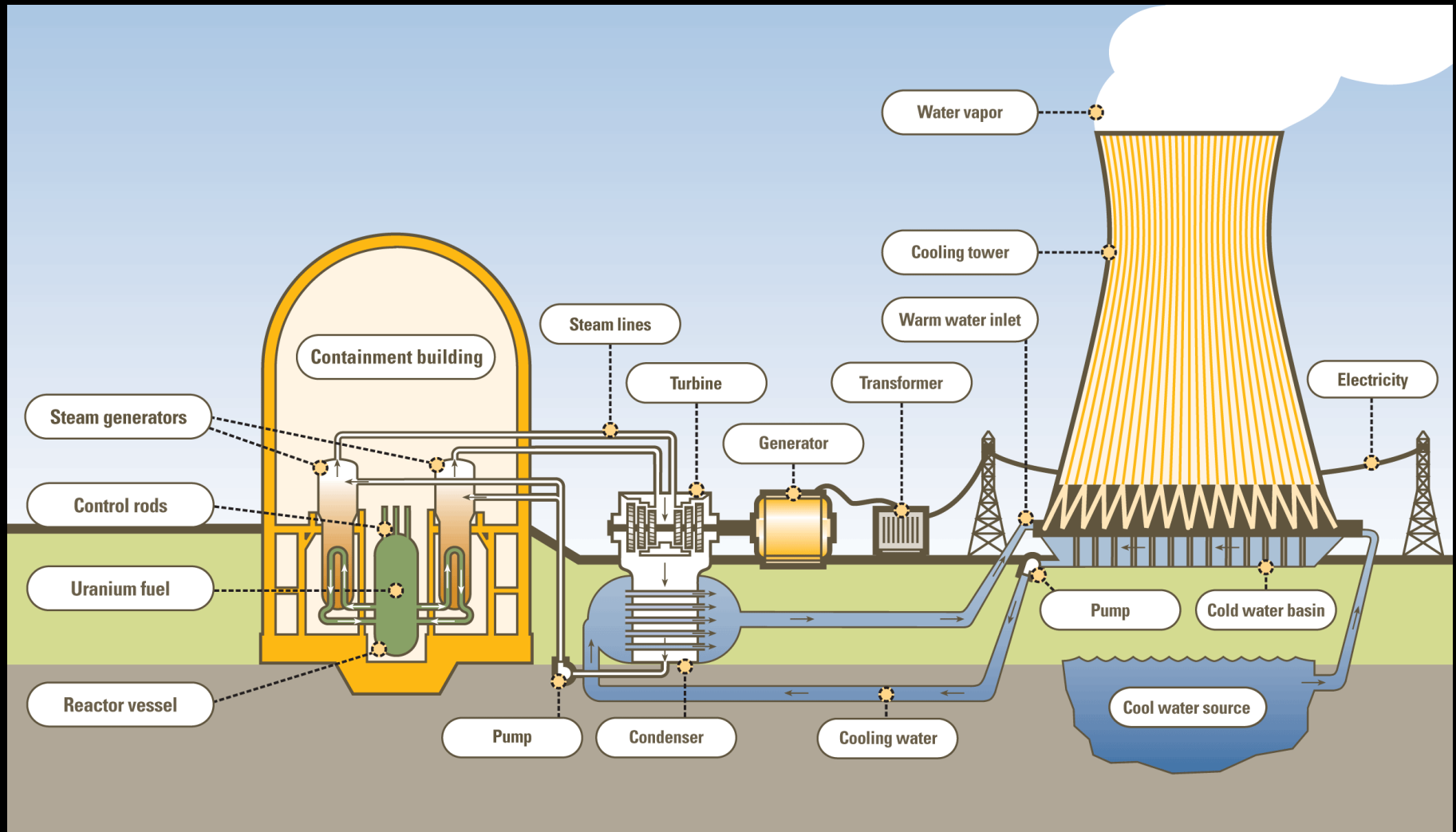
Neutron flux controlled by neutron absorbing material

Heat released used to boil water

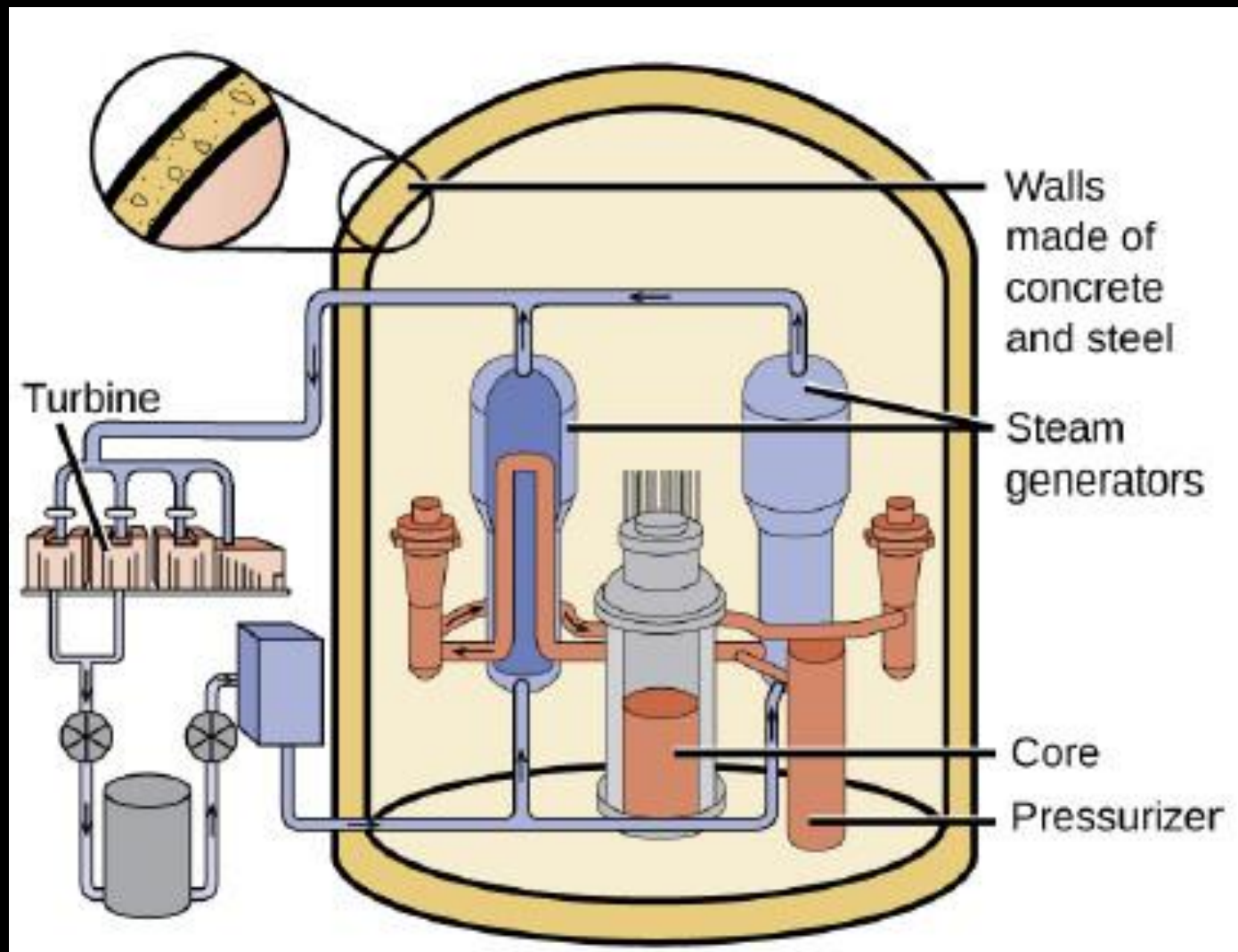
Steam turns turbine and generator



Power Plant Schematic



Fission Reactor Schematic



Nuclear Fuel Rods

Need fissionable material to sustain chain reaction

U-235

Uranium ore: $\sim 0.05 - 0.3\%$ U-235

Fuel rods: enriched uranium $\sim 5\%$ U-235

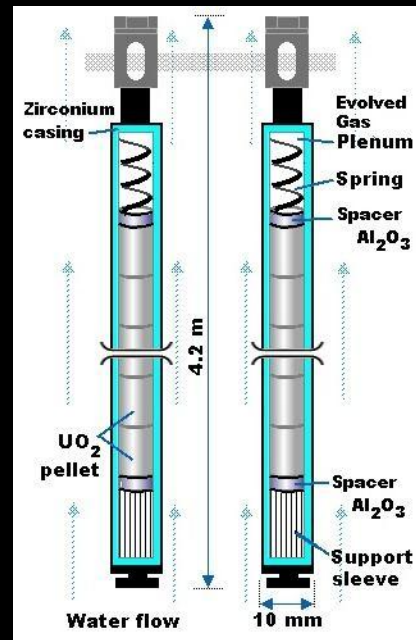
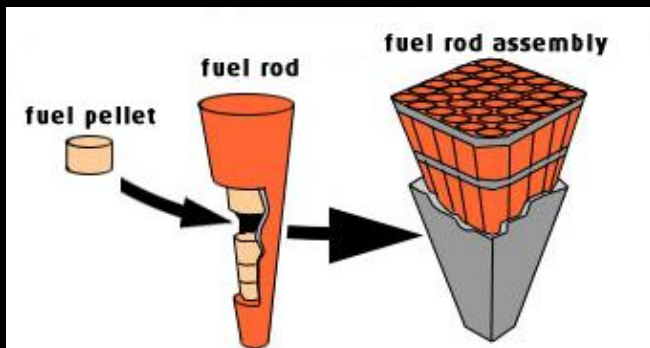
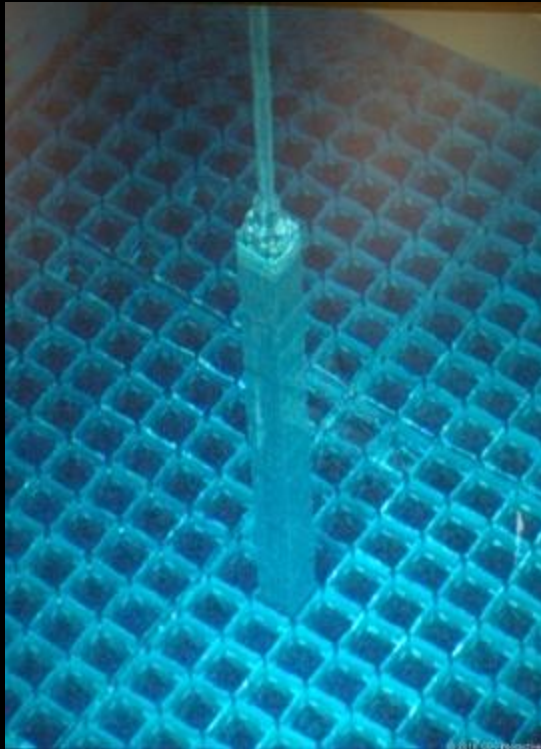
Well below level needed for atomic weapons

Eventually, consumed fuel drops efficiency

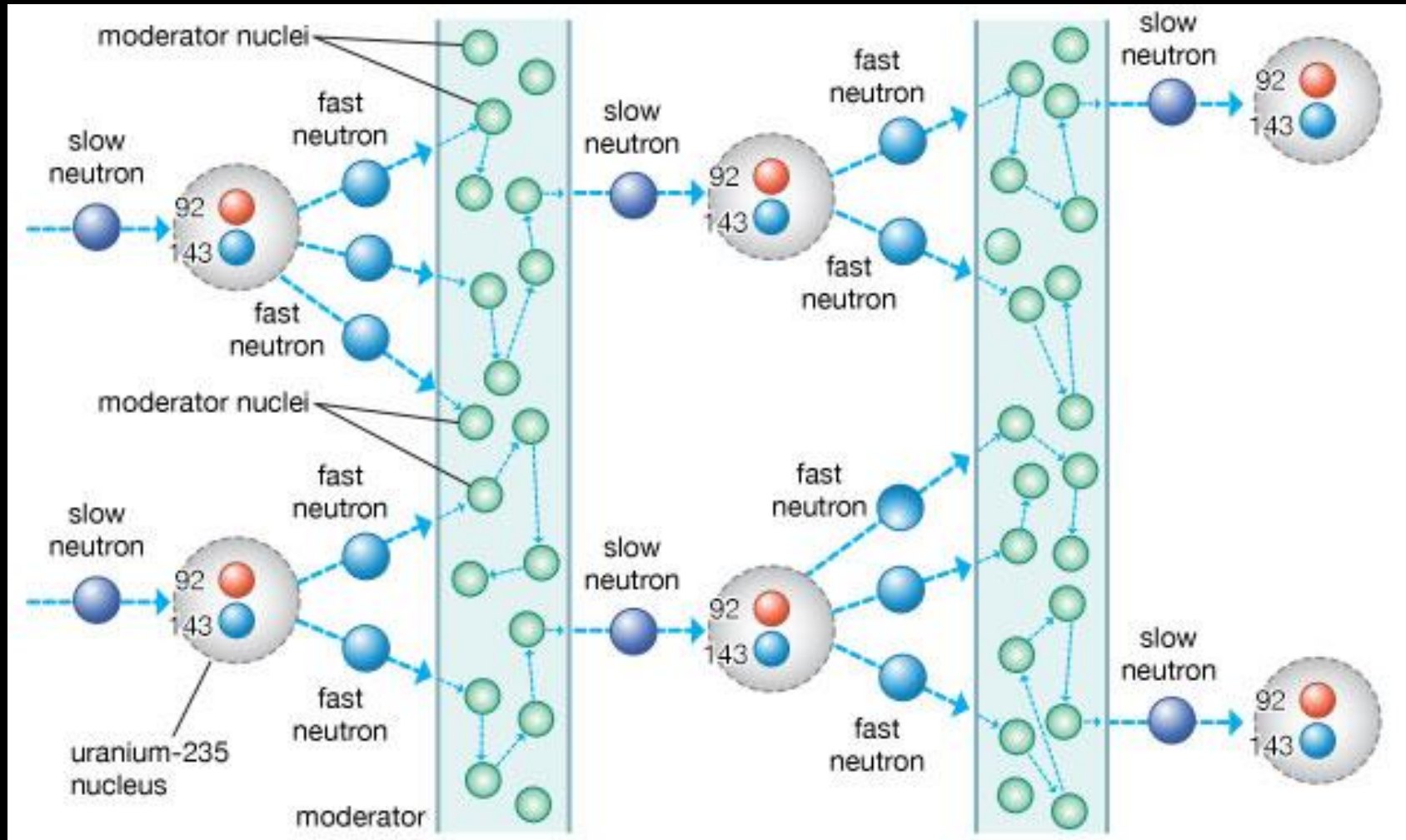
Depleted rods stored (thousands of years to decay)

Some material used for “depleted uranium” artillery shells

Also used to isolate a variety of radio-isotopes



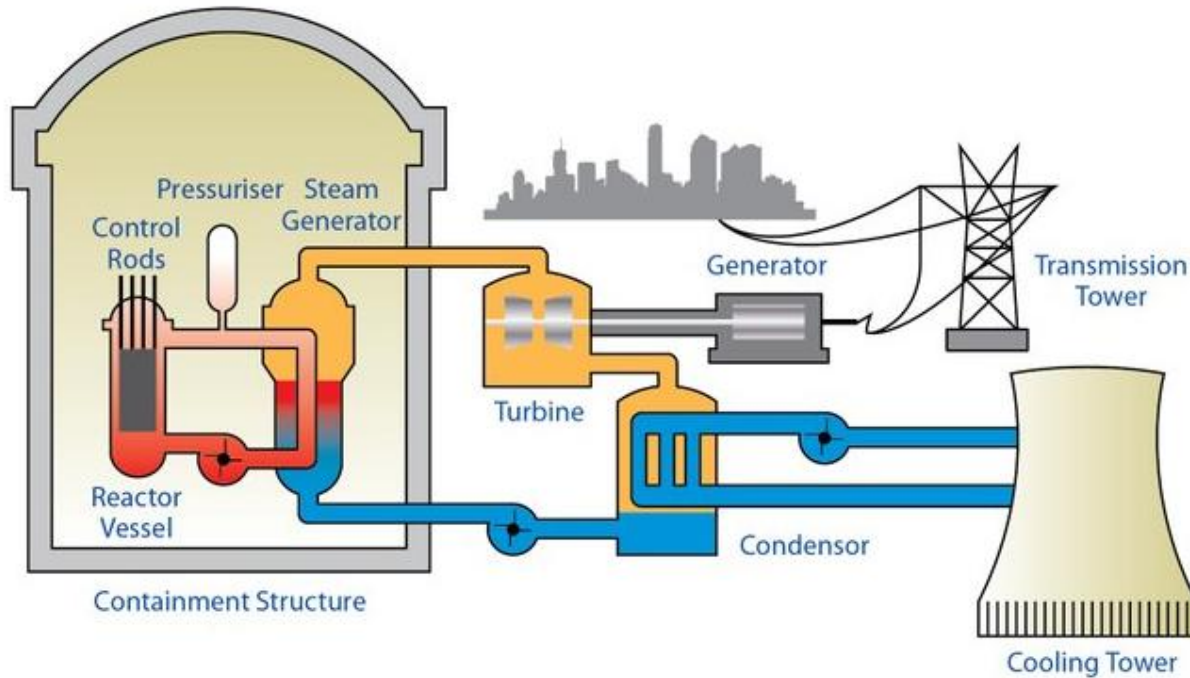
Nuclear Moderator



Slows neutron speed: reaction more efficient

Graphite, Be, CO₂, H₂O, D₂O

Reactor Coolant



Removes heat from reactor

Two/Three loops:

From reactor to boiler

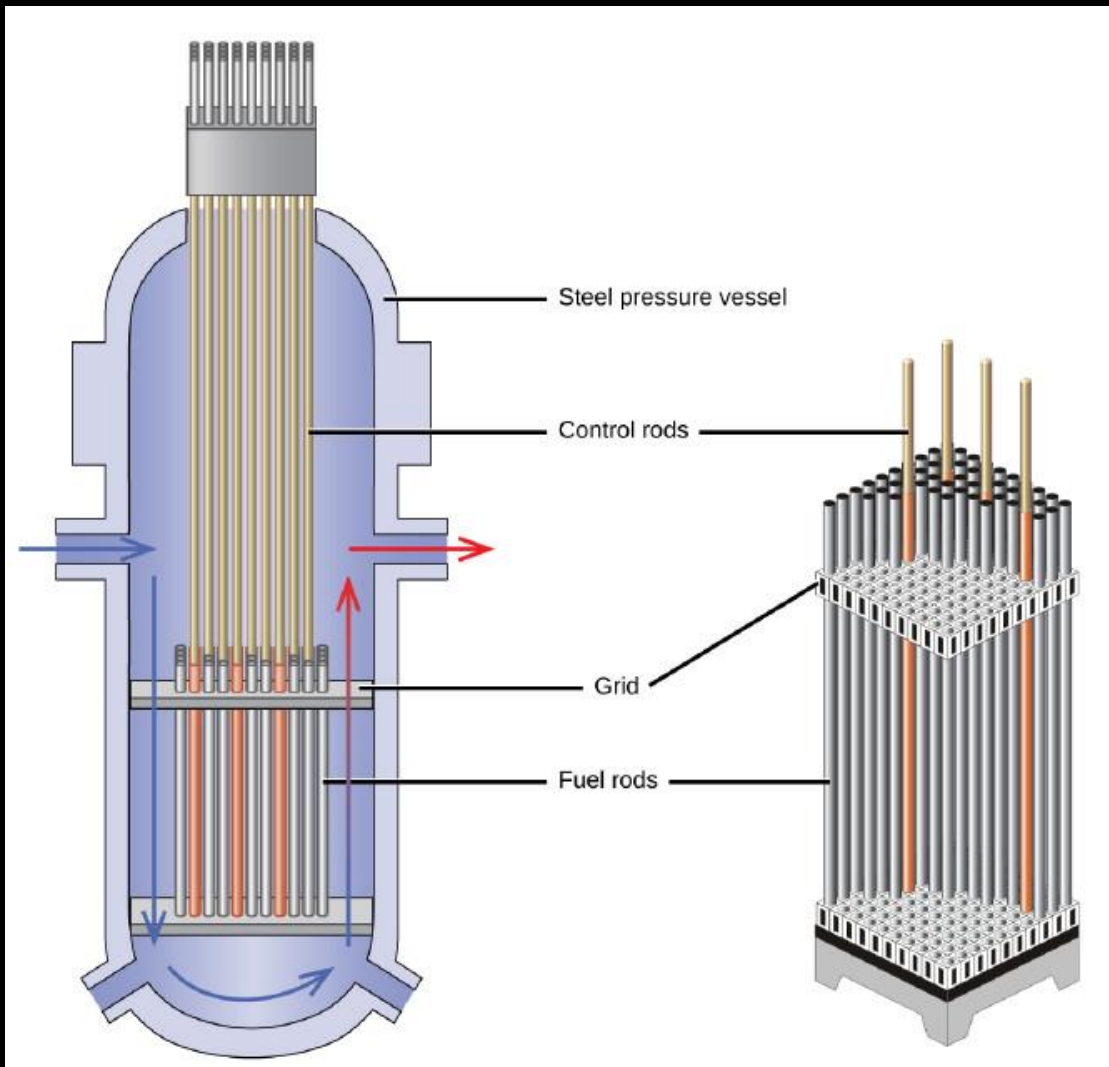
From boiler (steam) to turbine

From Turbine to Tower

H₂O, Na, Pb-Bi, molten salts



Control Rods

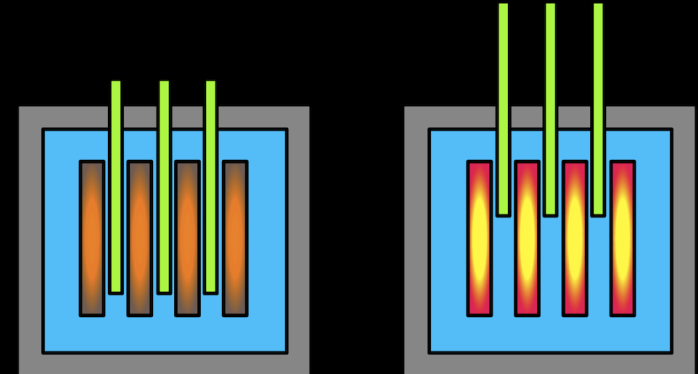


Neutron absorbing materials

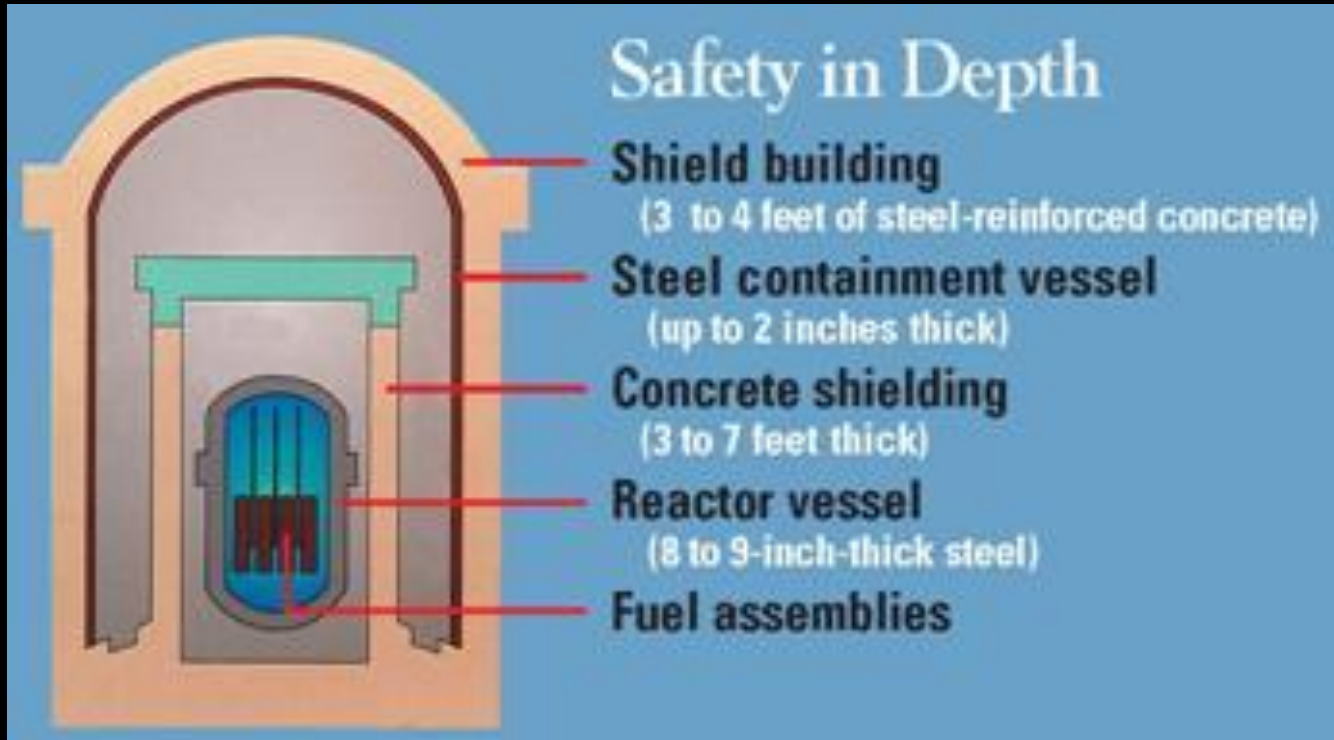
B, Cd, Hf

Control neutron flux

Control amount of fission



Shield and Containment System



Reactors contain and produce a variety of radioactive materials
Shielding and Containment: protects operating personnel

Reactors in USA must have containment shells

Nuclear Accidents: Three Mile Island

Nuclear power stations do not explode
Accidents release radiation into environment

March 1979

Cooling pump failed

Water flooded containment building

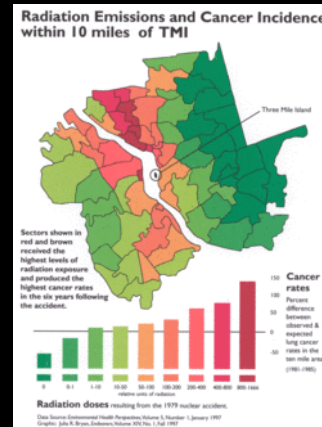
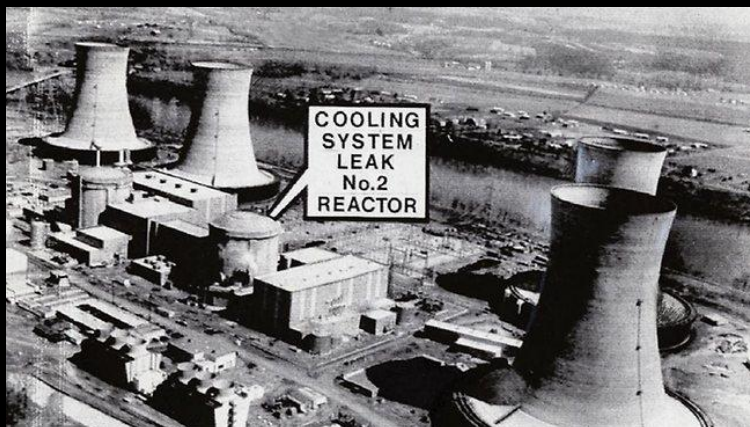
Water superheated

Produced H₂ gas

Gases (H₂, Kr, Xe) vented to atmosphere

Took 10 years to clean

Nothing “hazardous” was released



Nuclear Accidents: Chernobyl

April, 1986

Unauthorized experiment

Minimum safety devices

Reactor became uncontrolled

Superheated water ruptured reactor

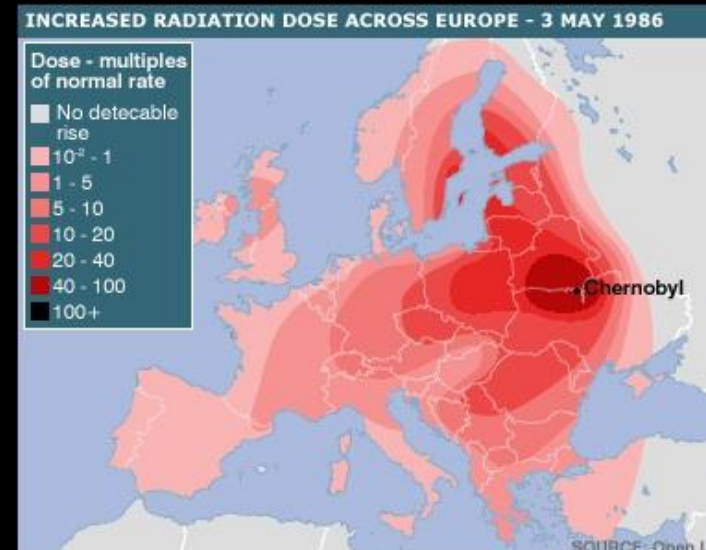
Massive release of radioactivity

Graphite moderator ignited

Thousands died

Area is still a wasteland

Russian reactors have no containment



Nuclear Accidents: Fukushima

March, 2011

9.0 Earthquake, followed by Tsunami

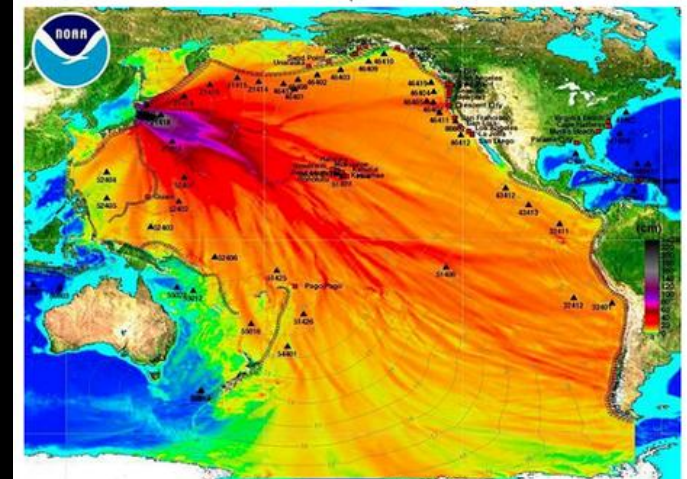
3 reactor meltdowns

Hydrogen explosions / fire

~ 200,000 people displaced



Fukushima nuclear radiation spread across the Pacific Ocean

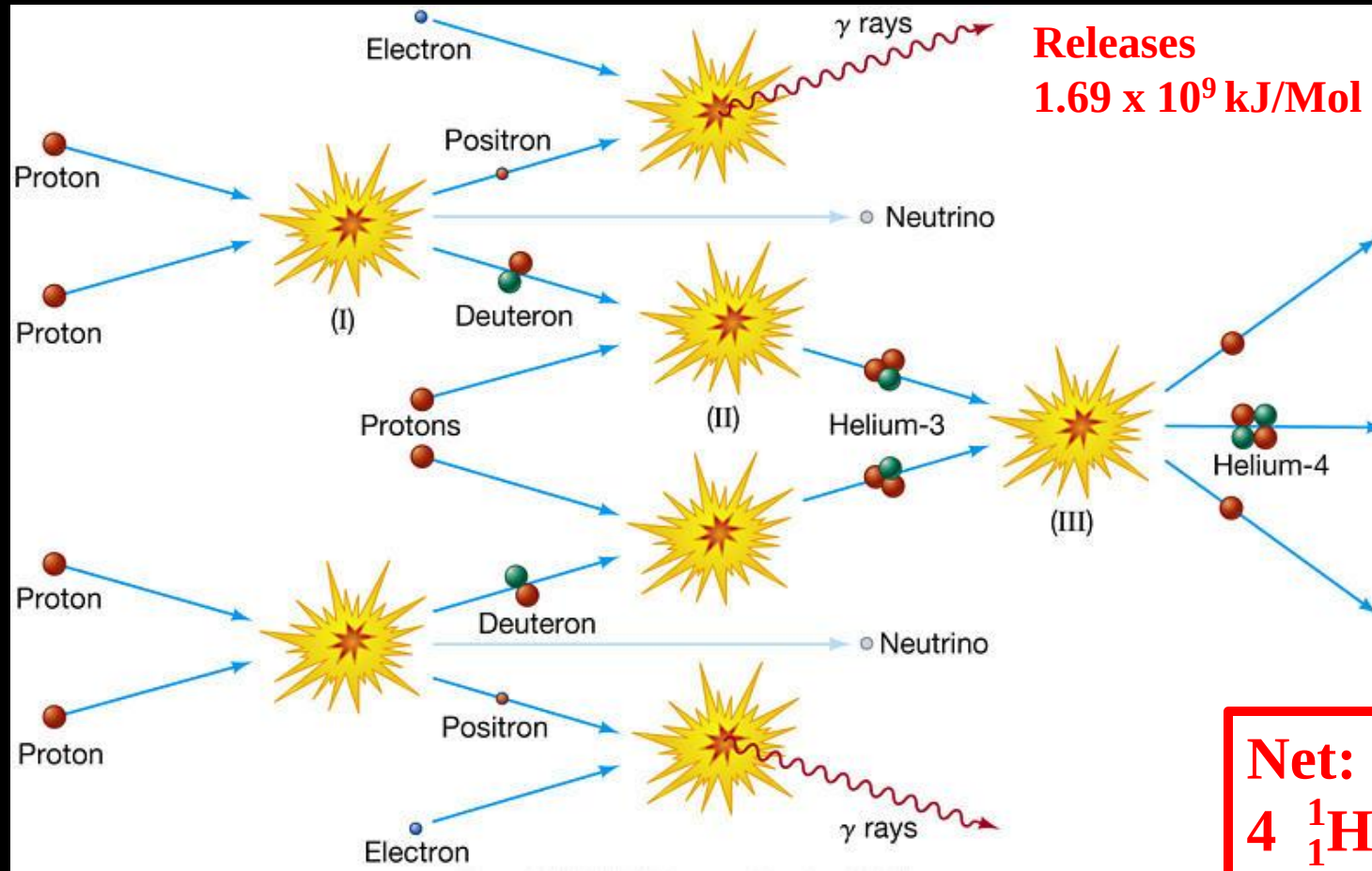


Thermonuclear Fusion

Fusion:

colliding small atoms together to form larger atoms

Energy source of stars



Releases

$1.69 \times 10^9 \text{ kJ/Mol}$

Requires:

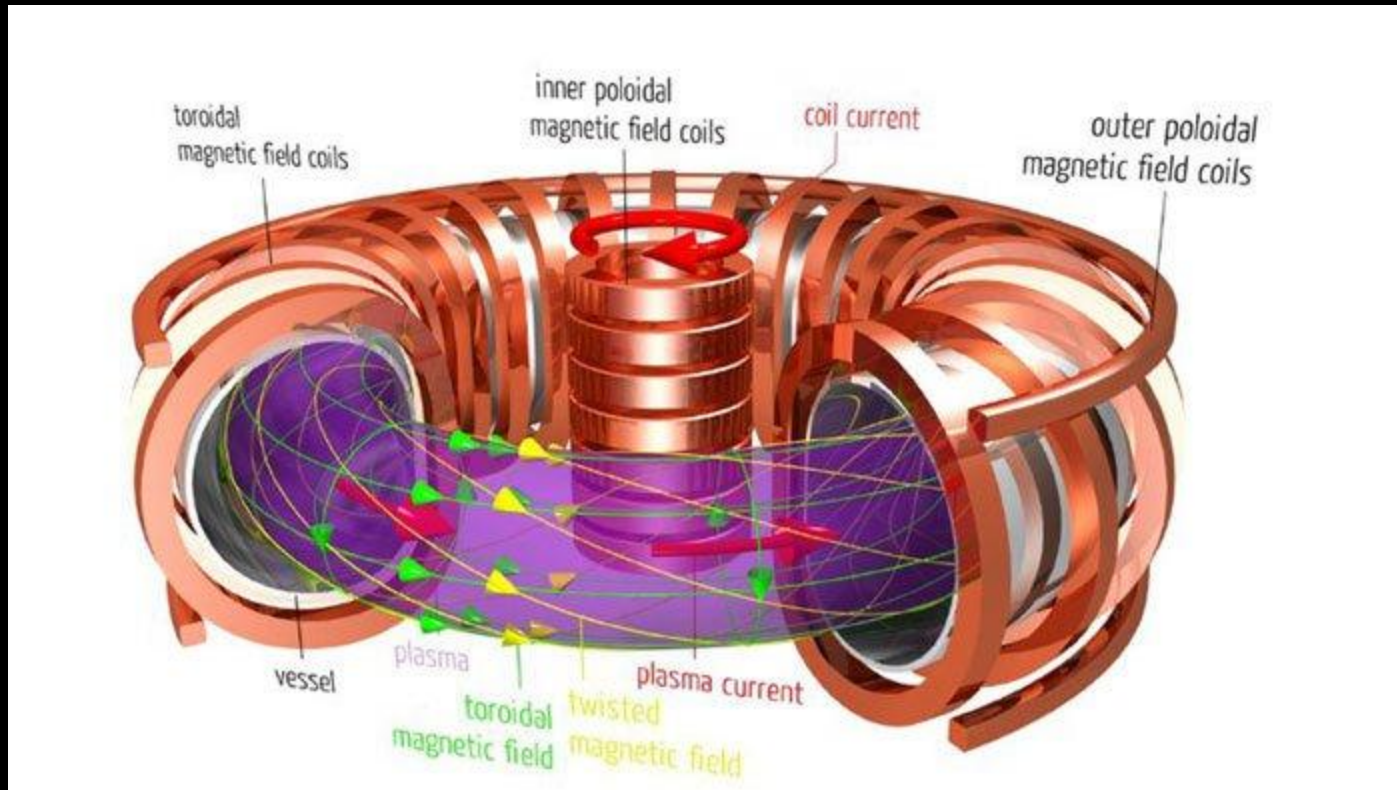
15 M K

Net:



Fusion Reactors

Theory: contain plasma in magnetic field
So far, only brief episodes successful



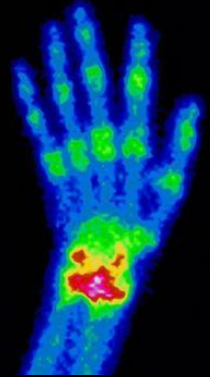
Uses of Radioisotopes

Radioisotopes have same chemistry as stable isotopes

Radioactivity allows detection

Allows radioisotopes to be effective tracers

Used in medical diagnosis and treatment



Common Medical Isotopes:

Technecium-99

Thallium-201

Iodine-131

Sodium-24

Annually:

10 million nuclear medicine procedures

100 million nuclear medical tests

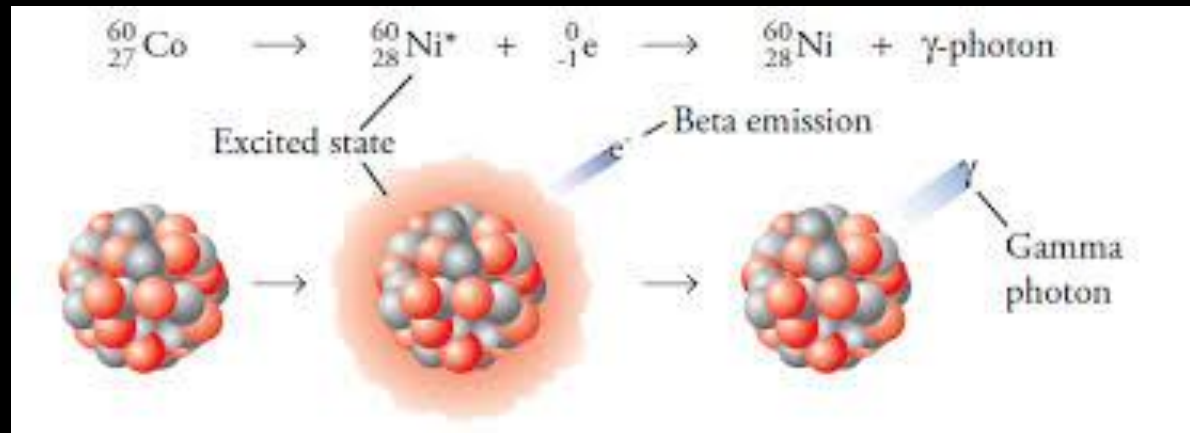
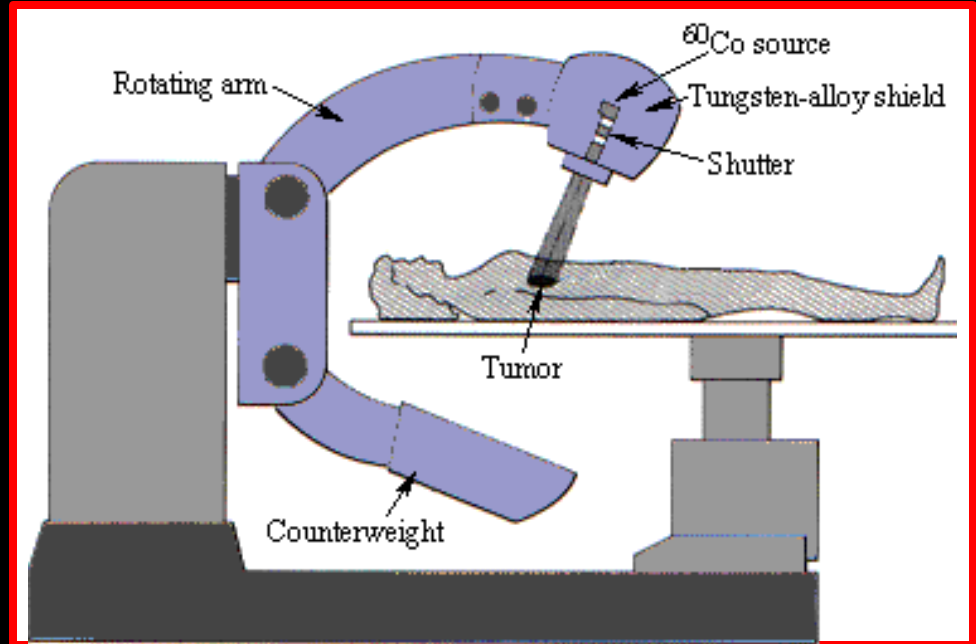


Uses of Radioisotopes

Radioactive Isotope	Applications in Medicine
Cobalt-60	Radiation therapy to prevent cancer
Iodine-131	Locate brain tumors, monitor cardiac, liver and thyroid activity
Carbon-14	Study metabolism changes for patients with diabetes, gout and anemia
Carbon-11	Tagged onto glucose to monitor organs during a PET scan
Sodium-24	Study blood circulation
Thallium-201	Determine damage in heart tissue, detection of tumors
Technetium-99m	Locate brain tumors and damaged heart cells, radiotracer in medical diagnostics (imaging of organs and blood flow studies)

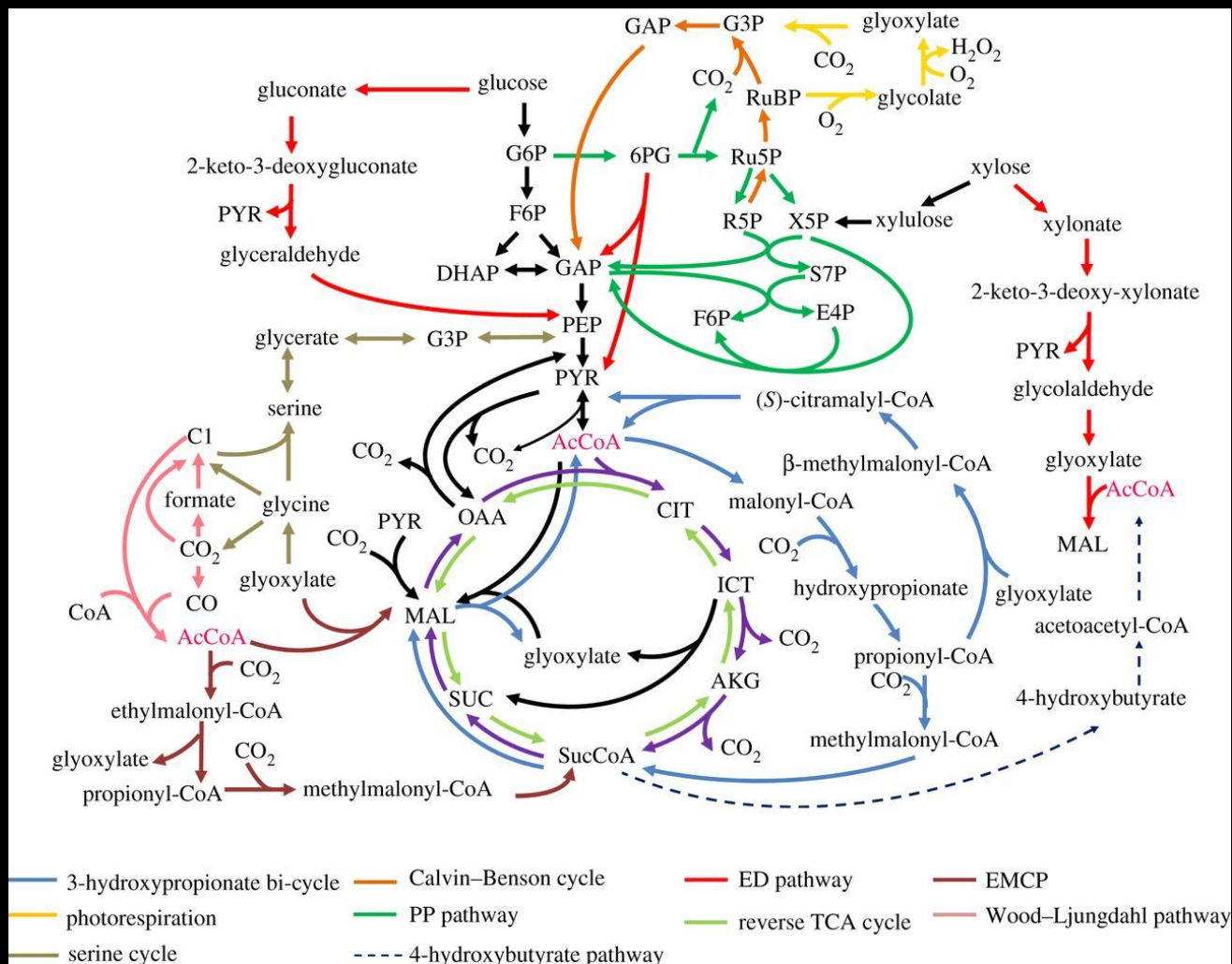
Short Half-lives ... often made on-site

Cobalt-60 Used in Radiation Therapy

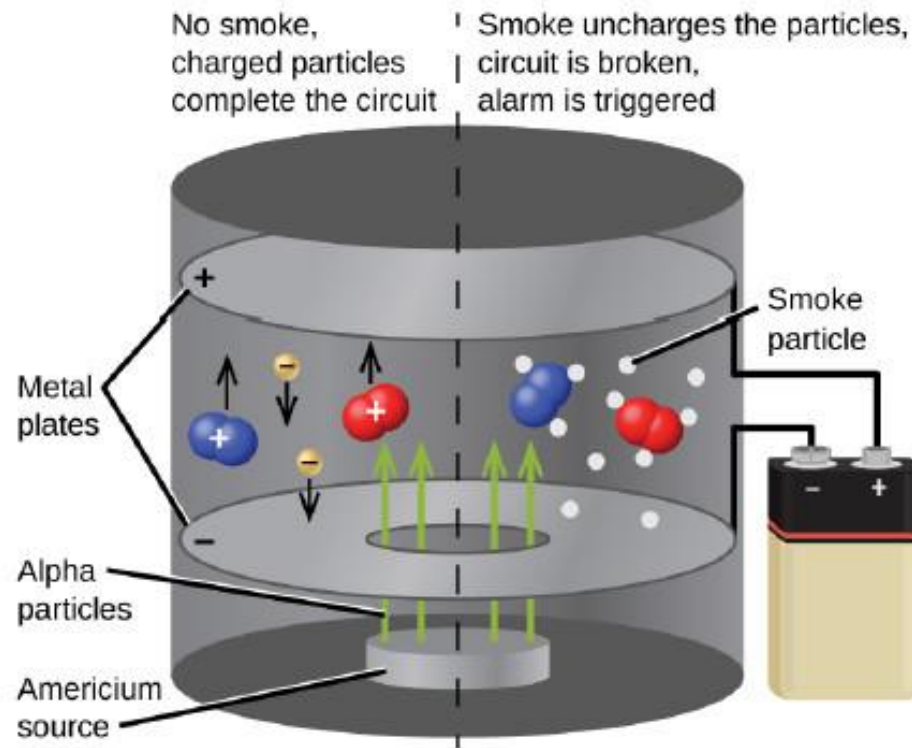
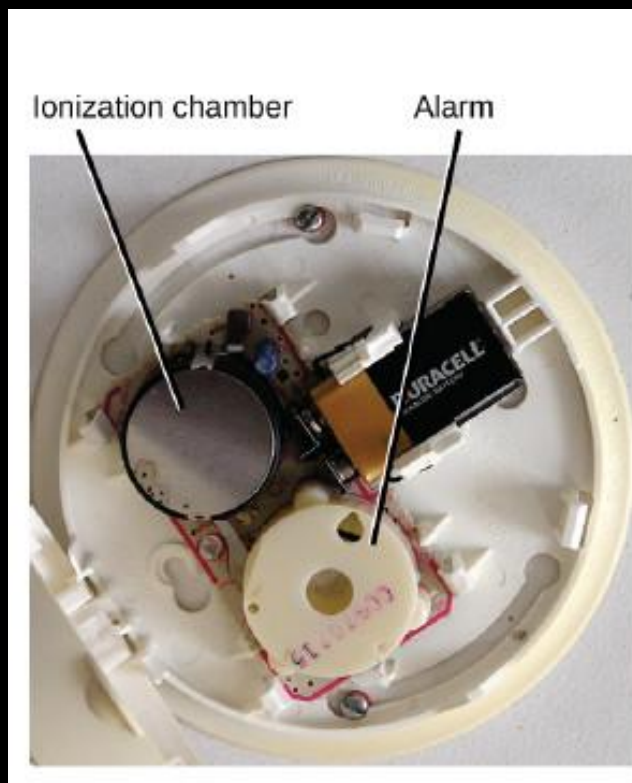


Gamma radiation kills cancer cells

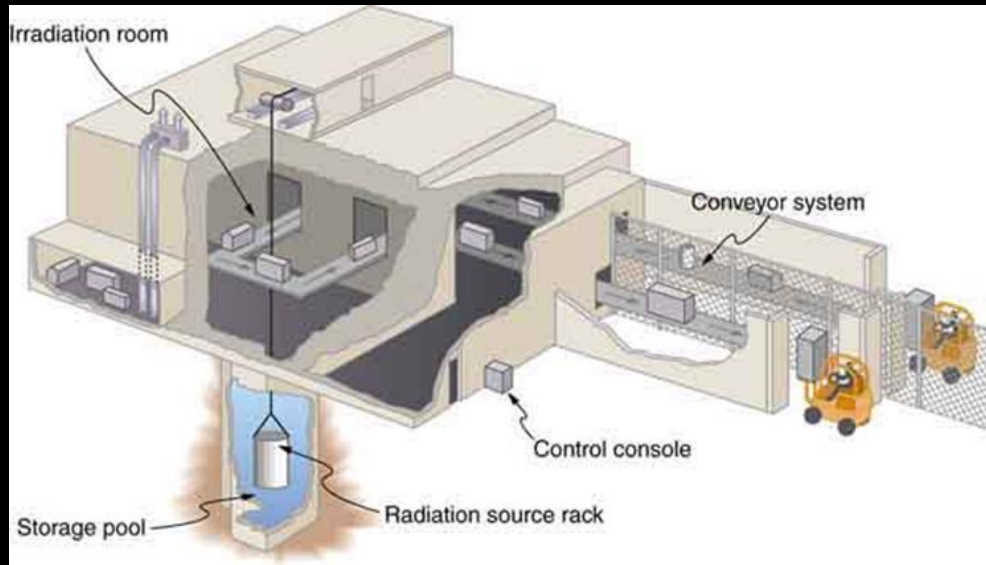
Carbon-14 Used as Tracer: Determines Biochemical Pathways



Americium-241: Smoke Detectors



Food Sterilization: C - 60 & Cs - 137



Objects exposed to γ -radiation

Food

Medical instruments

Bulk Cargo

Ions produced attack DNA

Biological dies



Biological Effects of Radiation

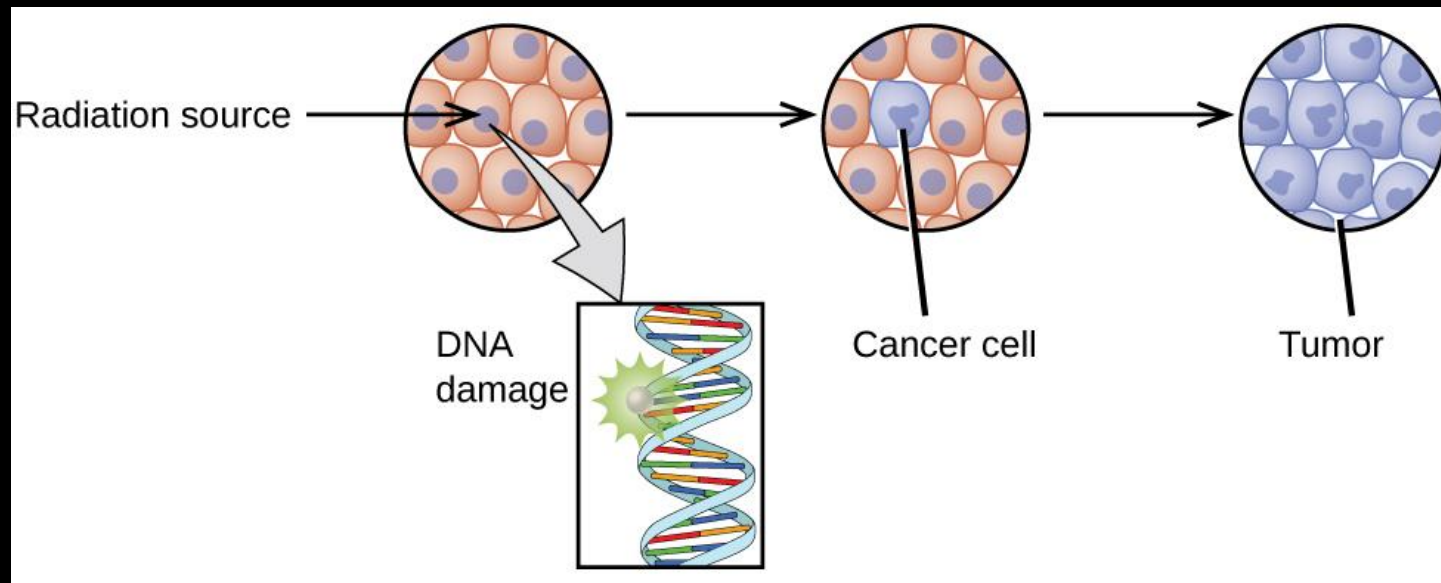
All radioactive materials emit high energy particles or electromagnetic waves

This radiation may:

thermal heating

chemical bond breaking

ionize molecules



No natural defenses against incoming radiation

Ionizing vs. Non-Ionizing Radiation

Ionizing

Higher energy electromagnetic waves (x-ray, γ – ray)

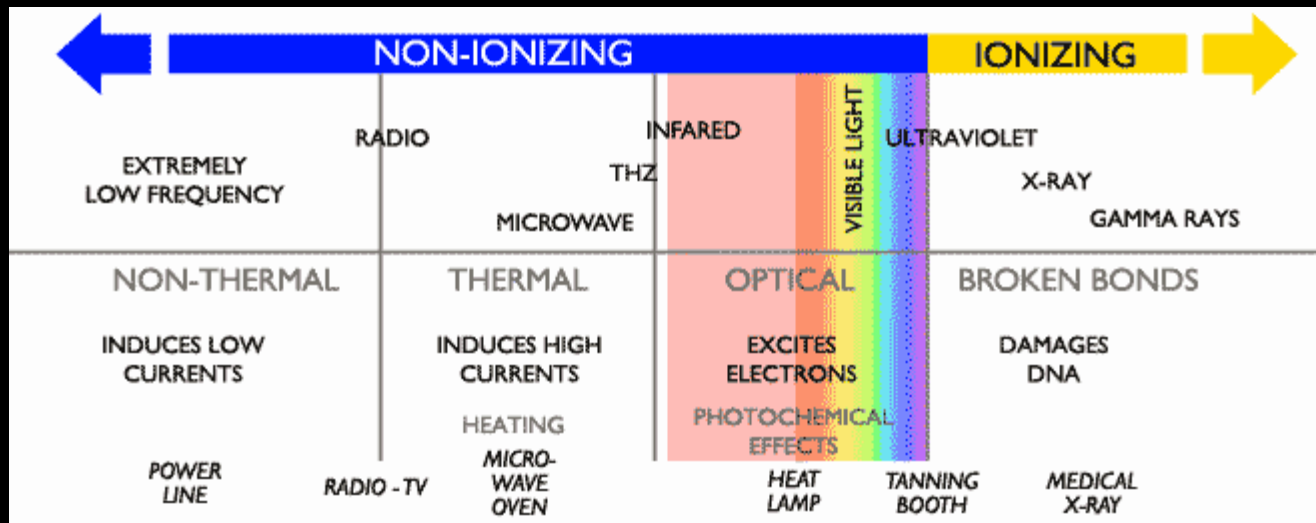
High energy particles (α , β particles)

Energy sufficient to pull electrons away from nucleus

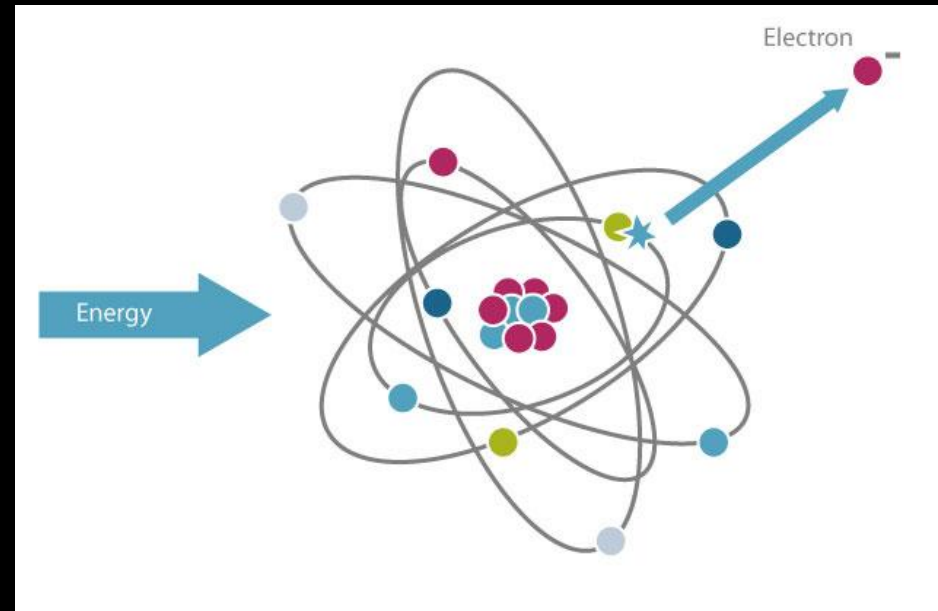
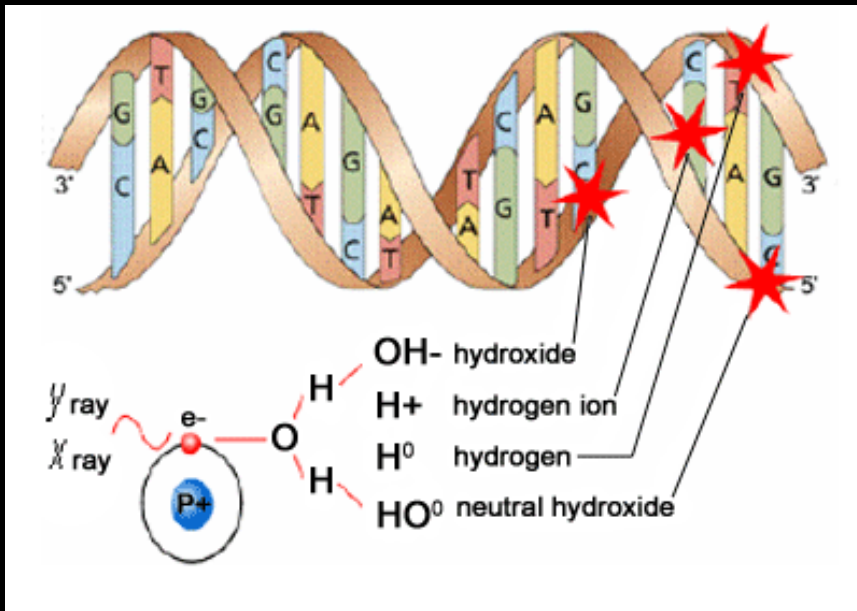
Non-Ionizing

Lower energy electromagnetic waves

Insufficient energy to ionize, but does excite electrons



Ionizing Radiation: Direct Effect

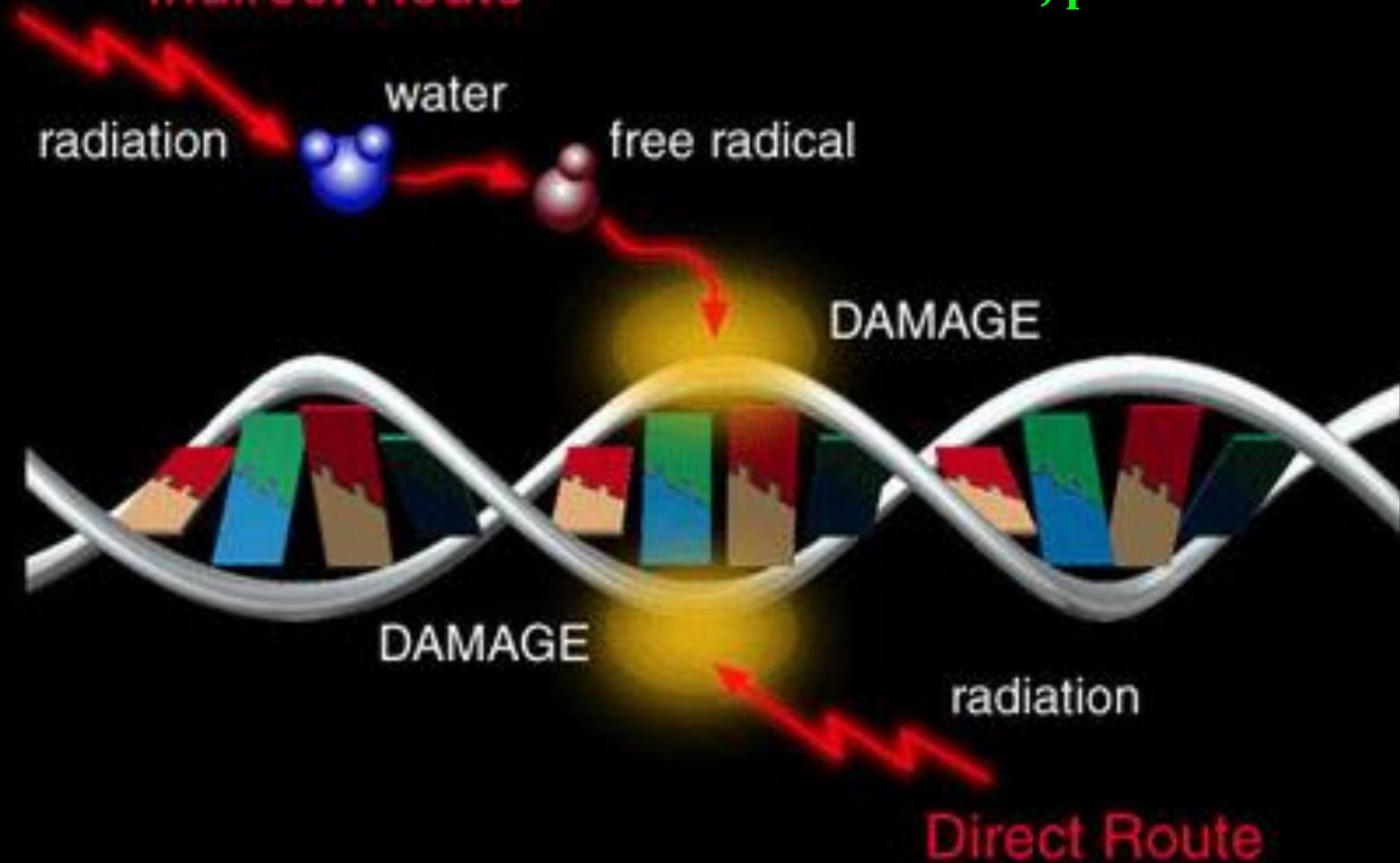


High energy radiation removes electron
Resultant ion has high energy
Biological structure altered

DNA is the primary target of radiation exposure

Non-Ionizing Radiation: Indirect Effect

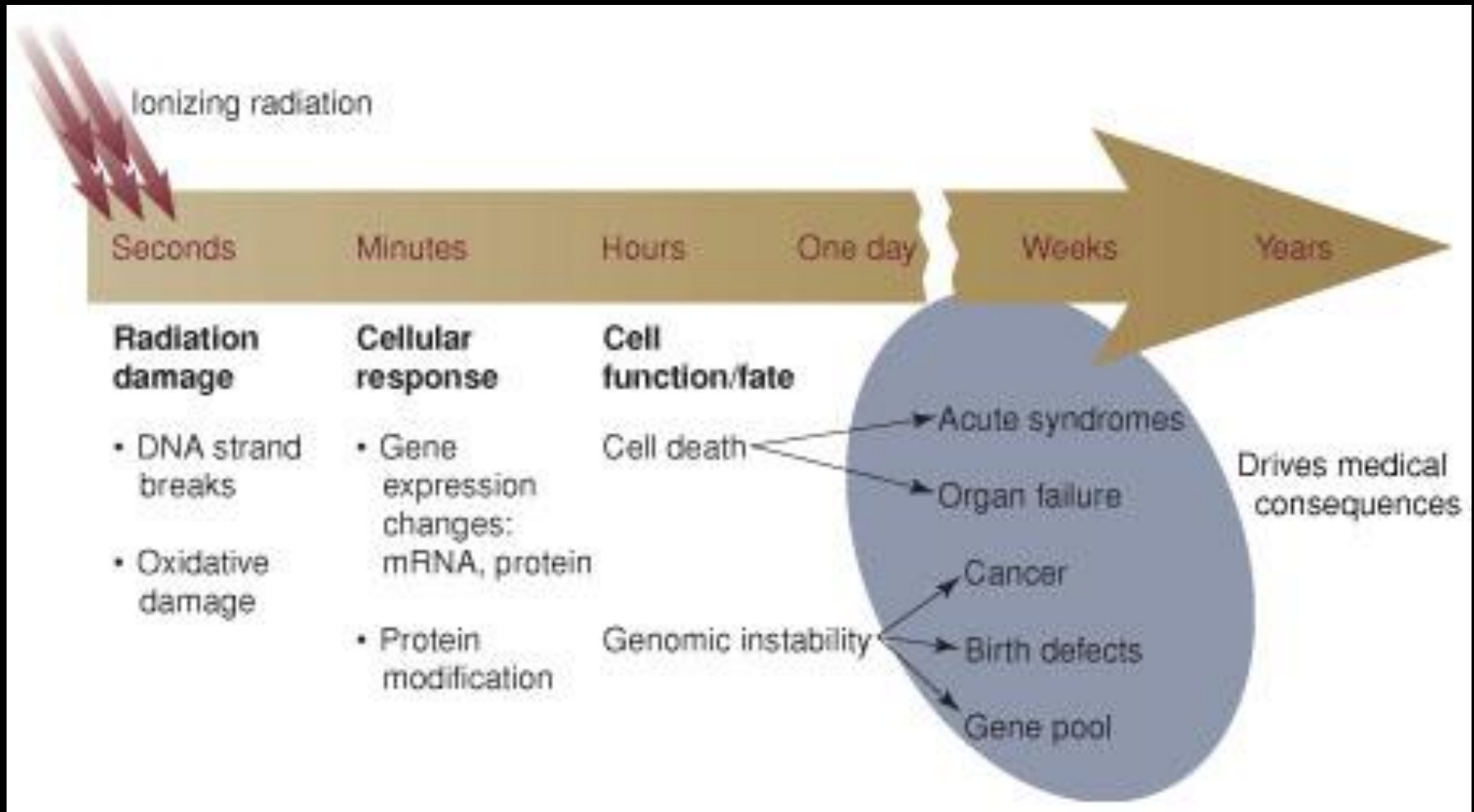
Indirect Route Radiation excites water; products attack DNA



Direct Route
Radiation directly attacks DNA

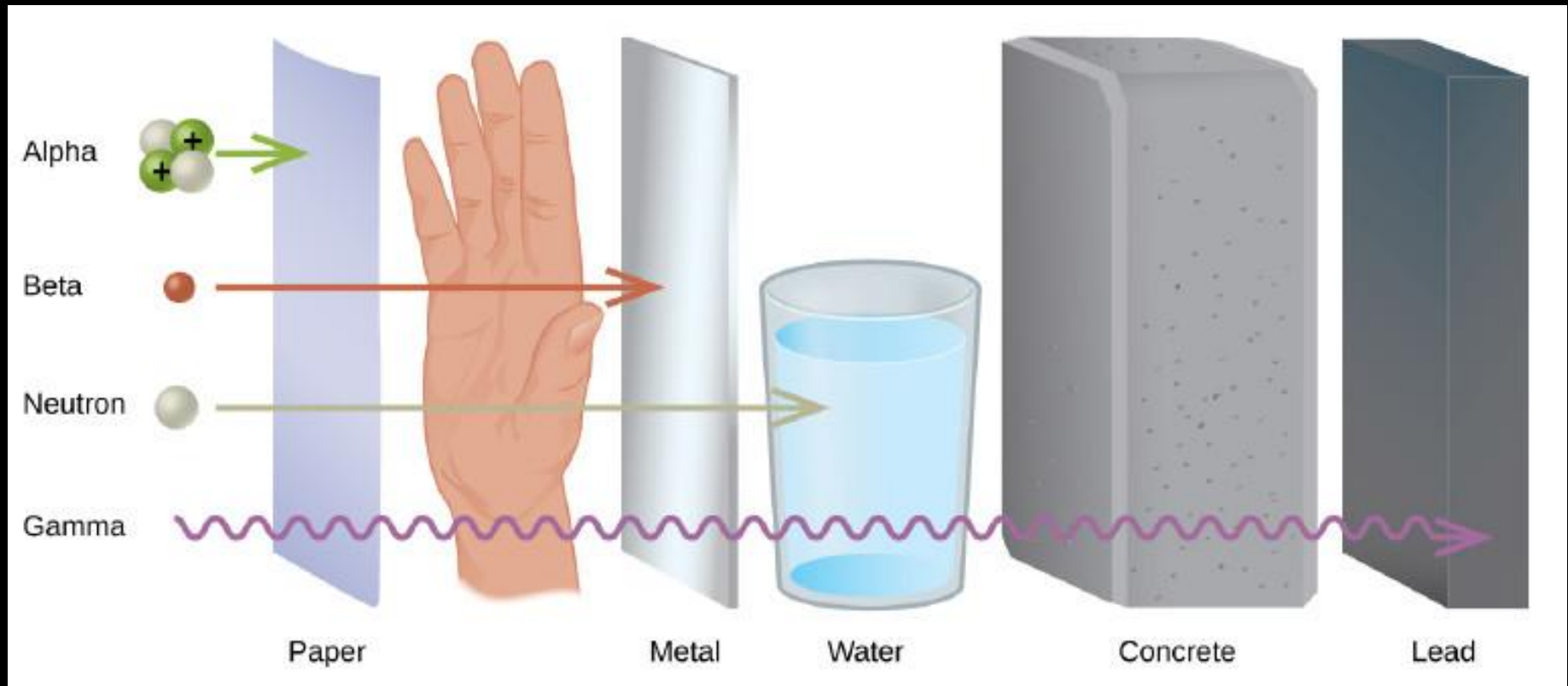
Biological Effects of Radiation

Insidious since damage may take years to manifest

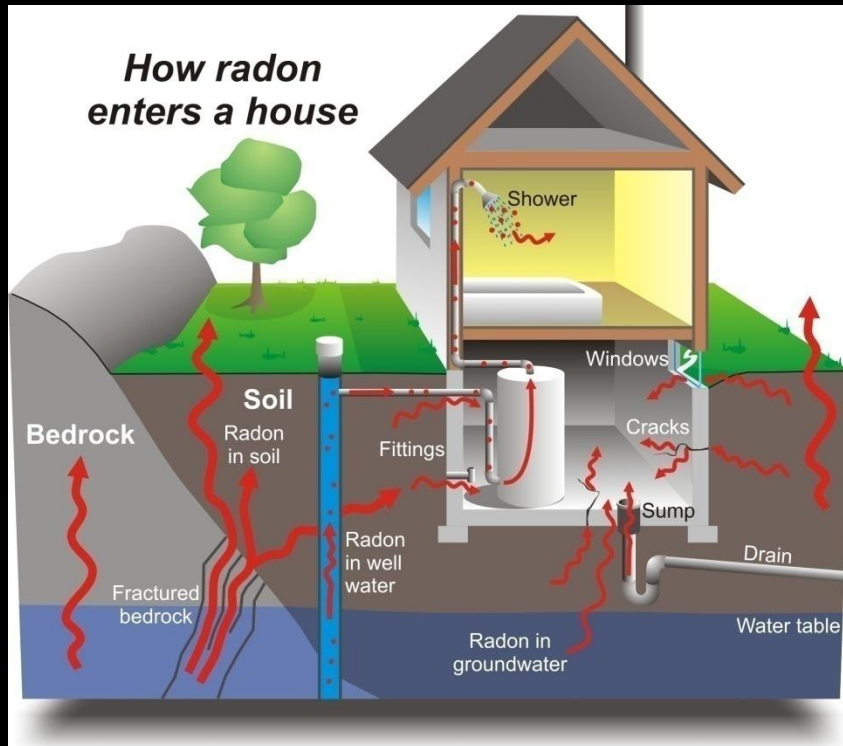
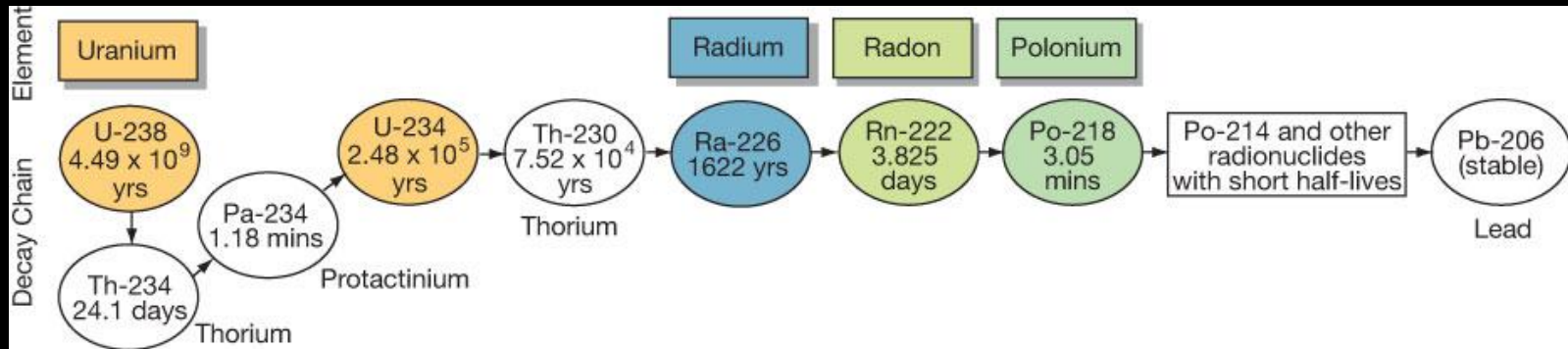


Radiation Penetration

> The energy, the > the penetration



Radon: Health Hazard in Michigan

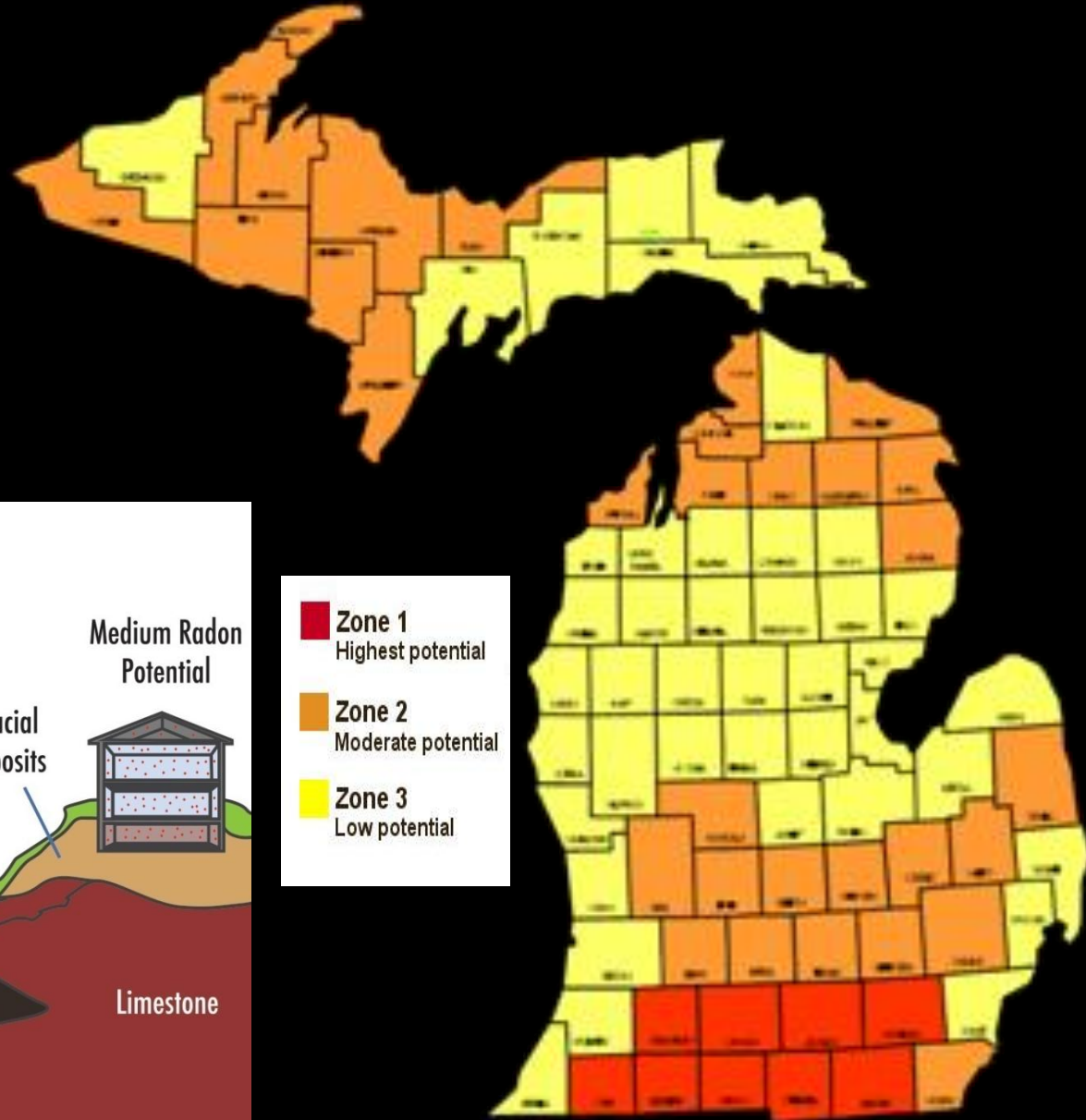
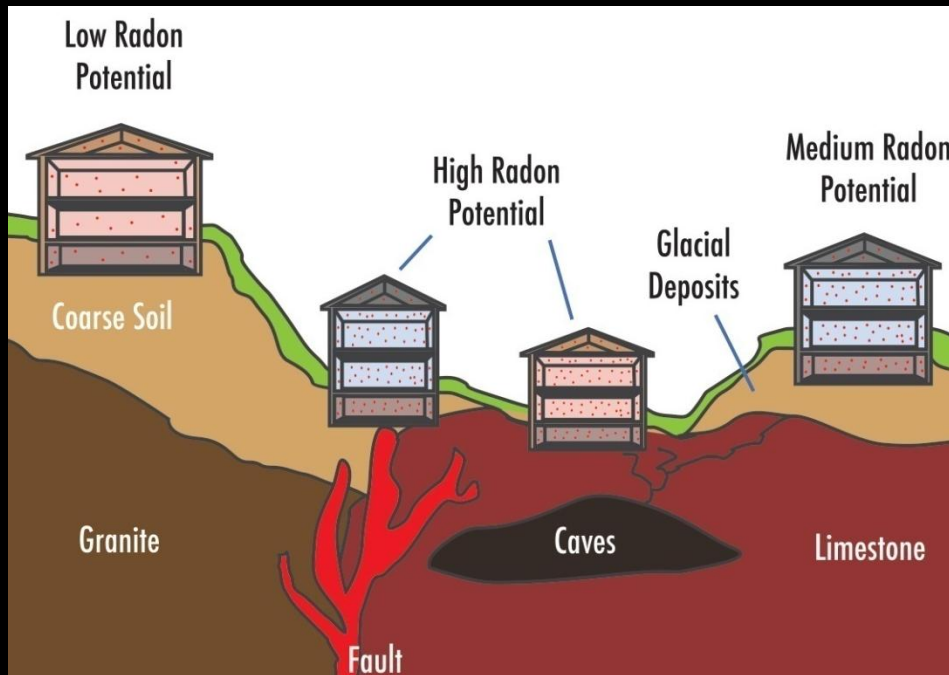
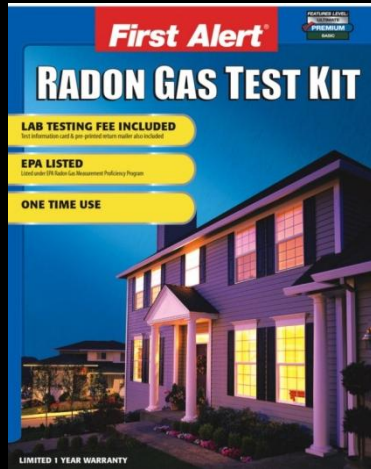


Major radiation exposure source
A Gas: An alpha emitter
Considered major risk for lung cancer

Sources of Radon

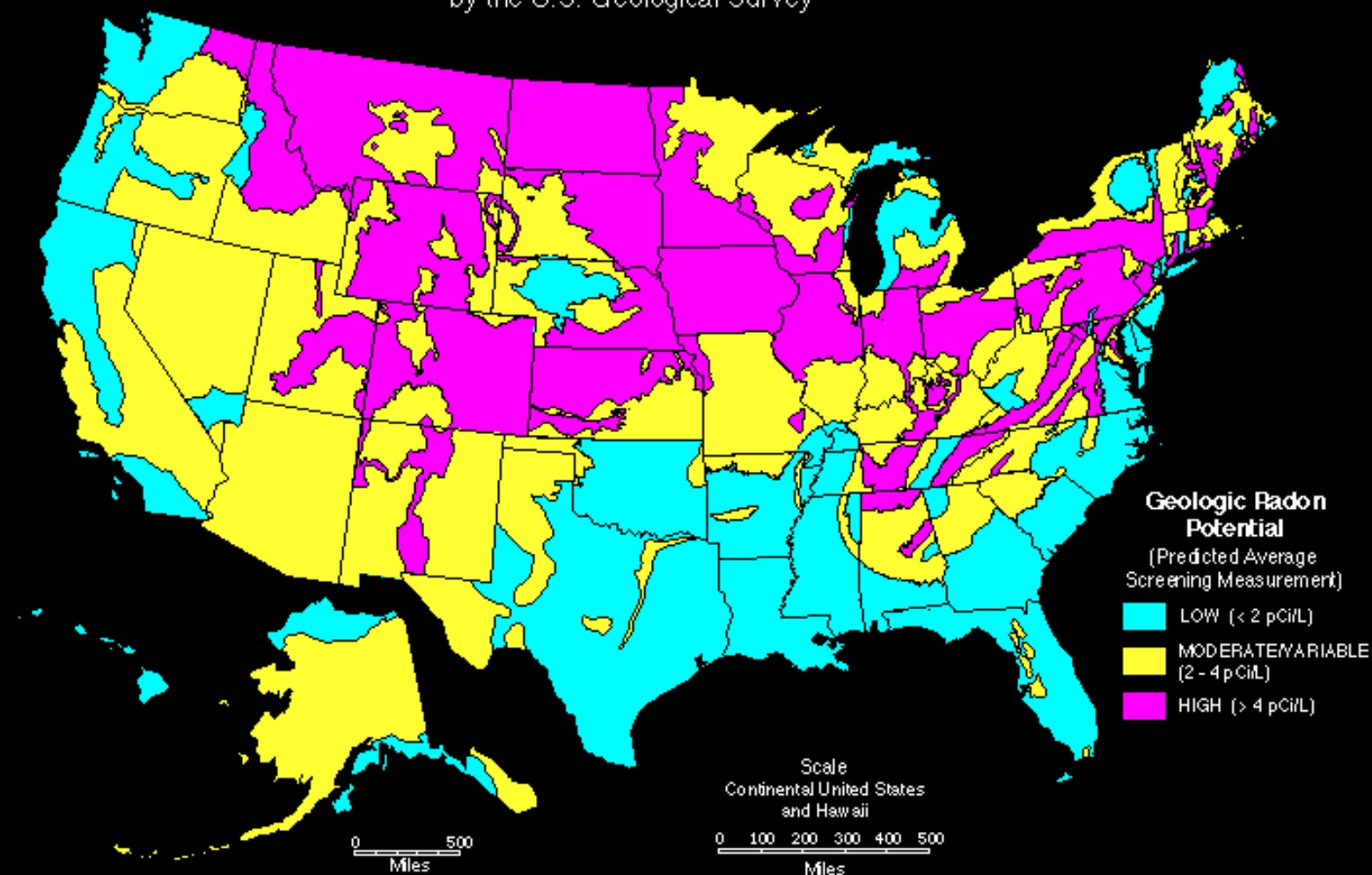
	Soil: 69.3%
	Well Water: 18.5%
	Outdoor Air: 9.2%
	Building Materials: 2.5%
	Public Water Supply: 0.5%

Radon: Health Hazard in Michigan



Radon: Health Hazard in United States

by the U.S. Geological Survey



Measuring Radiation Exposure

Geiger Counter

Gas filled tube

Radiation ionized the gas

Amount of radiation proportion to level of ionization



Scintillation Detector

Substance emits light when struck by radiation

Amount of radiation proportion to level of emitted light



Dosimeter (Dose meter)

Measure personal exposure

Many types:

film badge

electronic

thermoluminescence

quartz fiber

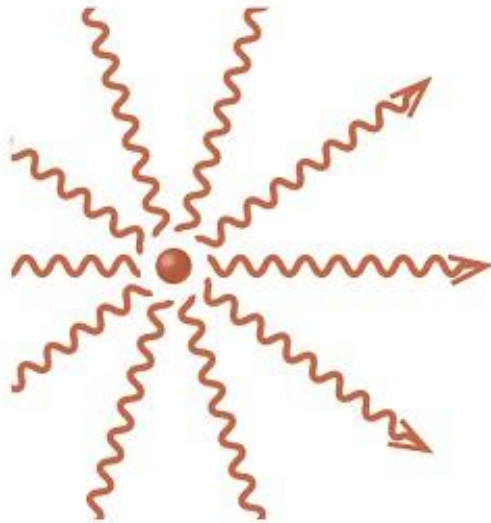


Units of Radiation Measurement

Measurement Purpose	Unit	Quantity Measured	Description
activity of source	becquerel (Bq)	radioactive decays or emissions	amount of sample that undergoes 1 decay/second
	curie (Ci)		amount of sample that undergoes 3.7×10^{10} decays/second
absorbed dose	gray (Gy)	energy absorbed per kg of tissue	1 Gy = 1 J/kg tissue
	radiation absorbed dose (rad)		1 rad = 0.01 J/kg tissue
biologically effective dose	sievert (Sv)	tissue damage	$Sv = RBE \times Gy$
	roentgen equivalent for man (rem)		$Rem = RBE \times rad$



Units of Radiation Measurement



Rate of radioactive decay
measured in becquerels or curies



Film badge or dosimeter
measures tissue damage
exposure in rems or sieverts

Absorbed dose measured in grays or rads

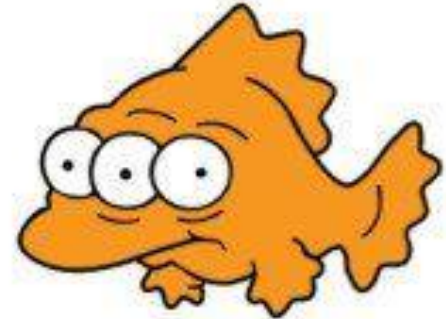
SIMPSONS GUIDE TO NUCLEAR RADIATION



Bequerel [Bq]
How brightly your
Cesium glows



Gray [Gy]
How brightly
Cesium will make
you glow



Sieverts [Sv]
How many extra
eyes will you have
after glowing?





Amount of Radiation

Co-60 ($t^{1/2} = 5.26$ y) is the γ source for cancer therapy
For a 5.00 gram sample, what is the activity?



Activity:

$$\lambda N = (\ln 2 / t^{1/2}) N = (\ln 2 / 5.26 \text{ y}) 5.00 \text{ g} = 0.659 \text{ g/y that decays}$$

Converting this activity to decays per second:

$$\frac{0.659 \text{ g}}{\text{year}} \times \frac{1 \text{ year}}{365 \text{ d}} \times \frac{1 \text{ day}}{24 \text{ hr}} \times \frac{1 \text{ hour}}{3600 \text{ s}} \times \frac{1 \text{ mol}}{59.9 \text{ g}} \times \frac{6.02 \times 10^{23} \text{ atoms}}{\text{mol}} \times \frac{1 \text{ decay}}{\text{atom}} = 2.10 \times 10^{14} \frac{\text{decay}}{\text{s}}$$

Since 1 decay/sec = 1 Bq, activity in Bq = 2.10×10^{14} Bq

Converting Bq to Ci:

$$2.10 \times 10^{14} \frac{\text{decay}}{\text{s}} \times \frac{1 \text{ Ci}}{3.7 \times 10^{11} \text{ decay/s}} = 5.7 \times 10^2 \text{ Ci}$$

Problem: Amount of Radiation

Tritium (H-3, $t^{1/2} = 12.32$ y); atomic mass= 3.016 amu is used in a variety of self-illuminating devices; Beta particle from the radioactive decay strikes a phosphor that produces a glow. What is the activity of a 1.00 mg sample?

Activity:

$$\lambda N = (\ln 2 / t^{1/2}) N = (\ln 2 / 12.32 \text{ y}) 1.00 \text{ mg} \times \frac{1}{1000} \frac{\text{g}}{\text{mg}} = 5.63 \times 10^{-5} \text{ g/y that decays}$$

Converting this activity to decays per second:

$$\frac{5.63 \times 10^{-5} \text{ g}}{\text{year}} \times \frac{1 \text{ yr}}{365 \text{ d}} \times \frac{1 \text{ d}}{24 \text{ hr}} \times \frac{1 \text{ hr}}{3600 \text{ s}} \times \frac{1 \text{ mol}}{3.016 \text{ g}} \times \frac{6.02 \times 10^{23} \text{ atoms}}{\text{mol}} \times \frac{1 \text{ decay}}{\text{atom}} = 3.56 \times 10^{11} \frac{\text{decay}}{\text{s}}$$

Since 1 decay/sec = 1 Bq, activity in Bq = 3.56×10^{11} Bq

Converting Bq to Ci:

$$3.56 \times 10^{11} \frac{\text{decay}}{\text{s}} \times \frac{1 \text{ Ci}}{3.7 \times 10^{10} \text{ decay/s}} = 0.952 \text{ Ci}$$



Long Term Exposure to Radiation

Radiation Effects on tissues depends upon:

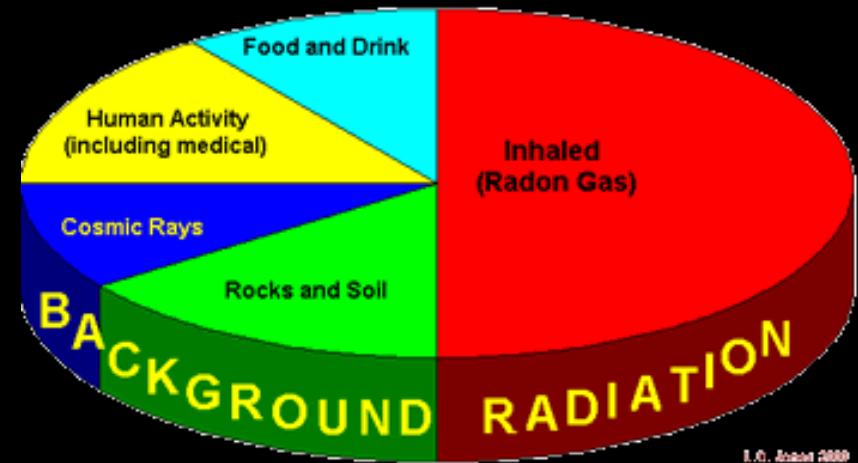
type of irradiation

energy at impact

internal vs. external

location of source

length of time of exposure



Sources:

Background: cosmic rays, radon, radioactive minerals

Medical: x-rays, CAT scans, tracer studies

Therapy: radioactive therapy agents

Cigarette Smoke

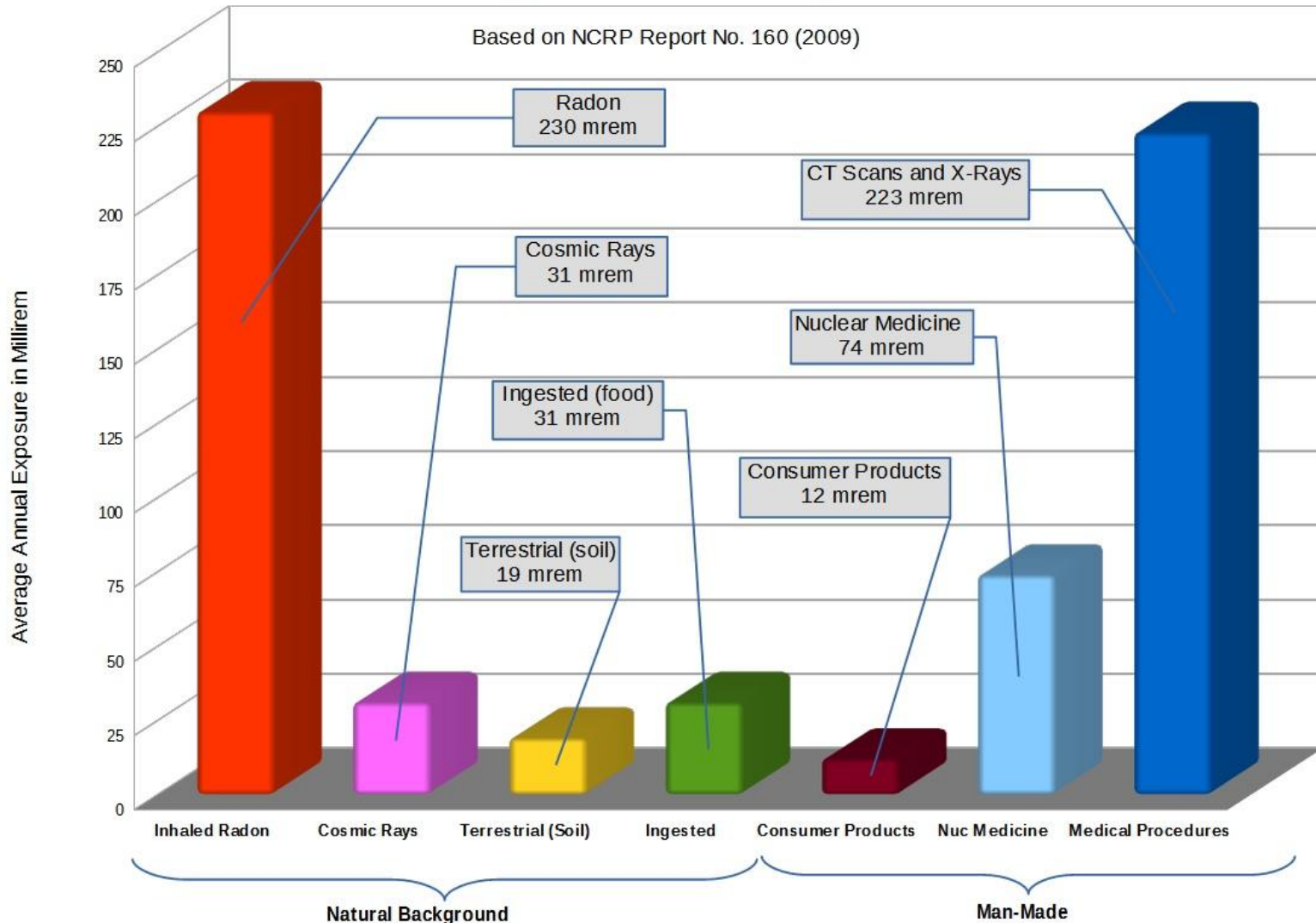
Food

Increased cosmic radiation at altitude (flying)

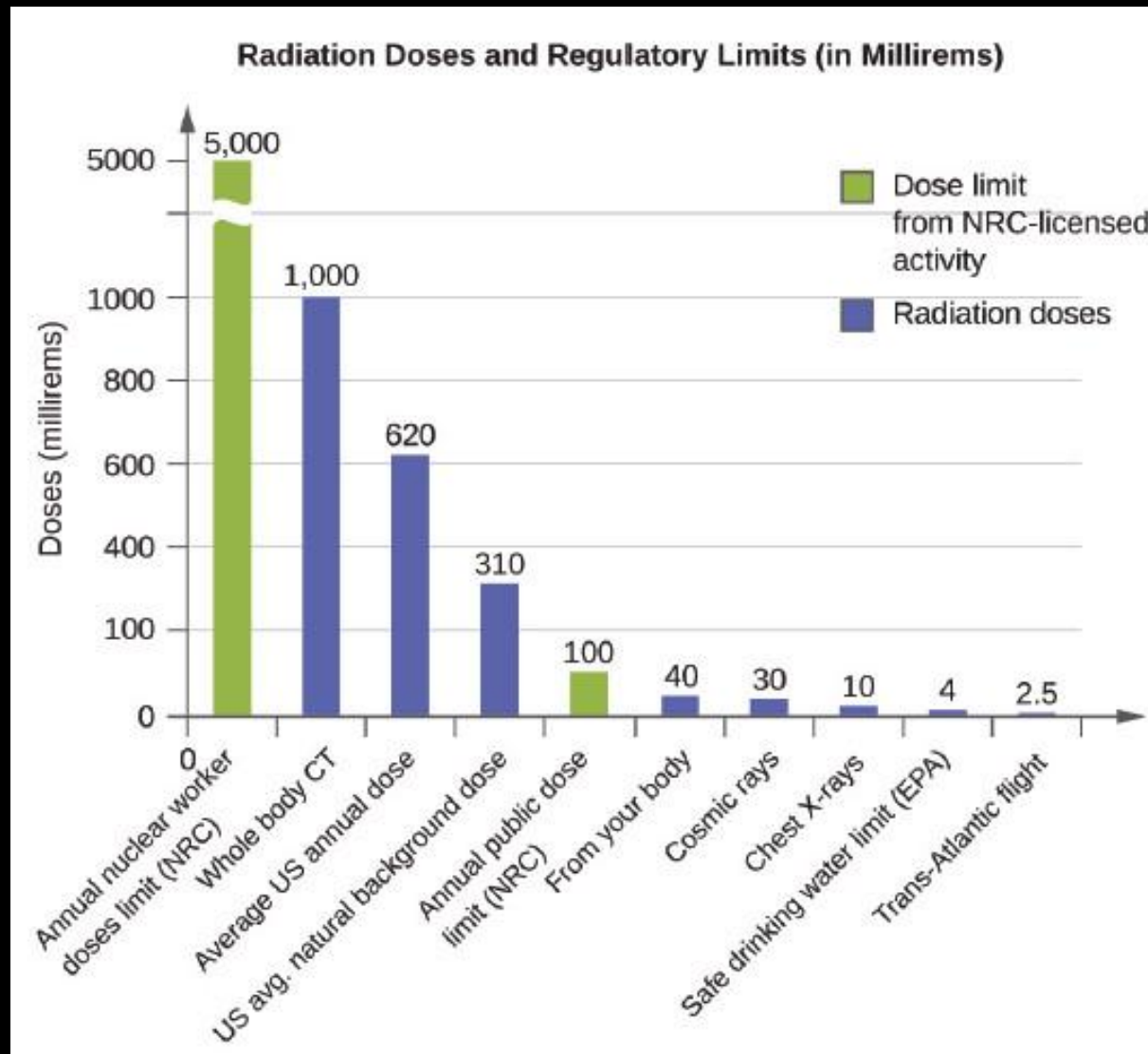
Long Term Exposure to Radiation

Sources of Average Annual Radiation Exposure in the US

Based on NCRP Report No. 160 (2009)



Long Term Exposure to Radiation



Long Term Exposure to Radiation

RADIATION EFFECTS

Measurements in millisieverts (mSv). Exposure is cumulative.

- **Potentially fatal radiation sickness.**
Much higher risk of cancer later in life.

10,000 mSv: Fatal within days.

5,000 mSv: Would kill half of those exposed within one month.

2,000 mSv: Acute radiation sickness.

- **No immediate symptoms. Increased risk of serious illness later in life.**

1,000 mSv: 5% higher chance of cancer.

400 mSv: Highest hourly radiation recorded at Fukushima .
Four hour exposure would cause radiation sickness.

100 mSv: Level at which higher risk of cancer is first noticeable

- **No symptoms. No detectable increased risk of cancer.**

20 mSv: Yearly limit for nuclear workers.

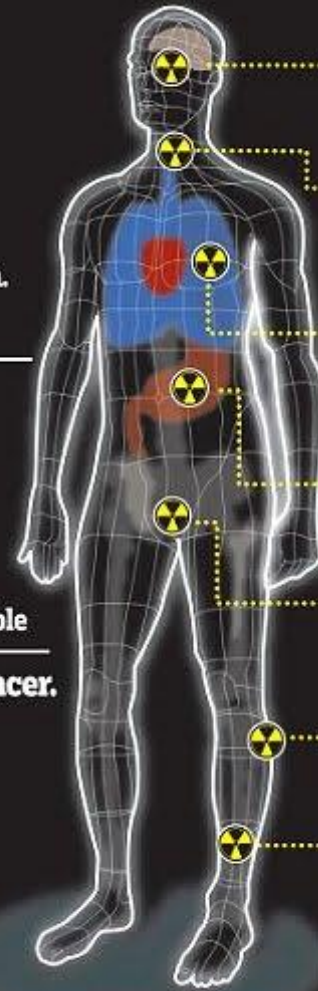
10 mSv: Average dose from a full body CT scan

9 mSv: Yearly dose for airline crews.

3 mSv: Single mammogram

2 mSv: Average yearly background radiation dose in UK

0.1 mSv: Single chest x-ray



EYES High doses can trigger cataracts months later.

THYROID Hormone glands vulnerable to cancer. Radioactive iodine builds up in thyroid. Children most at risk.

LUNGS Vulnerable to DNA damage when radioactive material is breathed in.

STOMACH Vulnerable if radioactive material is swallowed.

REPRODUCTIVE ORGANS High doses can cause sterility.

SKIN High doses cause redness and burning.

BONE MARROW Produces red and white blood cells. Radiation can lead to leukaemia and other immune system diseases.

Long Term Exposure to Radiation

Organ Sensitivity To The Effects Of Radiation

LOW

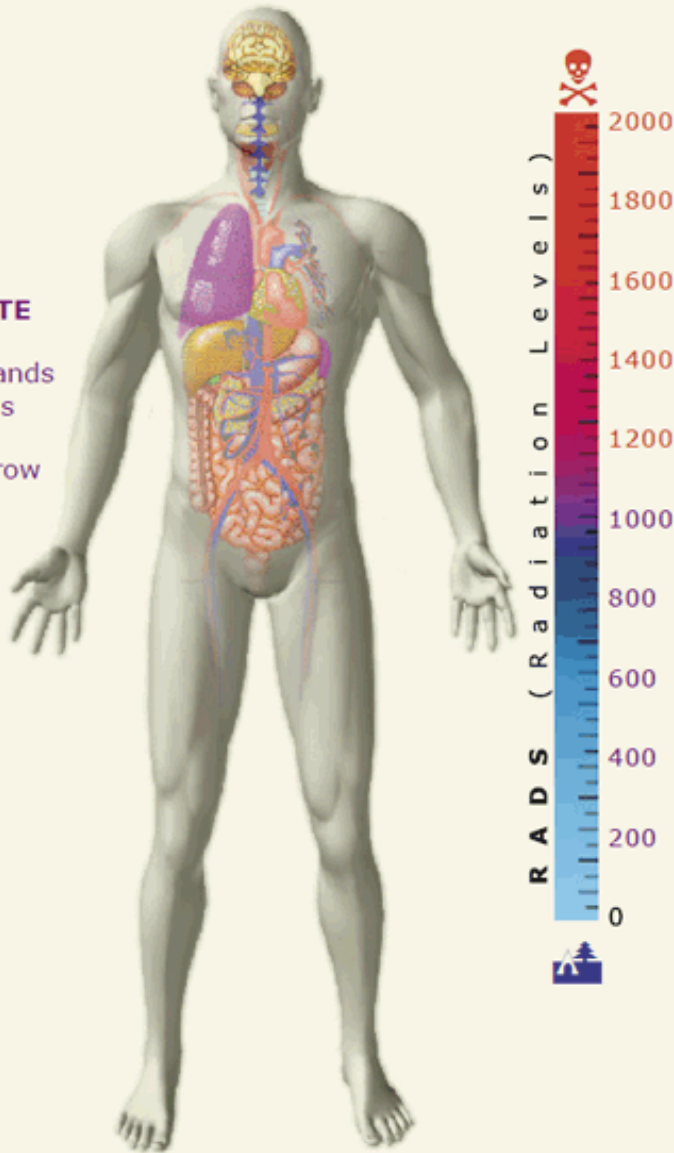
Skin
Bones
Spleen
Kidneys

MODERATE

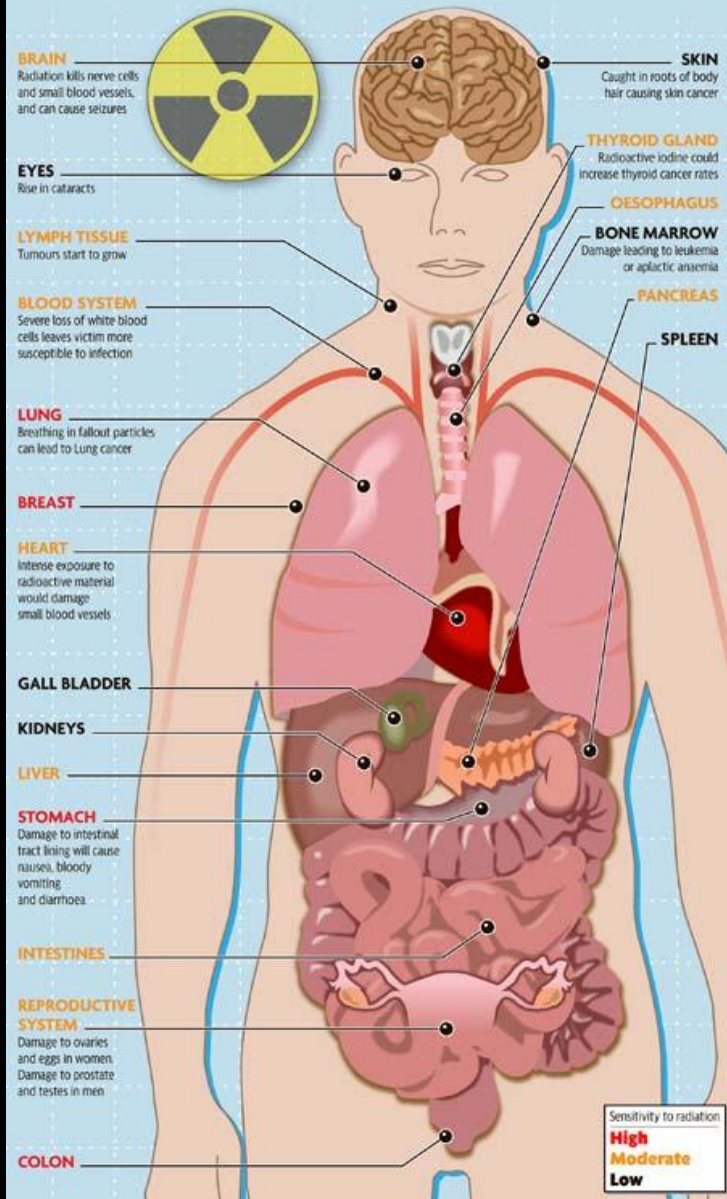
Brain
Lymph Glands
Esophagus
Thyroid
Bone Marrow
Liver
Pancreas
Intestines
Ovaries

HIGH

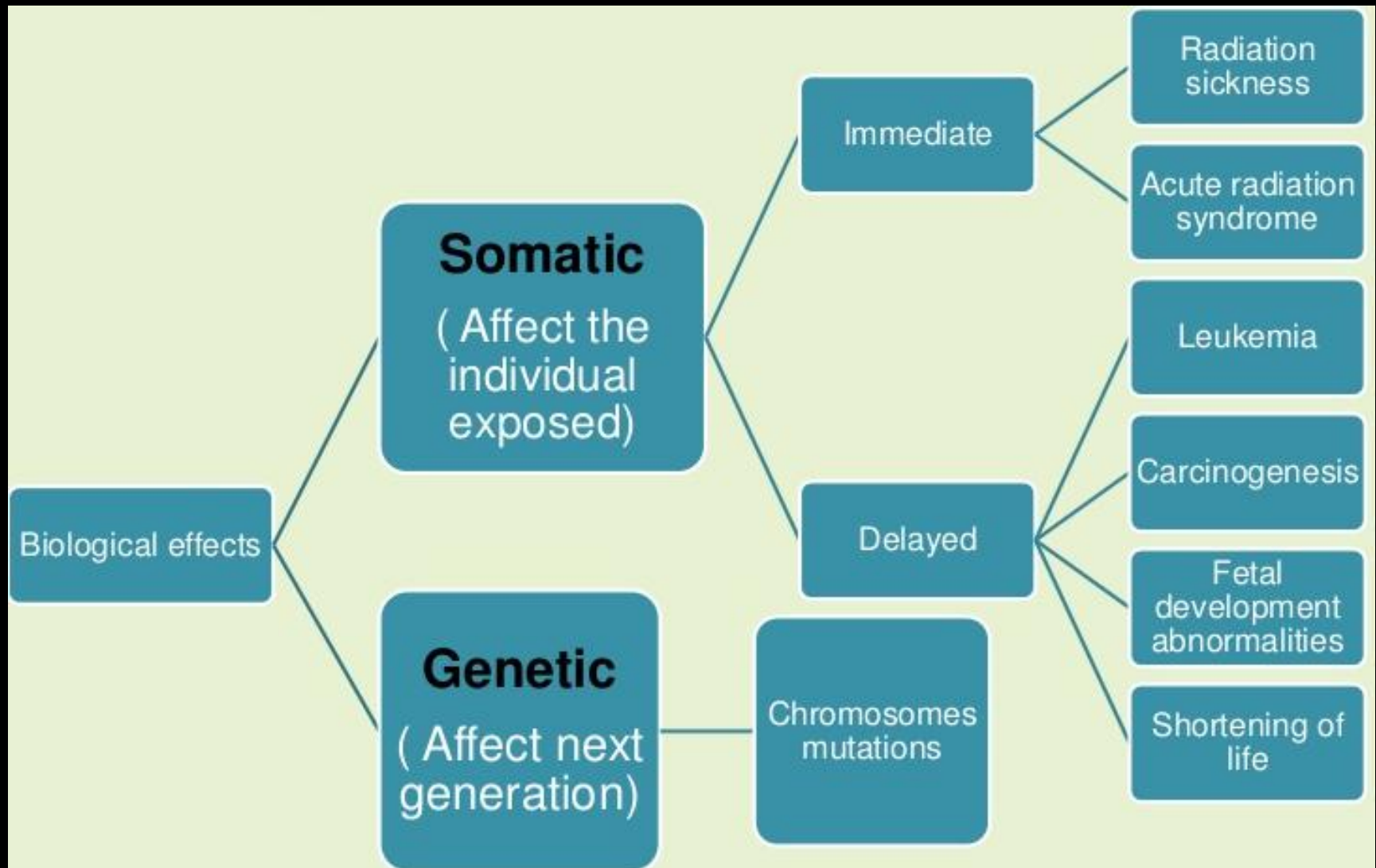
Lungs
Breasts
Stomach
Colon



EFFECTS OF NUCLEAR RADIATION ON THE HUMAN BODY



Biological Effects of Ionizing Radiation



Effects of Radioactivity Over Time

Hour:

1



4



6



9



12



14



17



19



21



23



26



29



31



36

